

Nature loves us, loves us not...



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IFT Madrid 11/2/2016



PLAN



1) Loves us not: **no-show** @ LHC

∞ Relaxions

2) Loves us: **present(s)** @ LHC

PLAN



1) Loves us not: **no-show @ LHC**

∞ Relaxions

2) Loves us: **present(s) @ LHC**

∞ Dilaton

1) Cosmological Relaxation mechanisms

A long long time ago,
in a far far away place,
the scale of the EW forces
(the Higgs mass)
were suppressed... by...
a cosmic conspiracy...

Motivation



- ❧ A new type of solution to the Hierarchy Problem has been identified (Graham, Kaplan & Rajendran '15)
- ❧ It doesn't require new particles @ TeV energies
- ❧ Instead, a fantastic cosmological history

Idea



Relaxation mechanism

Abbot '85

Dvali & Vilenkin'2004

Graham, Kaplan, Rajendran '15

$$M^2 H^2 \rightarrow M^2(\phi) H^2$$

ϕ rolls, scans, and stops when M small

Stopping can be
linked to EWSB

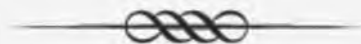
$$M^2 < 0$$

Cosmological Relaxation



$$V_{SM+PQ}(\phi, H) = M^2 H^2 + \lambda H^4 + \underbrace{\epsilon H \Lambda^3 \cos(\phi/f)}_{m_q \langle \bar{q}_R q_L \rangle}$$

Cosmological Relaxation



$$V_{SM+PQ}(\phi, H) = M^2 H^2 + \lambda H^4 + \varepsilon H \Lambda^3 \cos(\phi/f)$$

(Graham, Kaplan, Rajendran'15)

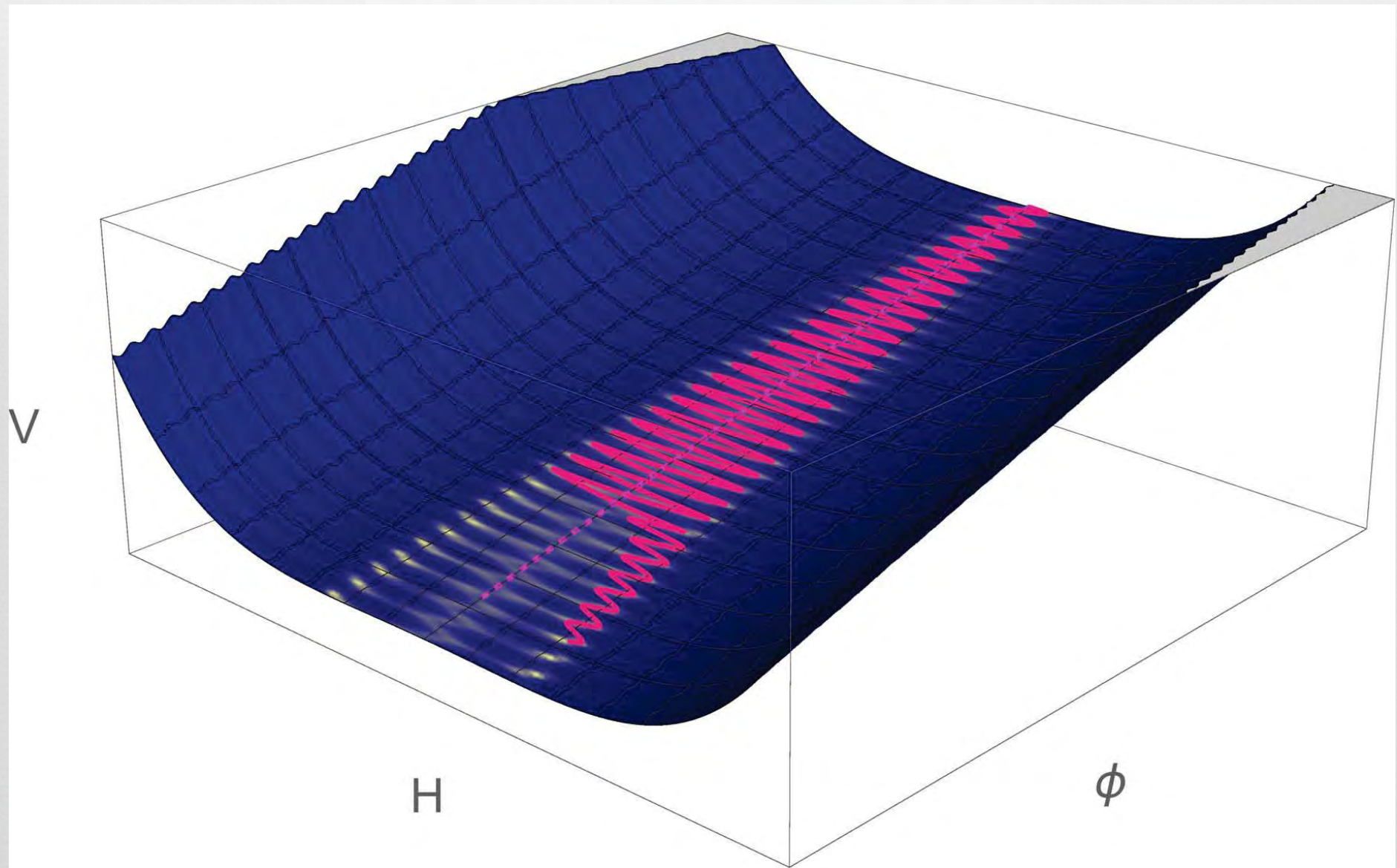
$$V_{GKR} = \underbrace{M^2(\phi)}_{\text{scanning}} H^2 + \lambda H^4 + \varepsilon H \Lambda^3 \cos(\phi/f) + \underbrace{V_0(\phi)}_{\text{rolling}}$$

Technically natural !!

$$V_0(\phi) = \Lambda^4 F(g\phi/\Lambda)$$

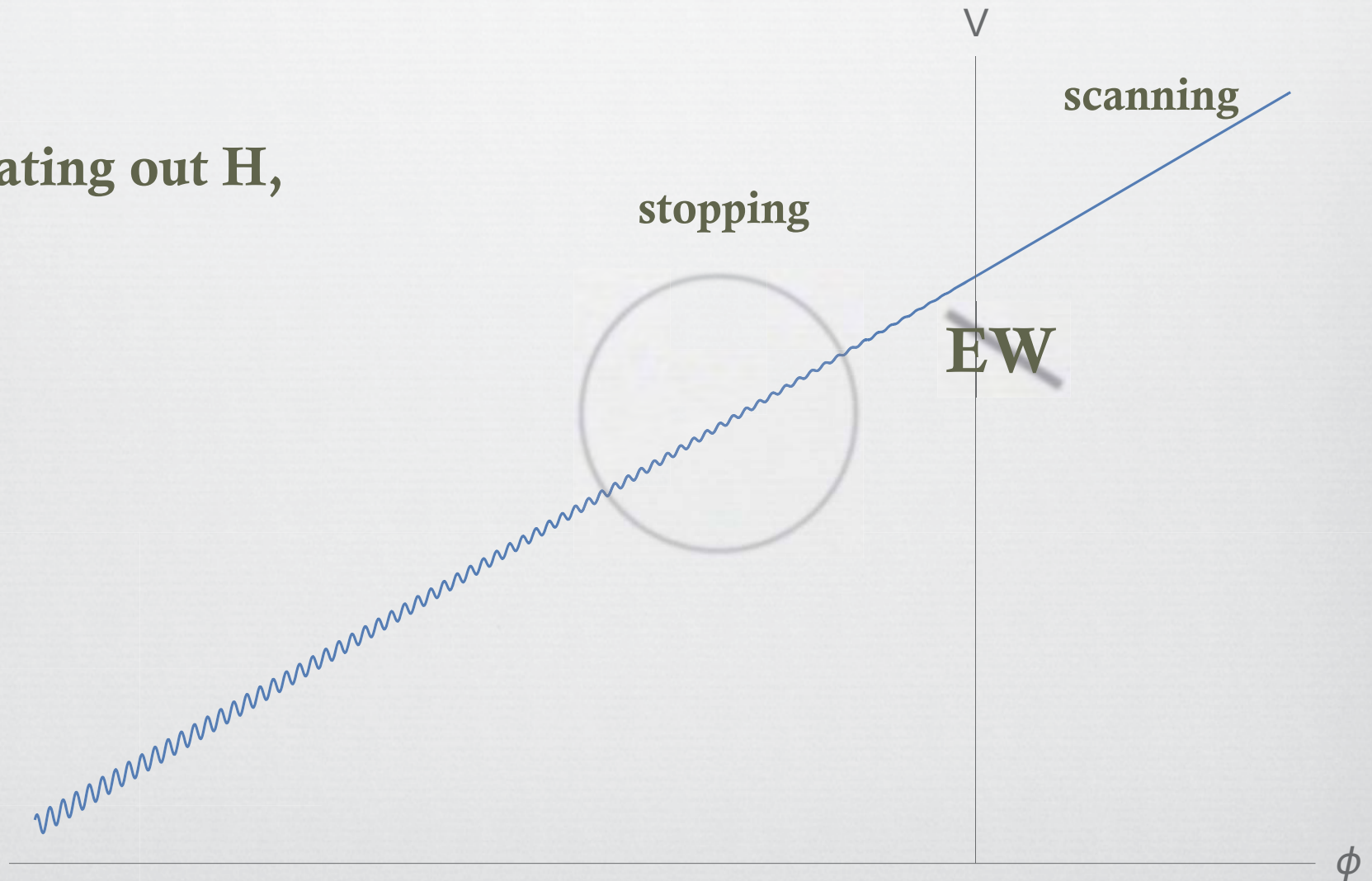
$$M^2(\phi) = \Lambda^2 F(g\phi/\Lambda)$$

GKR Relaxion



GKR Relaxion

Integrating out H,



GKR Relaxion

∞ conditions:

solve HP

∞ full scan $\Rightarrow \Delta\phi \geq \Lambda / g$

∞ many vacua $(\Delta M_H \ll \Lambda) \Rightarrow gf \ll \Lambda$

∞ right ew scale $\Rightarrow g \sim \epsilon v/f$

inflation

∞ friction \Rightarrow inflation (maybe more options...)

∞ subleading to inflaton $\Rightarrow \Lambda^4 < H_I^2 M_P^2$

∞ classical motion $\Rightarrow H_I^3 < g\Lambda^3$

$$\Lambda^3 < g M_P^3$$

GKR Relaxion

∞ results:

∞ raises NP scale to $\Lambda \leq 30\text{-}1000 \text{ TeV}$

∞ Long inflation epoch $N_{\text{efolds}} \sim 10^{40}$

∞ Long field excursion $\Delta\phi \gg \Lambda$

∞ low inflation scale $g^{1/3} \Lambda^2 M_P^2$

(Reminder) Ingredients below $\Lambda = \{ \text{SM} + \text{QCD axion} + \text{expl. breaking of PQ} + \text{few assumptions on inflation} \}$

Cosmological Higgs-Axion Interplay – CHAIN

Cosmological Higgs-Axion Interplay for a Naturally Small
Electroweak Scale

J. R. Espinosa, C. Grojean, G. Panico, A. Pomarol, O. Pujolàs, and G. Servant
Phys. Rev. Lett. **115**, 251803 – Published 16 December 2015

CHAIN



$$V(H, \phi) = M^2(\phi) H^2 + \lambda H^4 + V_0(\phi) + H^n \Lambda^{4-n} \cos(\phi / f)$$

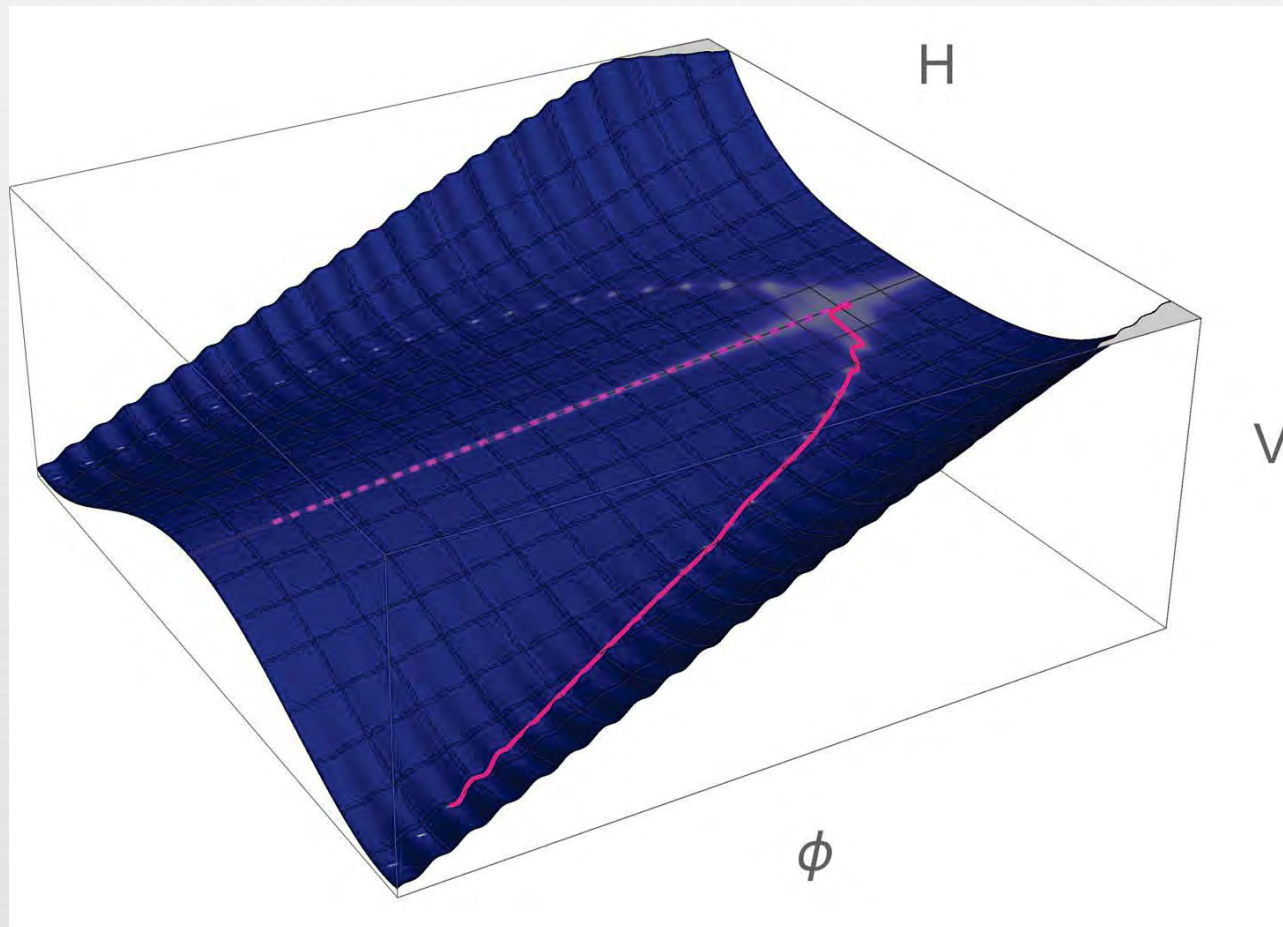
scanning

driving

stopping

$$V_{\text{stopping}} = |H|^2 \Lambda^2 \cos(\phi / f) \quad ??$$

CHAIN



$$V_{stopping} = H^2 \Lambda^2 \cos(\phi / f)$$

CHAIN



$$V(H, \phi) = M^2(\phi) H^2 + \lambda H^4 + V_0(\phi) + H^n \Lambda^{4-n} \cos(\phi / f)$$

scanning

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$$V_{\text{stopping}} = |H|^2 \Lambda^2 \cos(\phi / f) \quad ??$$

CHAIN



$$V(H, \phi) = M^2(\phi) H^2 + \lambda H^4 + V_0(\phi) + H^n \Lambda^{4-n} \cos(\phi / f)$$

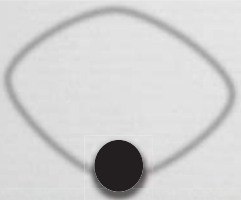
scanning

driving

stopping



$$V_{stopping} = |H|^2 \Lambda^2 \cos(\phi / f) \quad ??$$



$$V_{stopping} = (|H|^2 \Lambda^2 + \Lambda^4) \cos(\phi / f) \quad !!$$

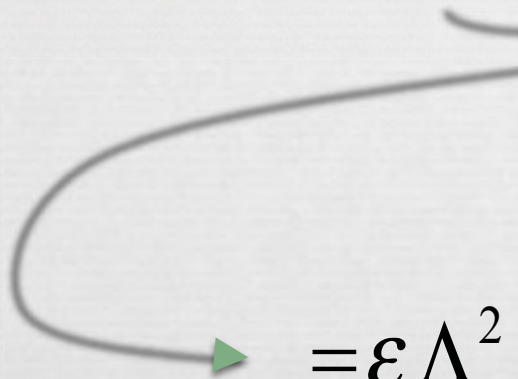
CHAIN



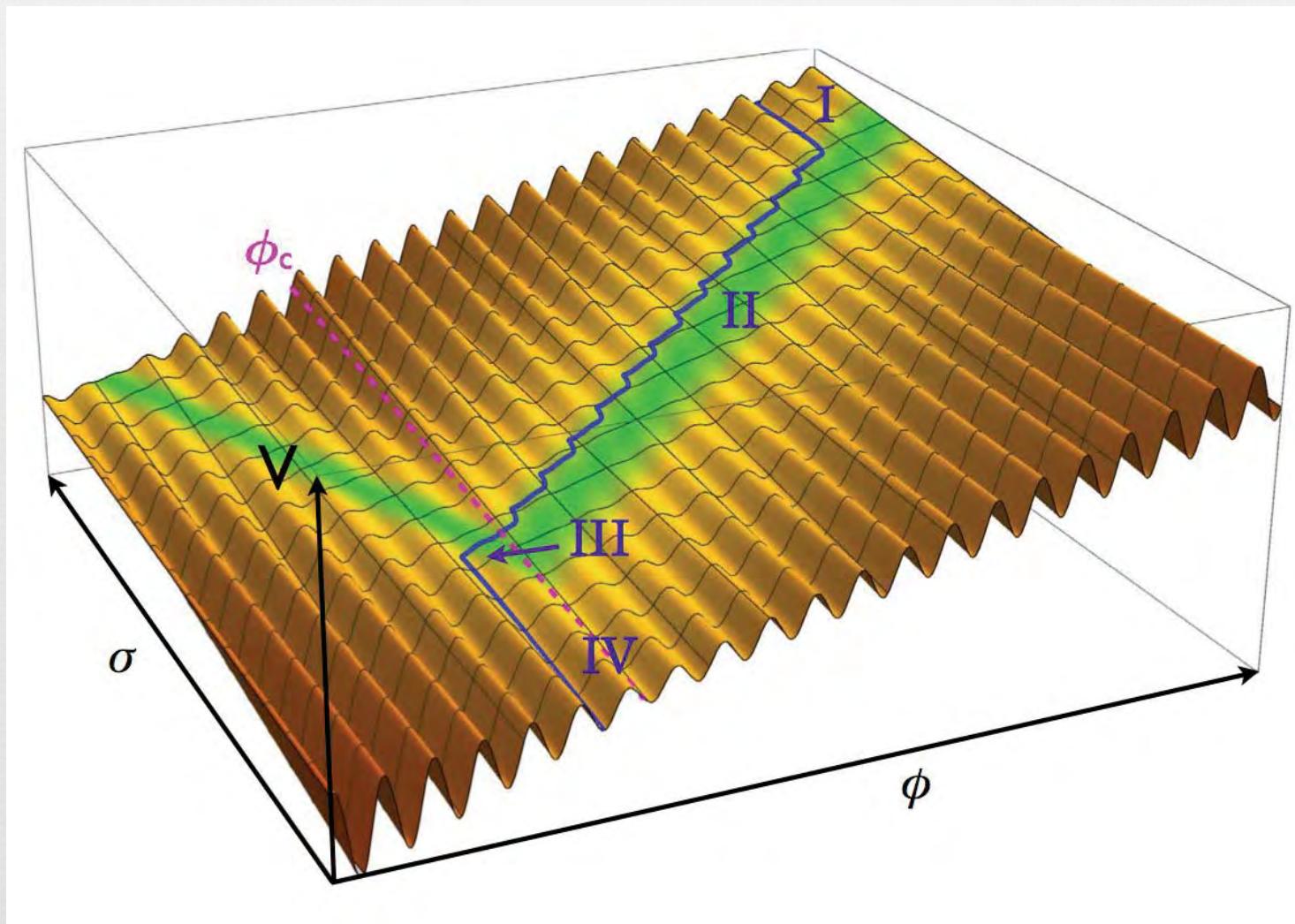
Resolution:

Technically natural !!

$$V(H, \phi, \sigma) = M^2(\phi) H^2 + \lambda H^4 \\ + A(H, \phi, \sigma) \cos(\phi/f) + V_0(\phi, \sigma)$$


$$= \varepsilon \Lambda^2 \left(\Lambda^2 + |H|^2 + g\Lambda\sigma + g\Lambda\phi + \dots \right)$$

CHAIN



CHAIN



∞ 2 light scalars,

ϕ

axion-like

$m_\phi \sim$ sub-GeV-ish

moderate weakly coupled (thermally produced)

impact on BBN

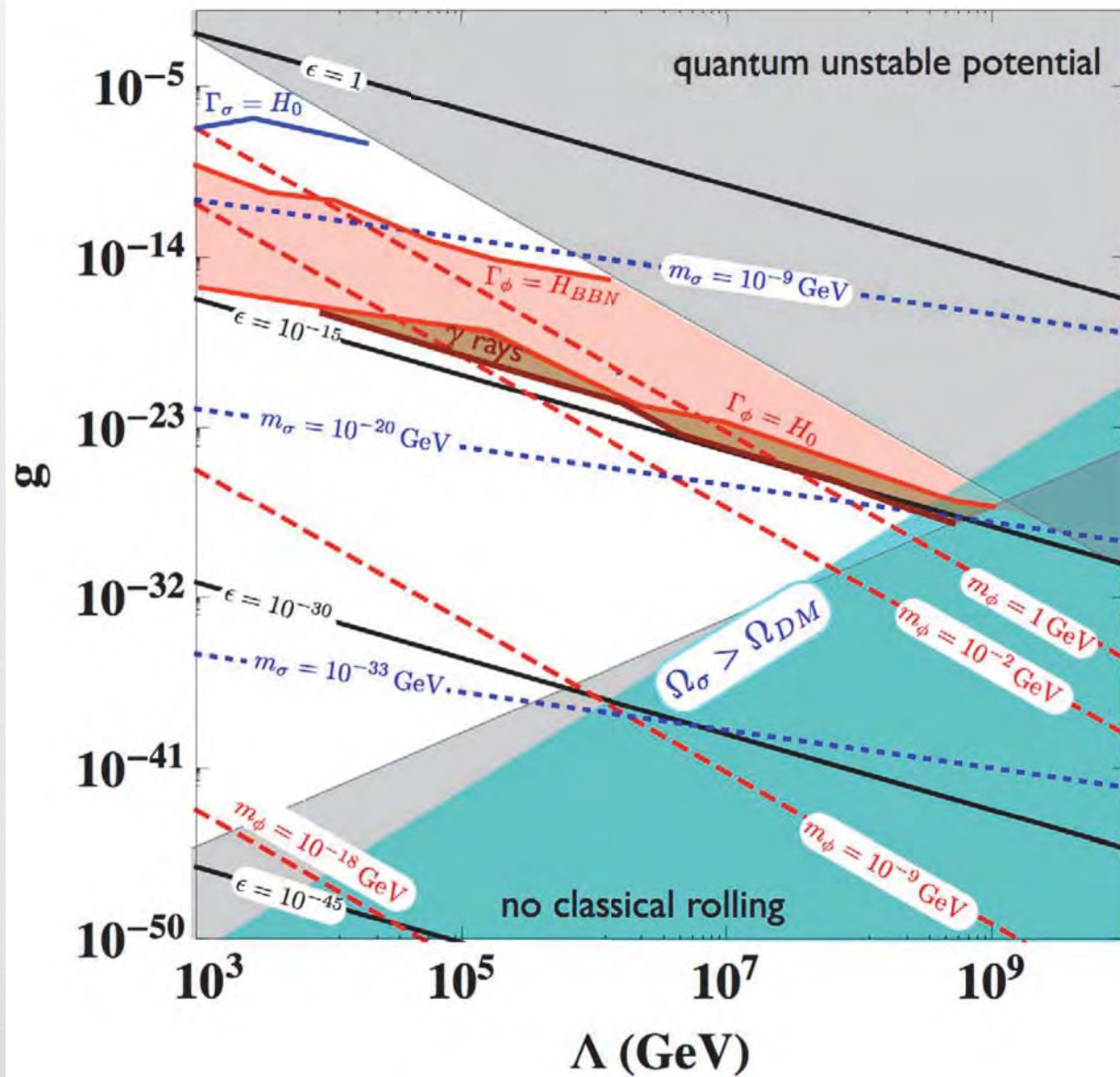
σ

ultra-light scalar, ultra weakly coupled

DM candidate (from ‘vac. misalignment’)

∞ No new particles until $10^5 - 10^9$ GeV

CHAIN



**BBN
constraints
can be
evaded if**

$$T_{reh} < m_H$$

Cosmological Relaxation



- ∞ Can raise the New Physics (NP) scale to $\sim 10^9$ GeV
- ∞ \Rightarrow *Partial* Solution to the Hierarchy problem
- ∞ \Rightarrow No new particles @ EW scale, but implications in cosmo
- ∞ Only remnant: 2 light and weakly coupled scalar singlets
- ∞ Long excursions/periods a problem?

Cosmological Relaxation



∞ Long excursions/periods a problem?

∞ Weak Gravity Conjecture^(hep-th/0601001) suggests $f < M_p$

Gupta Komargodski Perez Ubaldi 1509.00047

∞ However, the argument is naïve/quick

Choi & Im 1511.00132

Kaplan & Rattazzi 1511.01827

Ibanez Montero Uranga Valenzuela 1512.00025

Monodromy in Large N QCD

Witten 80's

Cosmological Relaxation



∞ Long excursions/periods a problem?

Exponentially large field periods can be realized in models with:

- order-few fields
- order-few charge assignments

$$\Delta\phi \sim e^N f$$

Choi & Im 1511.00132

Kaplan & Rattazzi 1511.01827

Open questions (?):

tension with Quantum Gravity/Weak Gravity Conjecture?

From EFT perspective, I see

- i) no problem ii) no deviation from QFT framework, and**
- iii) WGC and alike are *assumptions* about QG**

Cosmological Relaxation



∞ Long excursions/periods a problem?

It seems not (at least yet).

Rather, it makes relaxion-ology a probe of QG.

2) What if Nature loves us?

A nice present would be...

2) What if Nature loves us?

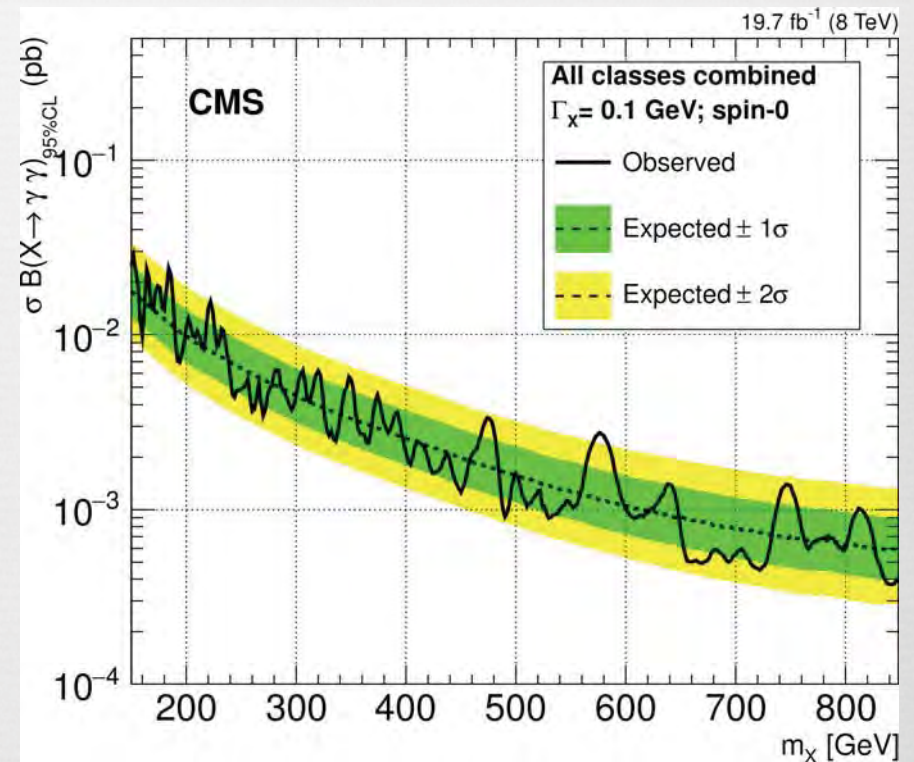
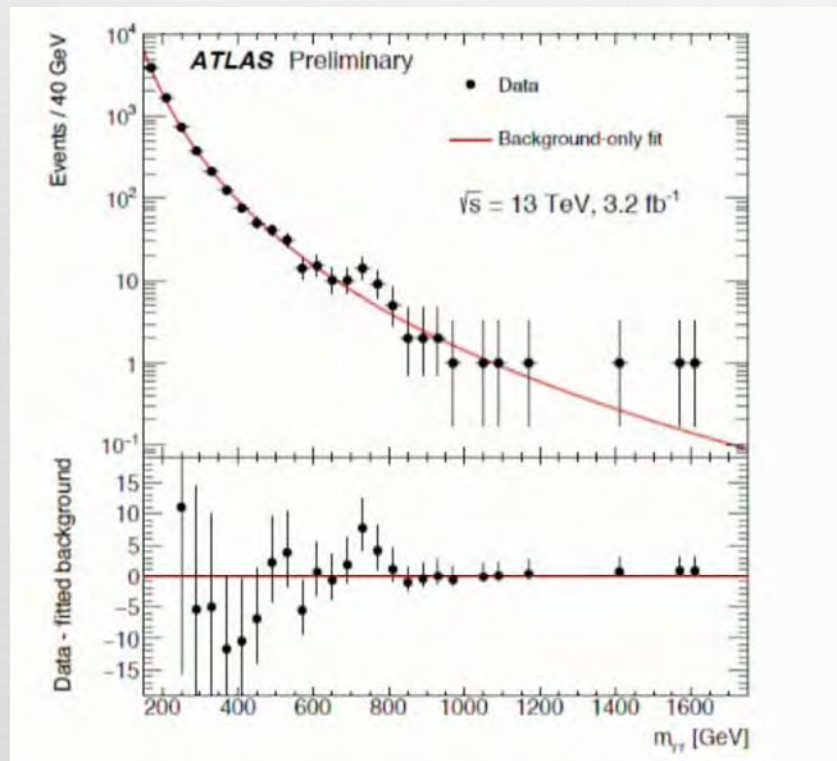
A nice present would be...

a dilaton !

“Naturally light dilatons”

w/ Eugenio Megías and Mariano Quirós

Diphoton excess



Dilaton?

- ⌘ Goldstone Boson of Spontaneously Broken Conformal Invariance (SBCI)
- ⌘ Contino-Pomarol-Rattazzi '2010 showed how to realize naturally SBCI and a light dilaton
- ⌘ Now we have examples of such beasts arising from (holographic) strong dynamics

Dilaton?

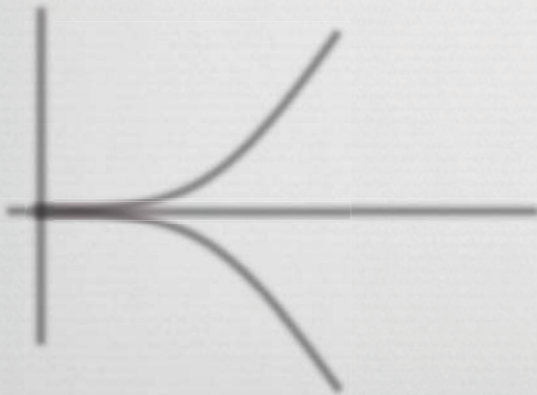
∞ EFT of SBCI

$$\begin{aligned}\phi &\rightarrow s\phi & \langle\phi\rangle &\neq 0 \\ x &\rightarrow s^{-1}x\end{aligned}$$

$$L = \phi^4 F\left[\frac{(\partial\phi)^2}{\phi^4}\right] = (\partial\phi)^2 - \lambda\phi^4 + \dots$$

Expected to be large (~ 1) \Rightarrow no natural SBCI

Analogous to Cosmological Constant problem



Dilaton?

∞ Simplest example: IR brane in Randall Sundrum

$\phi =$ warp factor at IR brane location

$$\lambda \sim \ell \Lambda_5 - \Lambda_4$$

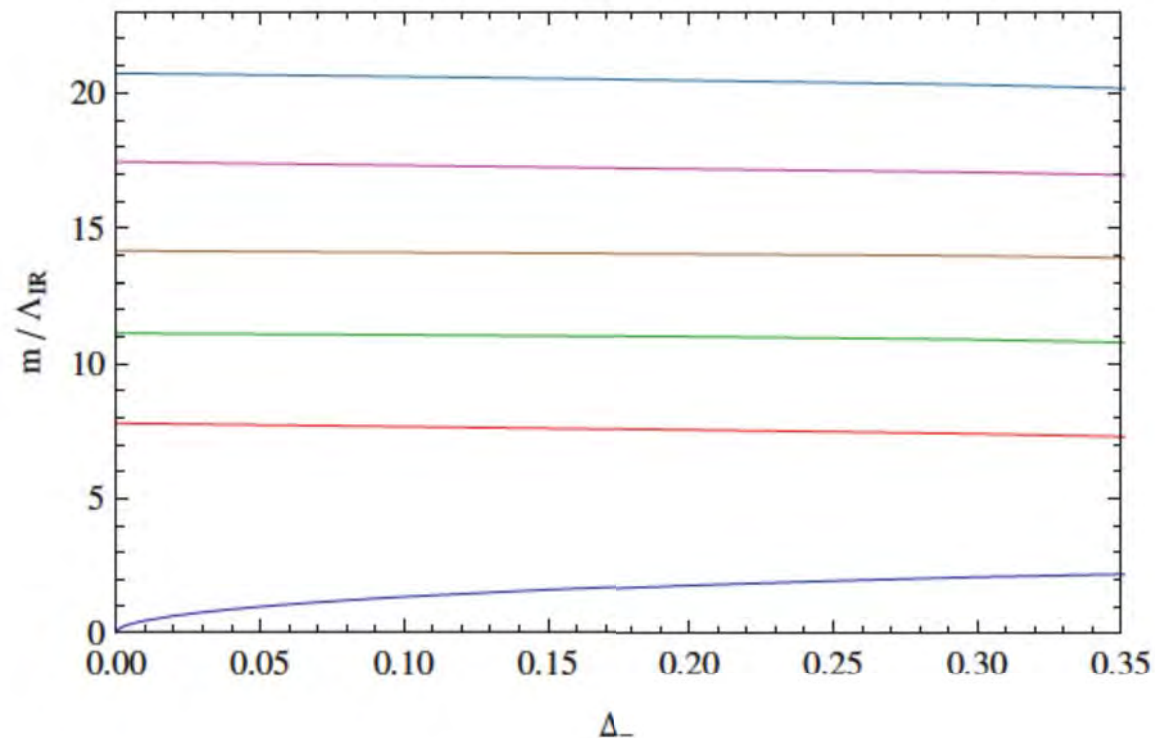
∞ CPR solution: add small explicit breaking
(by *nearly marginal deformation*)

∞ λ gets dynamically suppressed in the ground state
and a (naturally) light dilaton appears

Dilaton?

w/ Eugenio Megías 1401.4998

Also in
1305.3919
1306.4601



$$m_{dil}^2 \sim \beta_{IR} \Lambda_{IR}^2$$

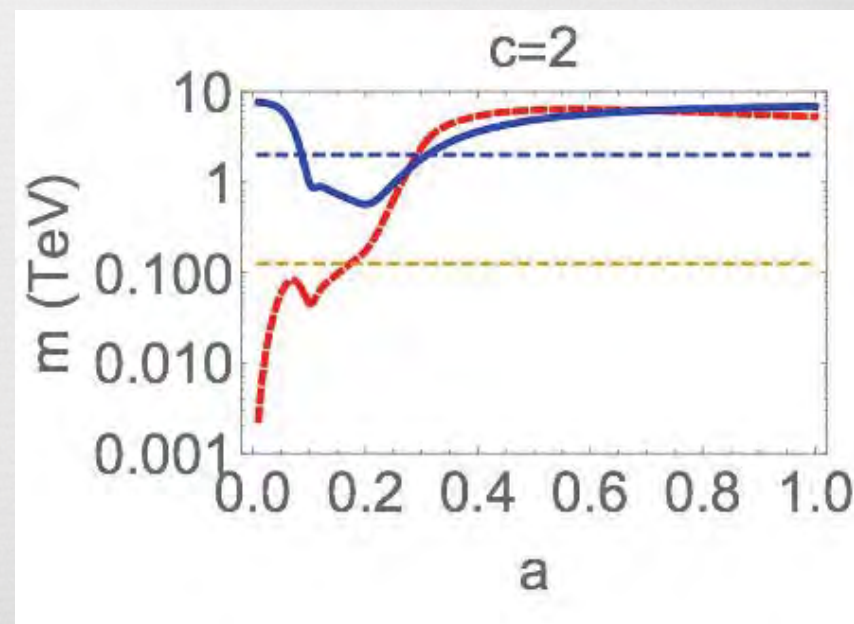
Diphoton as a dilaton?

Explicit & minimal Soft-Wall model with:

- KK resonances at most at 2 TeV
- *light* dilaton ~ 750 GeV
- large enough couplings to $\gamma\gamma$ and gg
- pass EWPT tests

'diphoton-dilaton' is light naturally
stands as a possibility

w/ E. Megías and M. Quirós
1512.06106



Conclusions

- ∞ Interesting times ahead...
- ∞ If she loves us => happy family, lots of children
- ∞ Loves us not =>

Conclusions

∞ Interesting times ahead...

∞ If she loves us => happy family, lots of children

∞ Loves us not => keep working on her, learning
how to lover her
– eventually she'll love us back!

Thanks!

(and apologies for many missing refs)