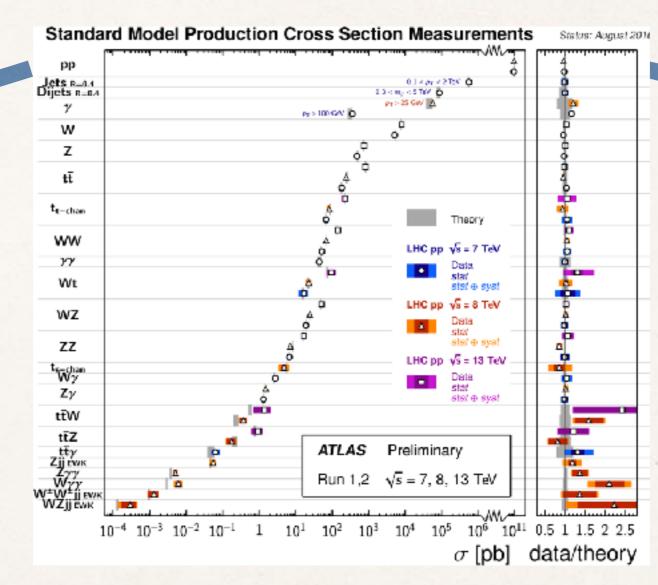
Beyond the Standard Model: The challenges ahead Veronica Sanz (Sussex)

UNiversity of Sussex

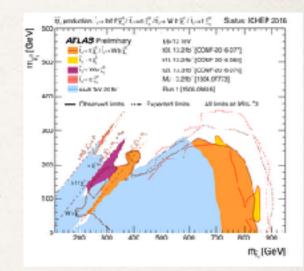
IFT seminar, April 2019

Let's start with the LHC The LHC is in a mature stage, already providing precision tests for the SM in most channels (excl the Higgs)

Precise tests of the full structure of the SM, based on QFT, symmetries (global/ gauge) and consistent ways to break them non-trivial tests of perturb.->non-perturb. QCD

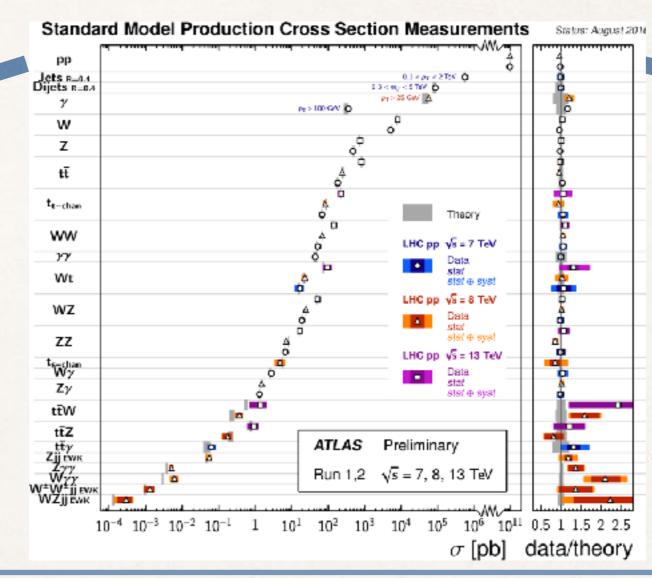


Absence of excesses: interpreted as new physics exclusions

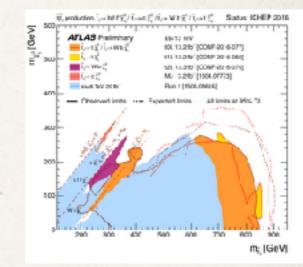


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Absence of excesses: interpreted as new physics exclusions



exclusions: rather impressive, many at the TeV searches: outstanding coverage of possible topologies any hints: (like in flavor) extremely tempting

So here we are

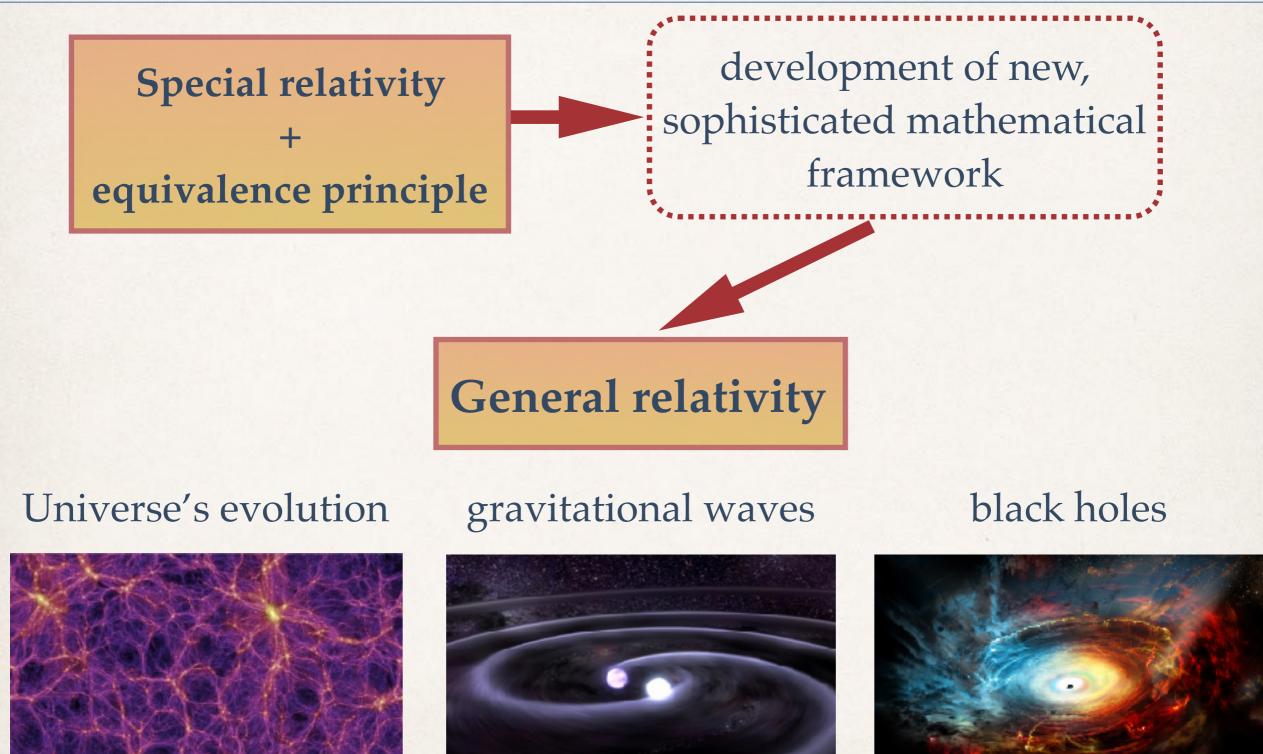
Light Higgs Inflation Neutrinos
Matter/Antimatter
Dark Energy CP QCD Dark Matter
Quantum Gravity

finding our path through **SYMMETRIES & DYNAMICS**

aiming for a UNIFIED FRAMEWORK

SM+GR

What we would hope for





Some years ago String theory, *the* final theory Mathematical consistency (anomalies, SUSY) +guiding principles (QGrav, unification, 3 families) trickle down to the SM, a boundary condition

Light Higgs

Inflation

CP QCD

Matter/Antimatter

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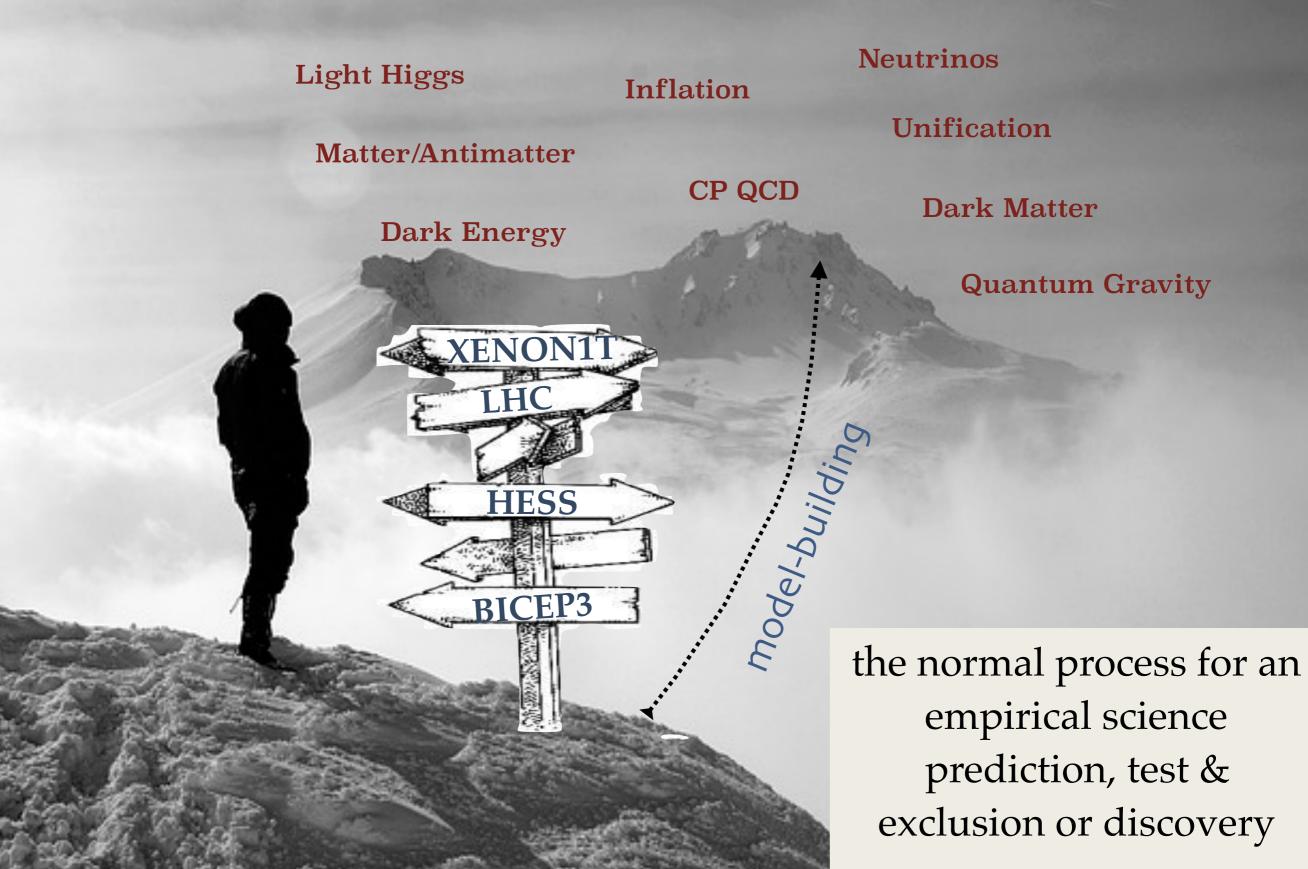
CP QCD

Dark Matter

Quantum Gravity

This program did not lead to identifying *the* theory (see string lanscape) instead, generated a **vast number of new ideas:** reformulations of gravity and QFT dualities incl AdS/CFT new scenarios for model-building incl duals of RS (composite higgs, clockwork), models for inflation

So here we are again, post-LHC Run2



New theoretical ideas



Experimental results challenge long-standing common lore in Particle Physics e.g. SUSY& naturalness, WIMP

New theoretical developments are very much needed. Simplified models are good *proofs of concept* but don't bring the field very far

*Extend the current models with more features e.g ALPs non-vanilla couplings *Find a rationale for hidden sectors * BE BOLD! this is the right environment for new ideas

One way forward: Connecting ideas/experiments

A cosmological Higgs

Dark Matter Higgs portal Higgs DM mediator UV sensitivity Naturalness heavy new physics Relaxation

Inflation Higgs inflation Inflaton vs Higgs

Phase transitions Baryogenesis gravitational waves

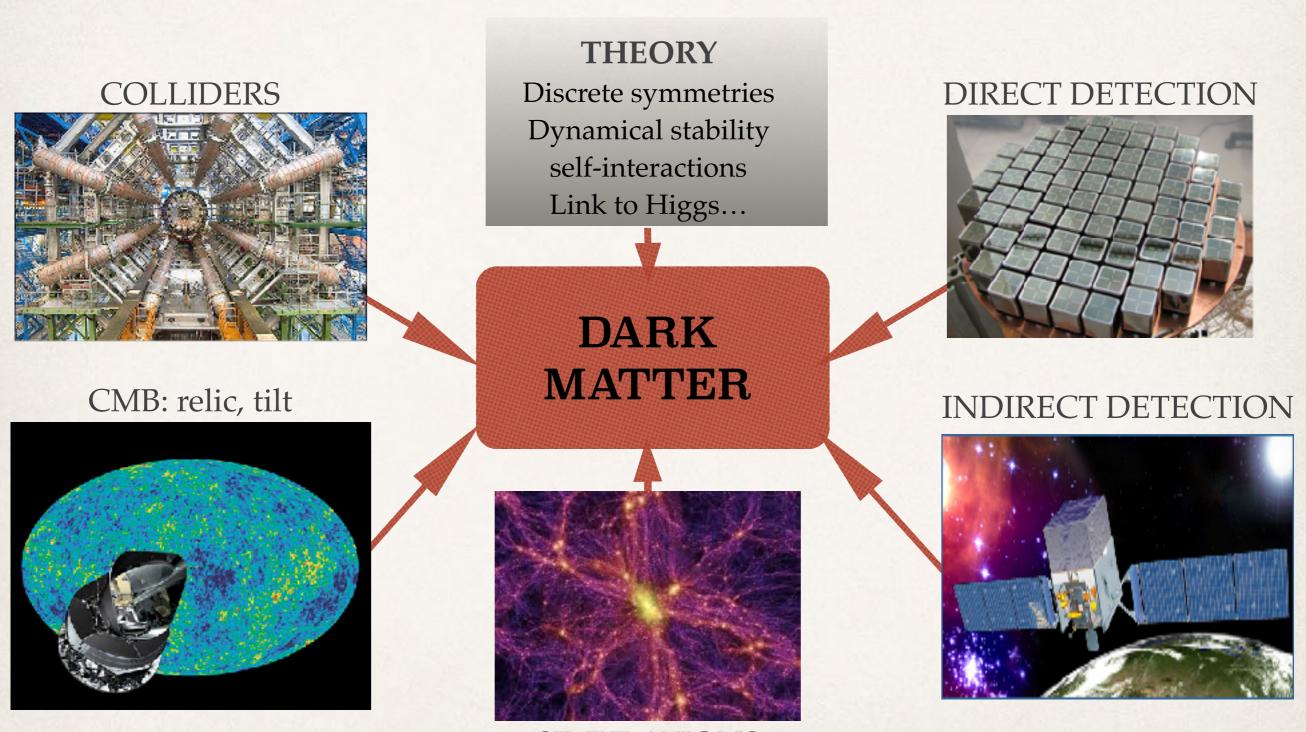
HIGGS

Fate of the Universe Stability

The LHC provides the most precise, controlled way of studying the Higgs and direct access to TeV scales Exploiting complementarity with cosmo/astro probes

Similar story for Axions and ALPs, scalars are versatile

Many faces of Dark Matter

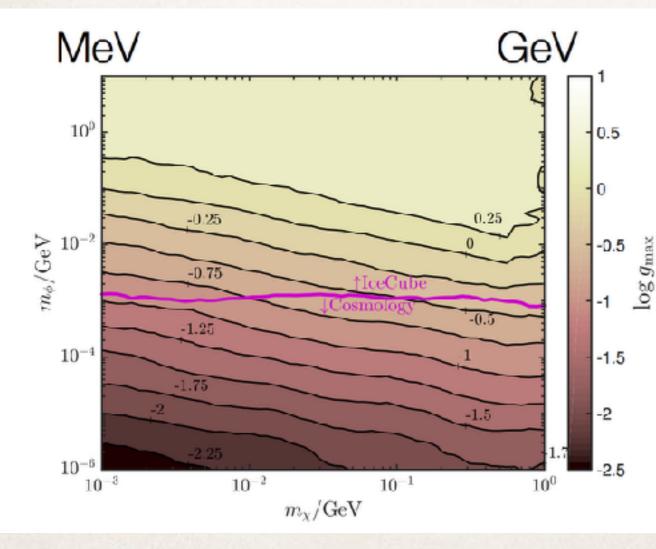


SIMULATIONS

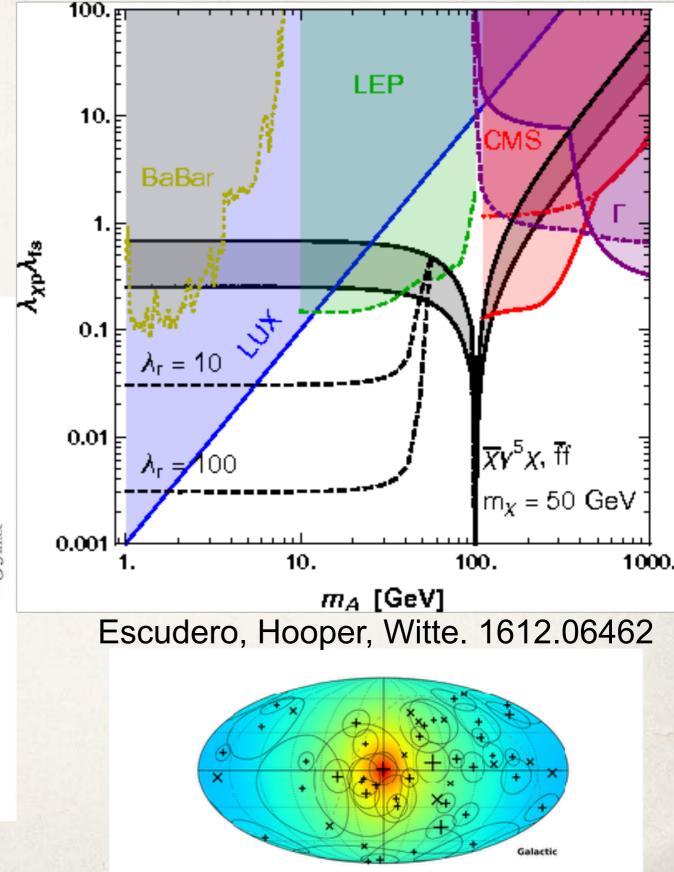
Examples of connections: Synergies colliders/non-colliders

Astrophysics/others

example: propose a solution to an astrophysical excess with a PP model, explore whether it is related to a coupling with neutrinos



Arguelles, Keirandish, Vincent. 1703.00451



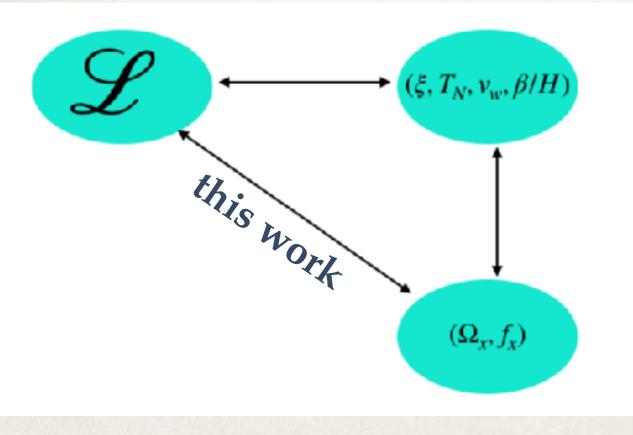
 $\log(\rho_{DM}/\text{GeVm}^{-2})$

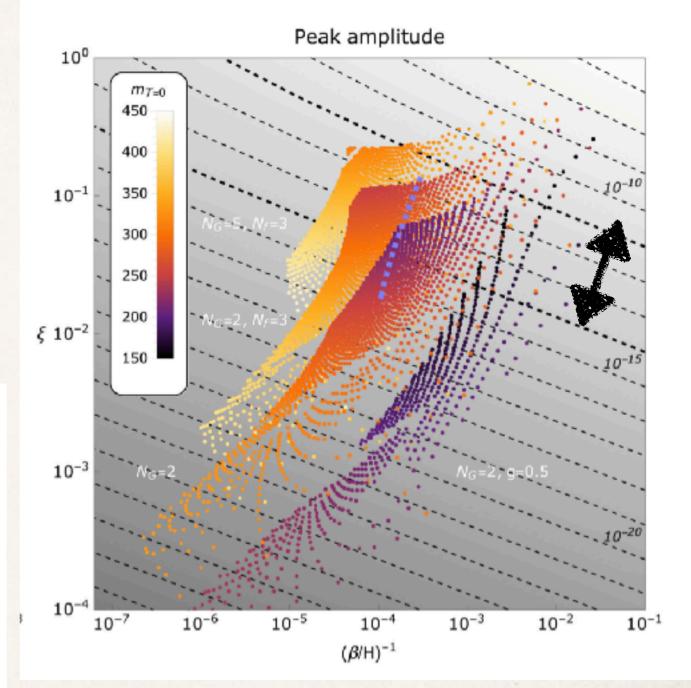
23.0201

Gravitational waves/others

another example: CROON, VS, WHITE. 1806.02332

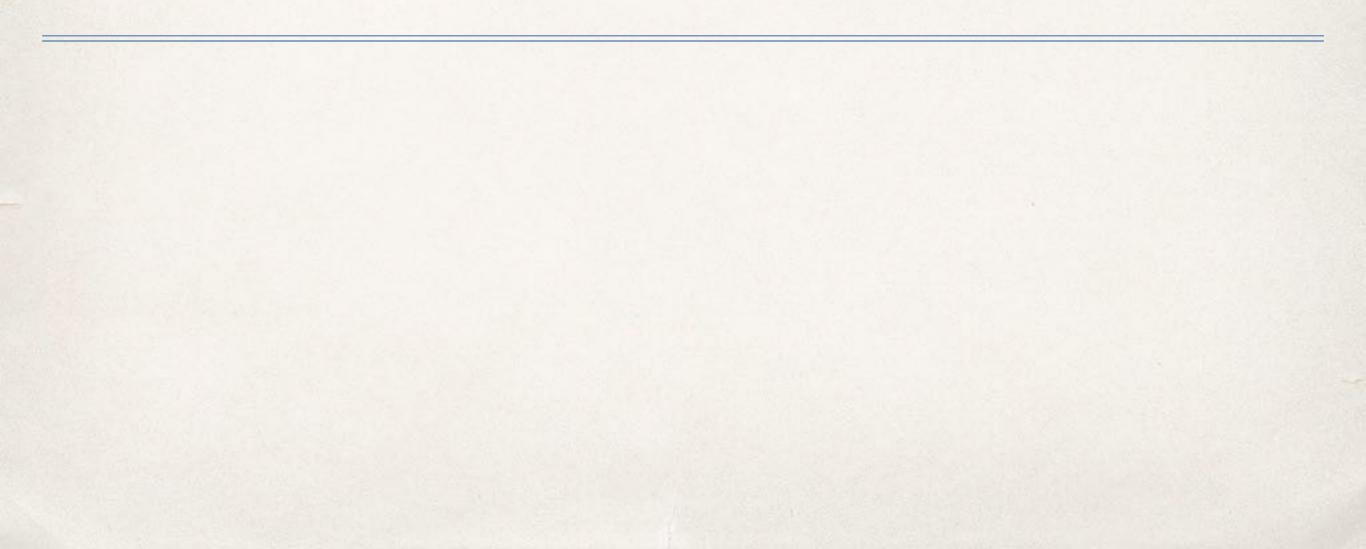
Dark sectors and GWs. Classify sectors with 1st order PT and compute their GW signatures. Map onto DM models.





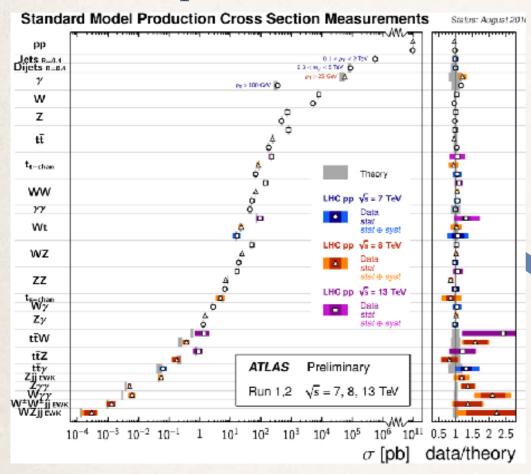
Regions: different dark sectors Arrow: ~ region LISA (1yr)

Back to the LHC: Direct versus indirect searches



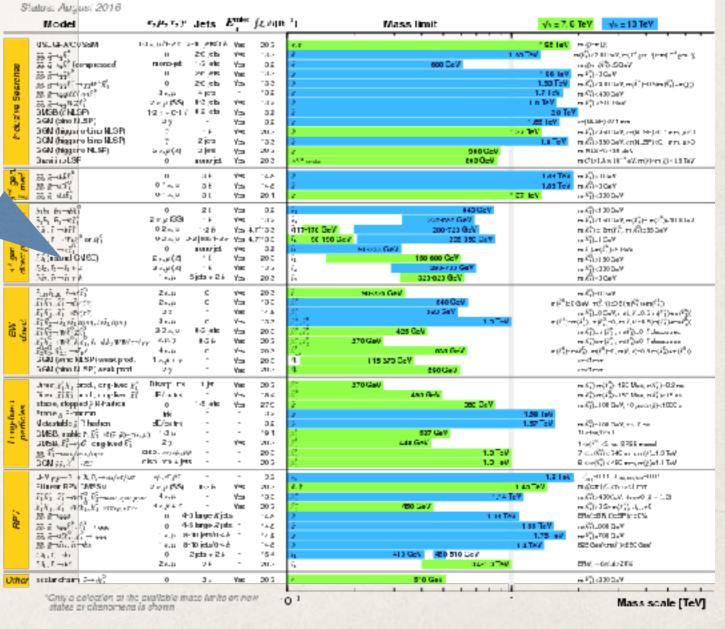
Direct searches for new phenomena

consistency of data vs SM predictions

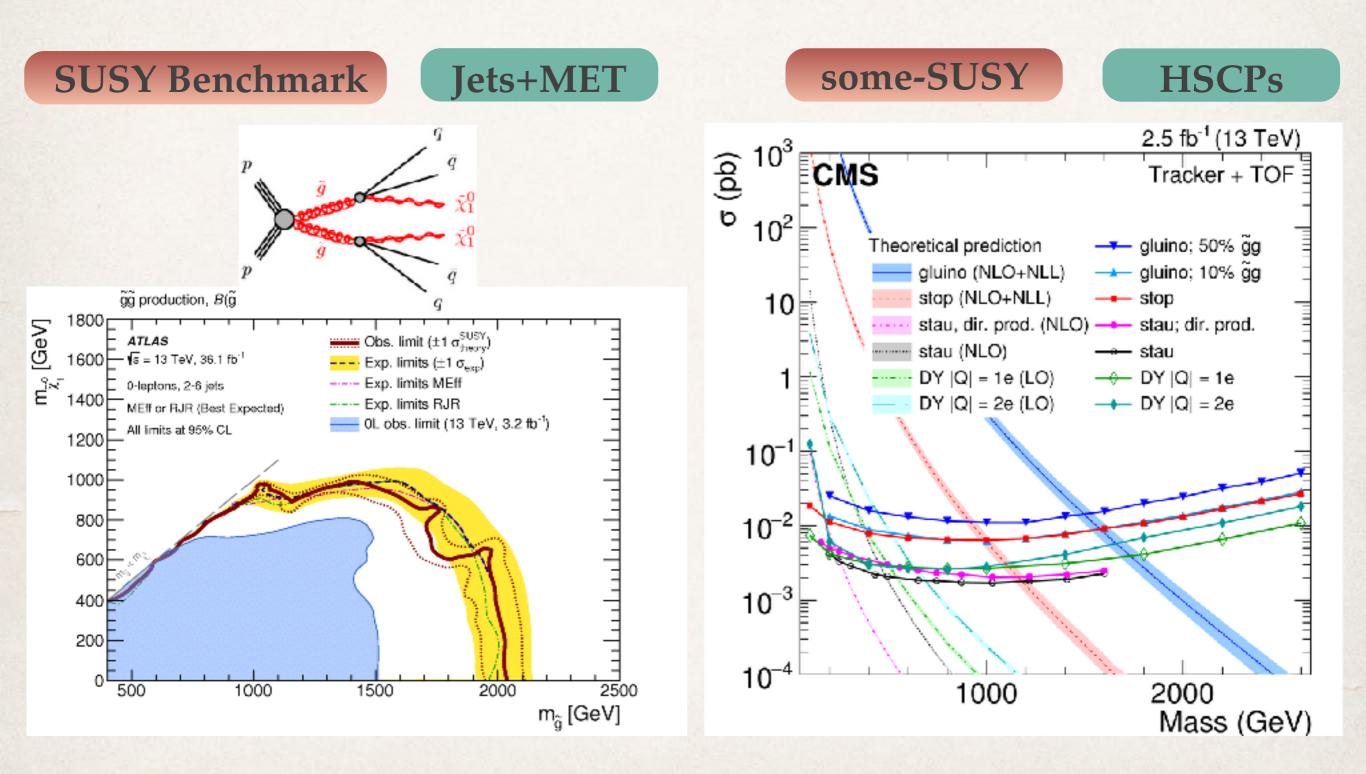


Interpretation in models: exclusion regions

ATLAS SUSY Searches* - 95% CL Lower Limits



Coloured states to the very exotic



Indirect searches

Focus on SM particles' behaviour precise determination of couplings and kinematics comparison with SM, search for deviations

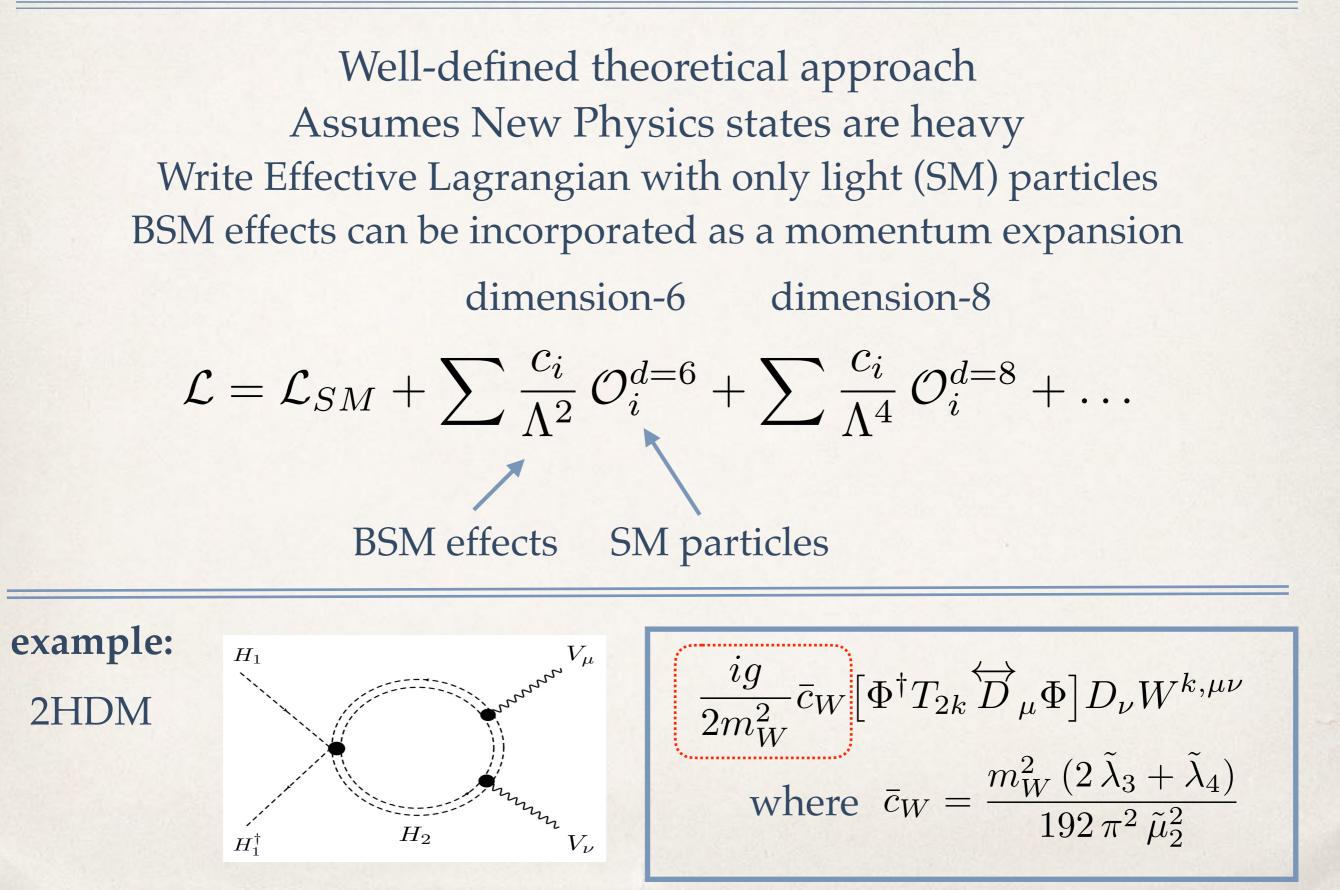
Indirect searches using the Higgs since 2012, relatively new Higgs as a window to NP expect deviations in its behaviour Run2 data and beyond precision Higgs Physics e.g. Anomalous trilinear gauge couplings, aka TGCs V2q **V1** q **V3** 5 GeV ATLAS Preliminary 800 (s = 8 TeV, Ldt = 20.3 fb⁻¹ other diboson Events / e[±]v u[∓]v channel 700È 600 500 400300 20080 120 60 100 140 p. (leading lepton) [GeV]

LEP, Tevatron, LHC

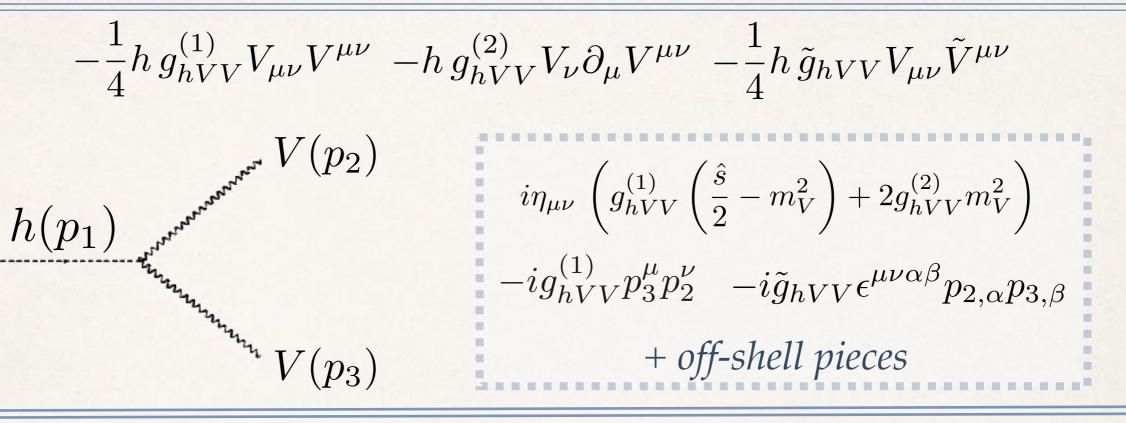
Casting a wide net & the new SM

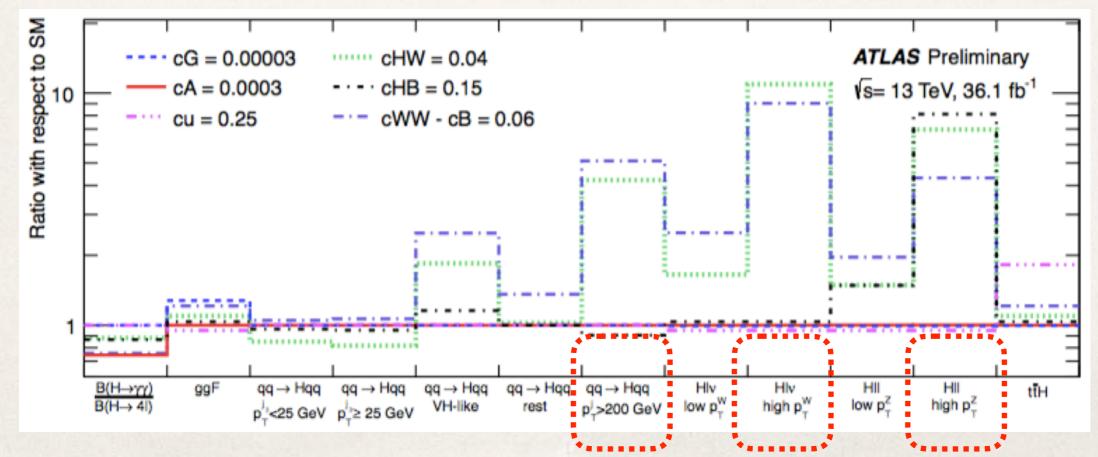


EFT approach

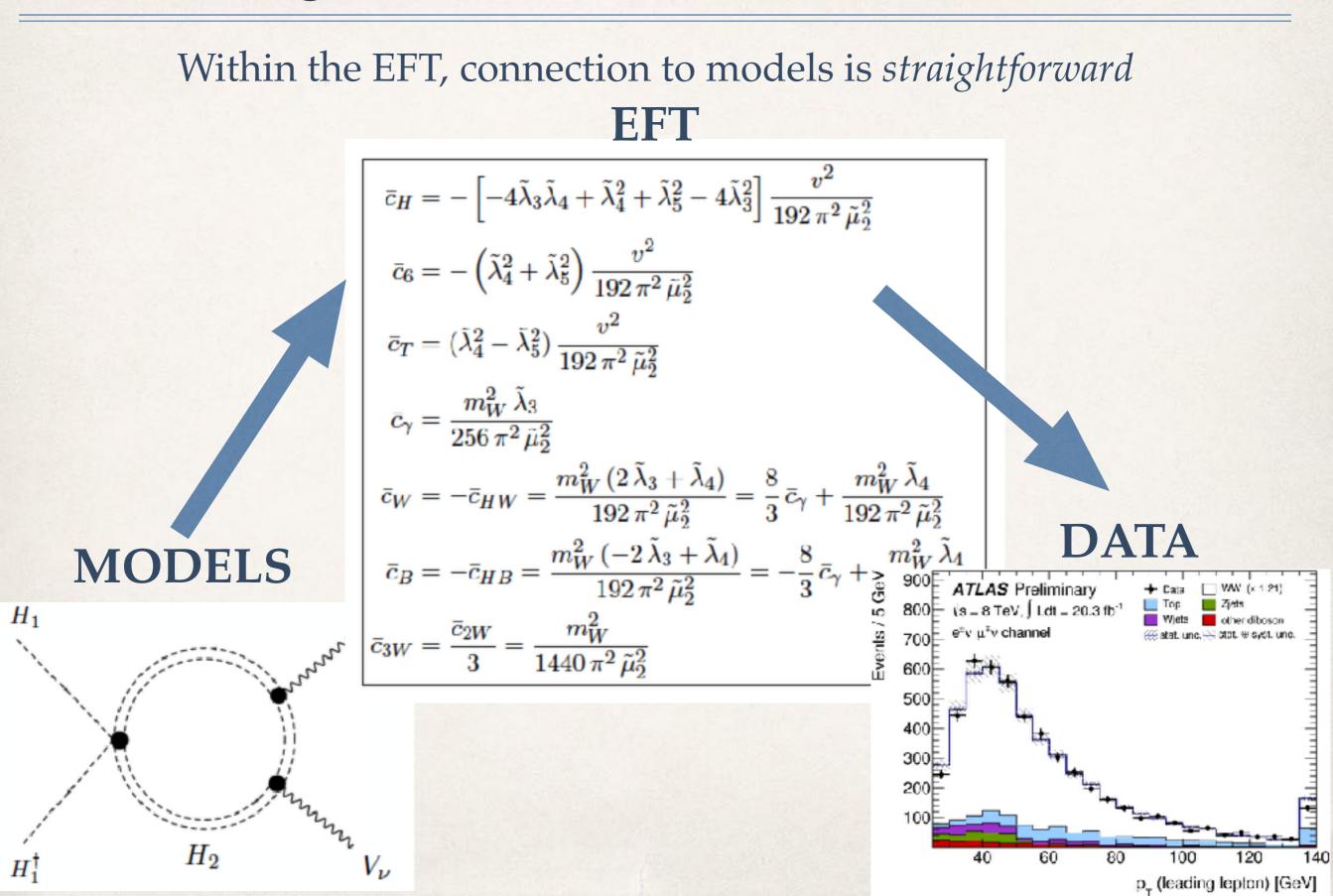


EFT and differential information





Matching to UV theories



SMEFT recent results ellis, murphy, vs, you. 1803.03252

In this work:

Use EWPT, Higgs and diboson data, incl use STXS Assume linear EWSB, CP-conservation and MFV Present results in Warsaw and SILH bases, 20 operators Matching to simplified UV models

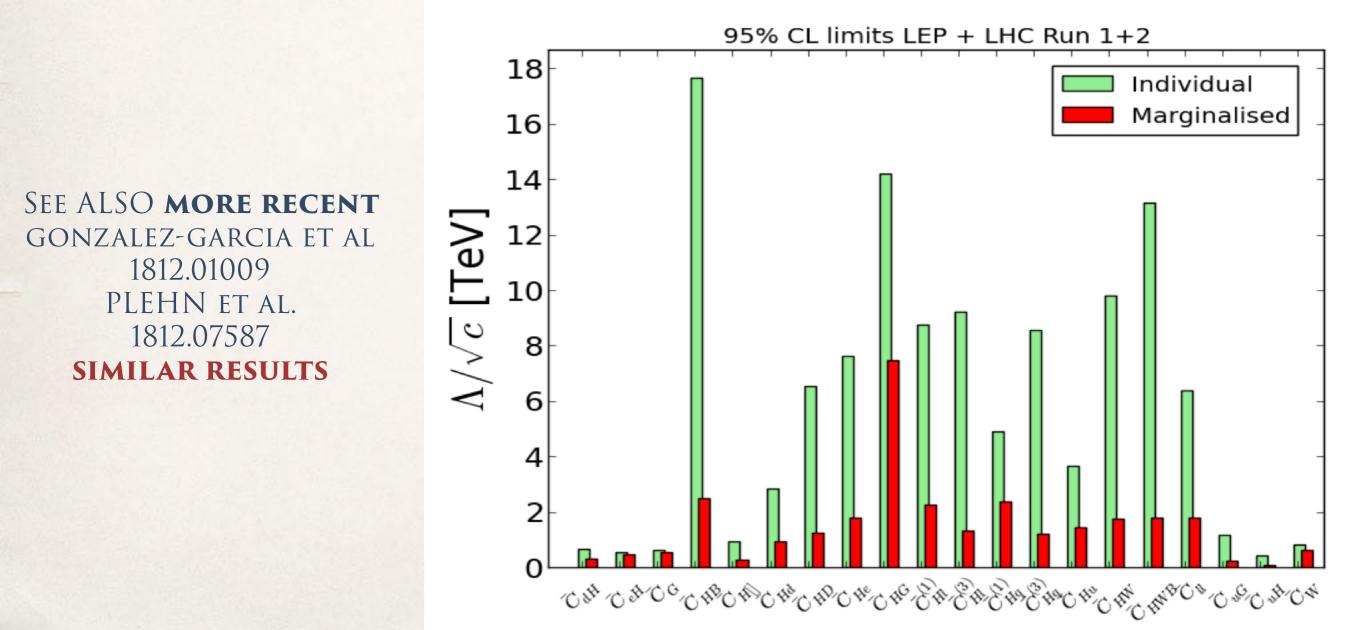
$$\begin{split} \mathcal{L}_{\rm SMEFT}^{\rm Warsaw} \supset \frac{\bar{C}_{Hl}^{(3)}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu}^{I} H) (\bar{l}\tau^{I} \gamma^{\mu} l) + \frac{\bar{C}_{Hl}^{(1)}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu} H) (\bar{l}\gamma^{\mu} l) + \frac{\bar{C}_{ll}}{v^2} (\bar{l}\gamma_{\mu} l) (\bar{l}\gamma^{\mu} l) \\ &+ \frac{\bar{C}_{HD}}{v^2} \left| H^{\dagger} D_{\mu} H \right|^2 + \frac{\bar{C}_{HWB}}{v^2} H^{\dagger} \tau^{I} H W_{\mu\nu}^{I} B^{\mu\nu} \\ &+ \frac{\bar{C}_{He}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu} H) (\bar{e}\gamma^{\mu} e) + \frac{\bar{C}_{Hu}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu} H) (\bar{u}\gamma^{\mu} u) + \frac{\bar{C}_{Hd}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu} H) (\bar{d}\gamma^{\mu} d) \\ &+ \frac{\bar{C}_{Hq}^{(3)}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu}^{I} H) (\bar{q}\tau^{I}\gamma^{\mu} q) + \frac{\bar{C}_{Hq}^{(1)}}{v^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu} H) (\bar{q}\gamma^{\mu} q) + \frac{\bar{C}_{W}}{v^2} \epsilon^{IJK} W_{\mu}^{I\nu} W_{\nu}^{J\rho} W_{\rho}^{K\mu} \\ \mathcal{L}_{\rm SMEFT}^{\rm Warsaw} \supset \frac{\bar{C}_{eH}}{v^2} (H^{\dagger} H) (\bar{l}eH) + \frac{\bar{C}_{dH}}{v^2} (H^{\dagger} H) (\bar{q}dH) + \frac{\bar{C}_{uH}}{v^2} (H^{\dagger} H) (\bar{q}u \widetilde{H}) \\ &+ \frac{\bar{C}_{G}}{v^2} f^{ABC} G_{\mu}^{A\nu} G_{\nu}^{B\rho} G_{\rho}^{C\mu} + \frac{\bar{C}_{H\Box}}{v^2} (H^{\dagger} H) \Box (H^{\dagger} H) + \frac{\bar{C}_{uG}}{v^2} (\bar{q}\sigma^{\mu\nu} T^A u) \widetilde{H} G_{\mu\nu}^A \\ &+ \frac{\bar{C}_{HW}}{v^2} H^{\dagger} H W_{\mu\nu}^I W^{I\mu\nu} + \frac{\bar{C}_{HB}}{v^2} H^{\dagger} H B_{\mu\nu} B^{\mu\nu} + \frac{\bar{C}_{HG}}{v^2} H^{\dagger} H G_{\mu\nu}^A G^{A\mu\nu} \,. \end{split}$$

e.g. WARSAW

SMEFT recent results Ellis, MURPHY, VS, YOU. 1803.03252

Theory	χ^2	$\chi^2/n_{ m d}$	<i>p</i> -value
SM	157	0.987	0.532
SMEFT	137	0.987	0.528
\mathbf{SMEFT}^{\star}	143	0.977	0.564

SMEFT: 20 deformations SMEFT*: 13 deformations (weakly coupled and renormalizable)



SMEFT recent results Ellis, Murphy, VS, YOU. 1803.03252

Constraints on simple extensions of the SM						
Model	χ^2	$\chi^2/n_{ m d}$	Coupling	Mass / TeV		
SM	157	0.987	-	-		
\mathcal{S}_1	156	0.986	$ y_{\mathcal{S}_1} ^2 = (6.3 \pm 5.9) \cdot 10^{-3}$	$M_{S_1} = (9.0, 49)$		
φ , Type I	156	0.986	$Z_6 \cdot \coseta = -0.64 \pm 0.59$	$M_{arphi} = (0.9, 4.3)$		
Ξ	155	0.984	$ \kappa_{\Xi} ^2 = (4.2 \pm 3.4) \cdot 10^{-3}$	$M_{\Xi} = (12, 35)$		
N	155	0.978	$ \lambda_N ^2 = (1.8 \pm 1.2) \cdot 10^{-2}$	$M_N = (5.8, 13)$		
\mathcal{W}_1	155	0.984	$\left \hat{g}^{\phi}_{\mathcal{W}_1} ight ^2 = (3.3 \pm 2.7) \cdot 10^{-3}$	$M_{\mathcal{W}_1} = (4.1,\ 13)$		
E	156.9	0.993	$ \lambda_E ^2 = (2.0 \pm 9.7) \cdot 10^{-3}$	$M_E=(9.2,\infty)$		
Δ_3	156	0.990	$ \lambda_{\Delta_3} ^2 = (0.8 \pm 1.1) \cdot 10^{-2}$	$M_{\Delta_3}=(7.3,\infty)$		
Σ	156.7	0.992	$ \lambda_{\Sigma} ^2 = (0.9 \pm 2.0) \cdot 10^{-2}$	$M_{\Sigma}=(5.9,\infty)$		
Q_5	156	0.990	$ \lambda_{Q_5} ^2 = 0.08 \pm 0.10$	$M_{Q_5}=(2.4,\infty)$		
T_2	156.8	0.992	$ \lambda_{T_2} ^2 = (2.0 \pm 5.1) \cdot 10^{-2}$	$M_{T_2}=(3.8,\infty)$		
S	157	0.993	$ y_{\mathcal{S}} ^2 < 0.32$	$M_S > 1.8$		
Δ_1	157	0.993	$ \lambda_{\Delta_1} ^2 < 5.7 \cdot 10^{-3}$	$M_{\Delta_1} > 13$		
Σ_1	157	0.993	$ \lambda_{\Sigma_1} ^2 < 7.3 \cdot 10^{-3}$	$M_{\Sigma_1} > 12$		
U	157	0.993	$\left \lambda_U ight ^2 < 2.8\cdot 10^{-2}$	$M_{U} > 6.0$		
D	157	0.993	$ \lambda_D ^2 < 1.4 \cdot 10^{-2}$	$M_D > 8.4$		
Q_7	157	0.993	$ \lambda_{Q_7} ^2 < 7.7 \cdot 10^{-2}$	$M_{Q_7} > 3.6$		
T_1	157	0.993	$ \lambda_{T_1} ^2 < 0.13$	$M_{T_1} > 3.0$		
\mathcal{B}_1	157	0.993	$\left \hat{g}^{\phi}_{\mathcal{B}_1} ight ^2 < 2.4 \cdot 10^{-3}$	$M_{{\mathcal B}_1}>21$		

Classification by DE BLAS, CRIADO, PEREZ-VICTORIA, SANTIAGO 1711.10391

EFT precision—next steps

incorporate higher-order QCD and EW effects

quantify higher-order EFT effects (dimension-8)

Lots of progress on this front, some projects involved in

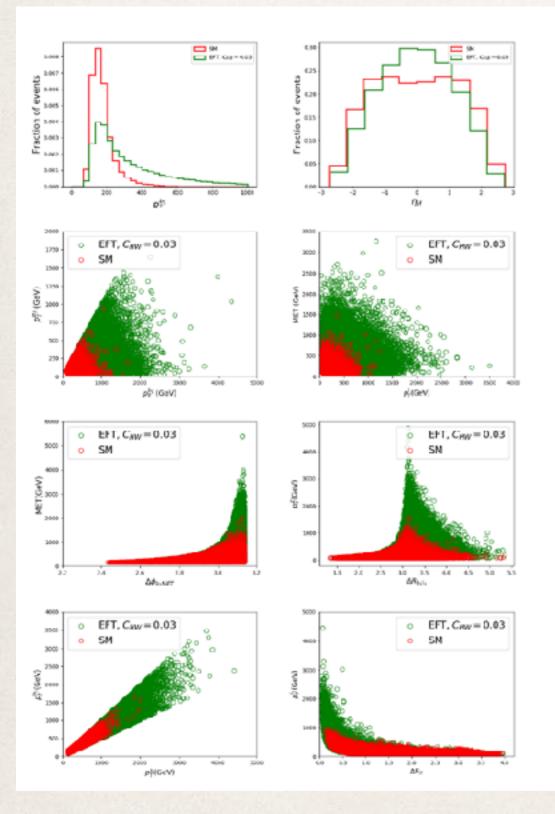
NLO QCD MC

DIMENSION-EIGHT

POWHEG-BOX MIMASU, VS, WILLIAMS. 1512.02572 aMC@NLO DEGRANDE, FUKS, MAWATARI, MIMASU, VS. 1609.04833 NEW: CP-VIOLATING TERMS— REQUEST Feynrules—> UFO—> aMC@NLO HAYS, MARTIN, VS, SETFORD. 1808.00442 Warsaw—>Other using Rosetta MIMASU ET AL. 1508.05895

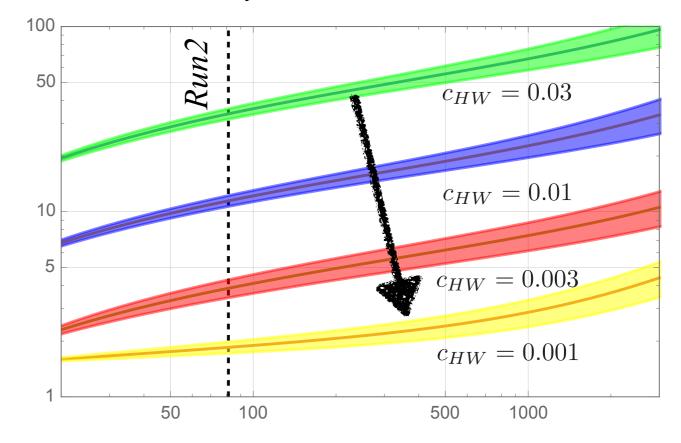
incorporate these tools to the experimental analyses

Next direction: Machine Learning



Capture *subtle* details in "images" **supervised** or **anomaly** detection lots of activity in the last months this is where the cutting-edge is

Asimov significance vs Luminosity systematics 50%

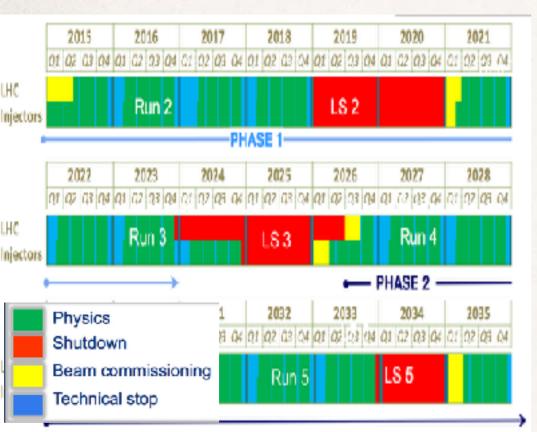


FREITAS, KAUR AND VS. 1902.05803

Experiments keep coming in: There is a lot to explore ahead of us

For the LHC, this is just the beginning

HL-LHC (High-Luminosity) LHC approved, to deliver 3000 inverse fb of data. Funding ensured until 2035.





LHC hopefuls gains from more data and better understanding of the environment

Testing non-standard kinematic features Reaching high-precision in Higgs physics Searches for invisible particles (monoX) Blind spots (DV, disap. tracks, quirks)

and, of course, **FLAVOUR** with Belle-II, NA62 complementing LHCb

Smaller experiments may be key

Narrower focus BUT cheaper, shorter time-scale develop creative experimental techniques often enlarge the initial physics focus

LZ	ACT	BICEP3	
DUNE ADMX	g-2 Icecube	Darwin Qbic HIP	ANTARES
NEXT nuSTORM	MADMA)	T2K Mu2e ATHUSLA	Euronu Moedal

Conclusions

- Here we are, looking for a way to advance our understanding of nature, to reach discovery
- Scaling back from an ambitious program to find *the* theory of everything.
 Facing the challenges / opportunities that more data brings
- Use of simplified models to organize/interpret searches, less model biased, and suitable to complementarity studies. Yet theoretical advances require more than simplified models, asking difficult questions from model building
- Keeping at the edge of the interpretation of data: bringing many towards precision (akin to SM) and to Artificial Intelligence techniques (NNs and the likes), but we should not lose track of our core mission:

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Understanding Nature (and having fun on the way!)

And what about the cool/crazy stuff?

Dark Energy and its interaction with us Alternatives to space-time symmetries (e.g. emergent gravity) Very light dark matter (new exp techniques) Dark moments in the Universe's history, pre-BBN Connections between IR and UV physics, e.g. BHs

We need to *challenge* the well-stablished paradigms, may be quickly ruled out but one **always** learn something new from these explorations