



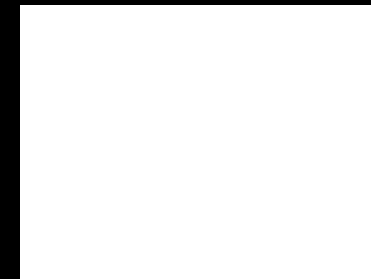
Neutrino Masses & DM candidate in Inverse See saw scenarios

Federica Bazzocchi
Vrije Universitet Amsterdam

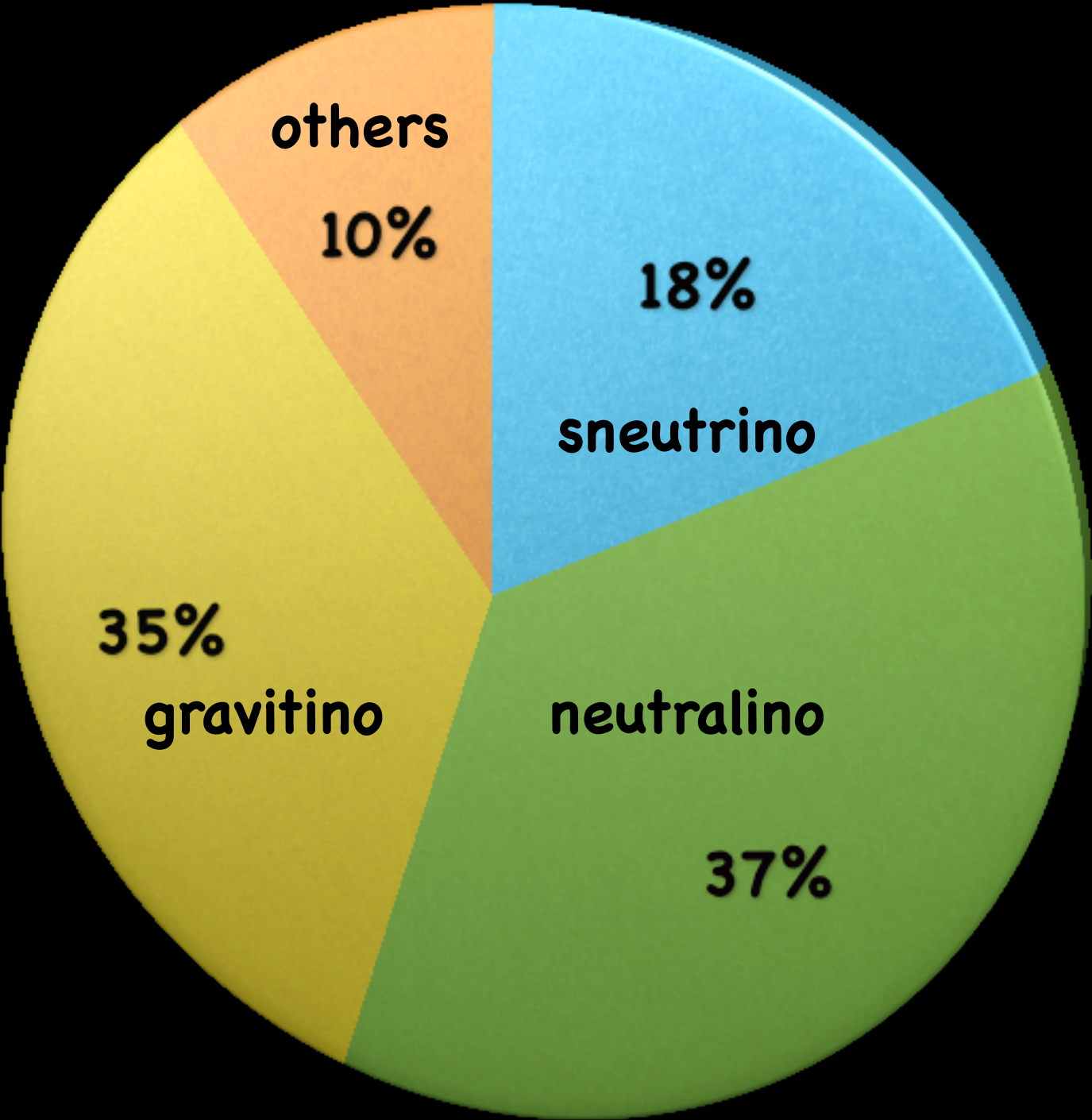
C. Arina, F.B., N. Fornengo, J. Romao, J. Valle PRL101 (2008)

FB,D. Cerdano, C. Munoz, J.F. Valle arXiv:0907.1262

Theoretical miniWorkshop on Dark Matter, Madrid, 16-18/09/9



2008-2009



sneutrino

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- in MSSM no good DM candidate

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from a single sector : both
neutrino masses & CDM

sneutrino as LSP & DM candidate

Arina & Fornengo, JHEP112007029

- ☐ MSSM
- ☒ MSSM + $\hat{N}$ (dirac neutrinos)
- ☐ MSSM + L-viol
- ☒ MSSM + $\hat{N}$ + L-viol (see-saw)

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“ Problems ”

📌 not compatible with mSUGRA universality

📌 scale M_R around TeV $\rightarrow$ unnatural small Y_D

Why

📌 $m_R^2 = m_0^2 + 0.15 m_{1/2}^2$, $m_L^2 = m_0^2 + 0.52 m_{1/2}^2 \rightarrow$ stau no LSP $\rightarrow$ neutralino LSP

📌 low see-saw scale $\rightarrow$ $>$ sterile component $\rightarrow$ $< \sigma_{\tilde{\nu}}$

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any model?

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InMSSM model

inverse see-saw : SM + ν_L^c, S with L (-1,1)

$$M^\nu = \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & 0 & M_{RS} \\ 0 & M_{RS} & \mu_S \end{pmatrix}$$

low scale mechanism

μ_{SS}

$$m_L^\nu \sim m_D \frac{1}{M_{RS}} \mu_S \frac{1}{M_{RS}^T} m_D^T$$

1 TeV

1 keV-MeV

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$$M_{RS} \hat{\nu}^c \hat{S}$$

$$\begin{pmatrix} \mathcal{O}(M_{RS}) \sim \mathcal{O}(\mu) \\ \mu \hat{H}_u \hat{H}_d \end{pmatrix}$$

$$\Phi^\dagger = (\tilde{\nu}_+^*, \tilde{\nu}_+^{c*}, \tilde{S}_+^*, \tilde{\nu}_-^*, \tilde{\nu}_-^{c*}, \tilde{S}_-^*) \quad \left(\begin{array}{cc} \mathcal{M}_+^2 & \mathbf{0} \\ \mathbf{0} & \mathcal{M}_-^2 \end{array} \right)$$

CP-eigenstate basis

$$\mathcal{M}_\pm^2 = \left(\begin{array}{ccc} m_L^2 + D + m_D^2 & m_D + A_{h\nu} v_u + \mu m_D \cot g\beta & m_D M_R \\ m_D + A_{h\nu} v_u + \mu m_D \cot g\beta & m_{\nu^c}^2 + M_R^2 + m_D^2 + m_{\nu^c}^2 & \mu_S M_R \pm B_{M_R} \\ m_D M_R & \mu_S M_R \pm B_{M_R} & m_S^2 + \mu_S^2 + M_R^2 \pm B_{\mu_S} \end{array} \right)$$

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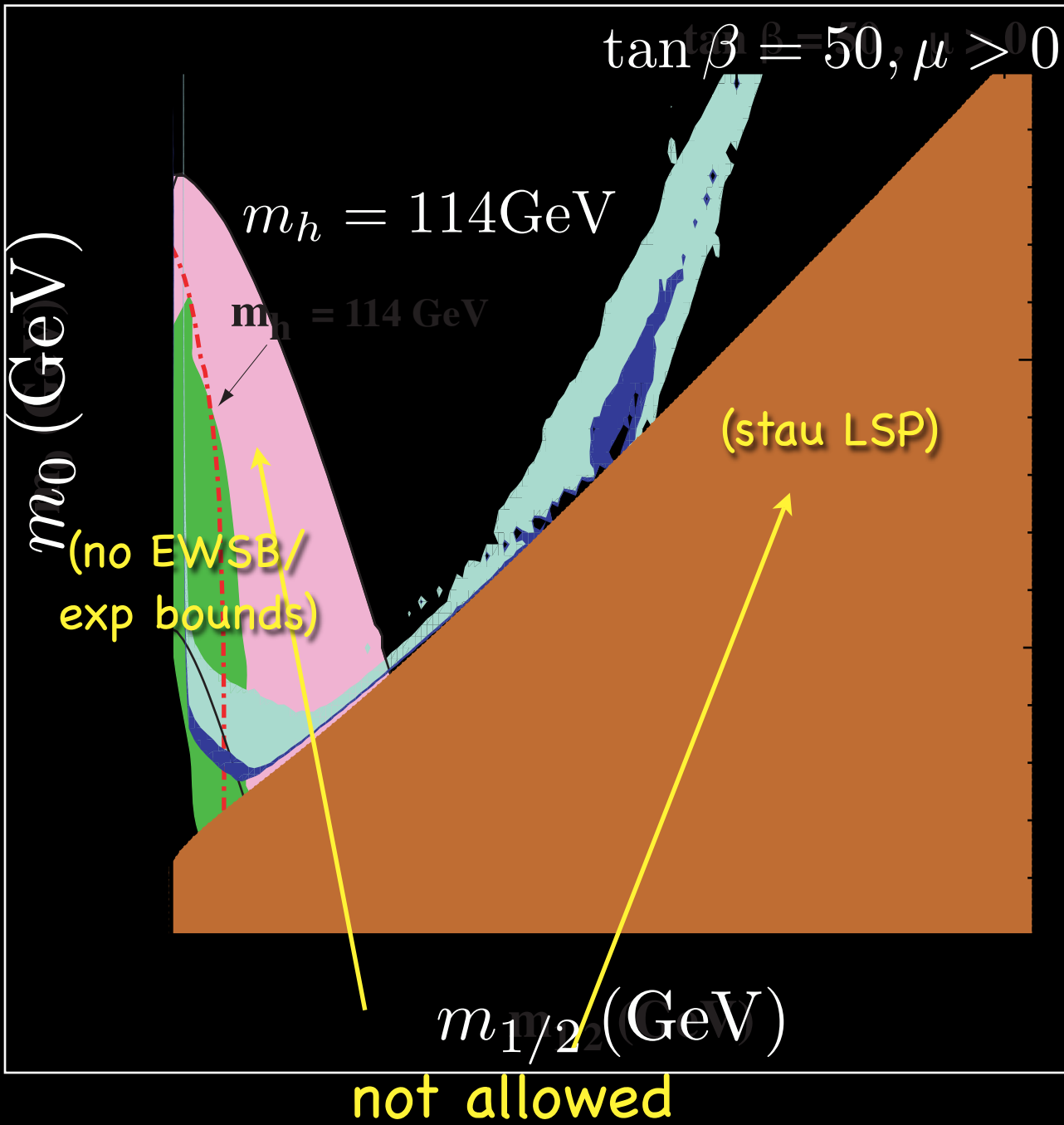
$$\mathcal{M}_\pm^2 = \begin{pmatrix} m_L^2 + D + m_D^2 & m_D + A_{h\nu} v_u + \mu m_D \cot g\beta & m_D M_R \\ m_D + A_{h\nu} v_u + \mu m_D \cot g\beta & m_{\nu^c}^2 + M_R^2 + m_D^2 + m_{\nu^c}^2 & \mu_S M_R \pm B_{M_R} \\ m_D M_R & \mu_S M_R \pm B_{M_R} & m_S^2 + \mu_S^2 + M_R^2 \pm B_{\mu_S} \end{pmatrix}$$

- mSUGRA universality $\rightarrow m_R^2 = m_0^2 + 0.15 m_{1/2}^2$, $m_L^2 = m_0^2 + 0.52 m_{1/2}^2$ but the sneutrino mass matrix has changed (more freedom!)
- soft L-viol $\rightarrow$ control splitting $m_+ - m_-$: inelastic scattering with nucleons suppression
- role of the singlet $\rightarrow \sigma_{\tilde{\nu}}$ suppression, freedom

What we expect ?

Large values of $m_{1/2}$ (neutralino heavier than sneutrino) $\rightarrow$ heavy spartners

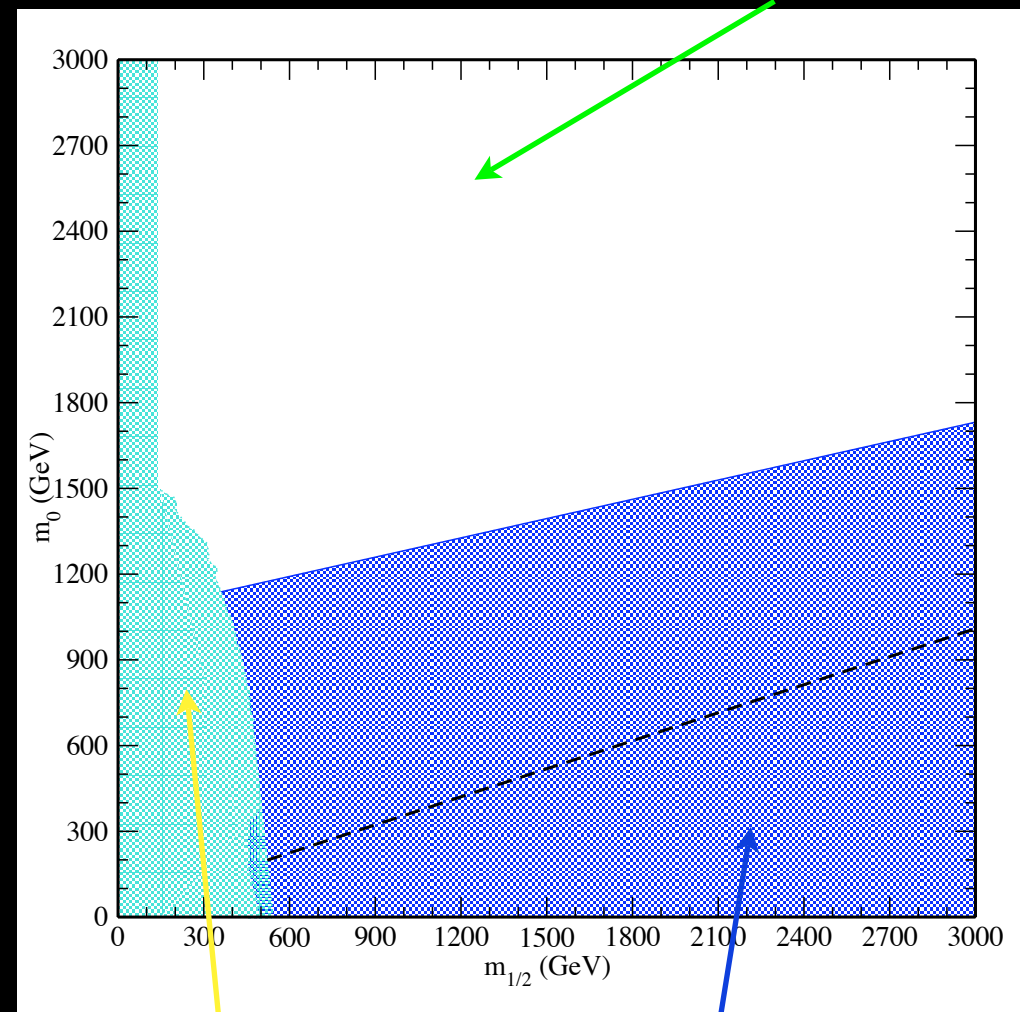
MSSM



InMSSM

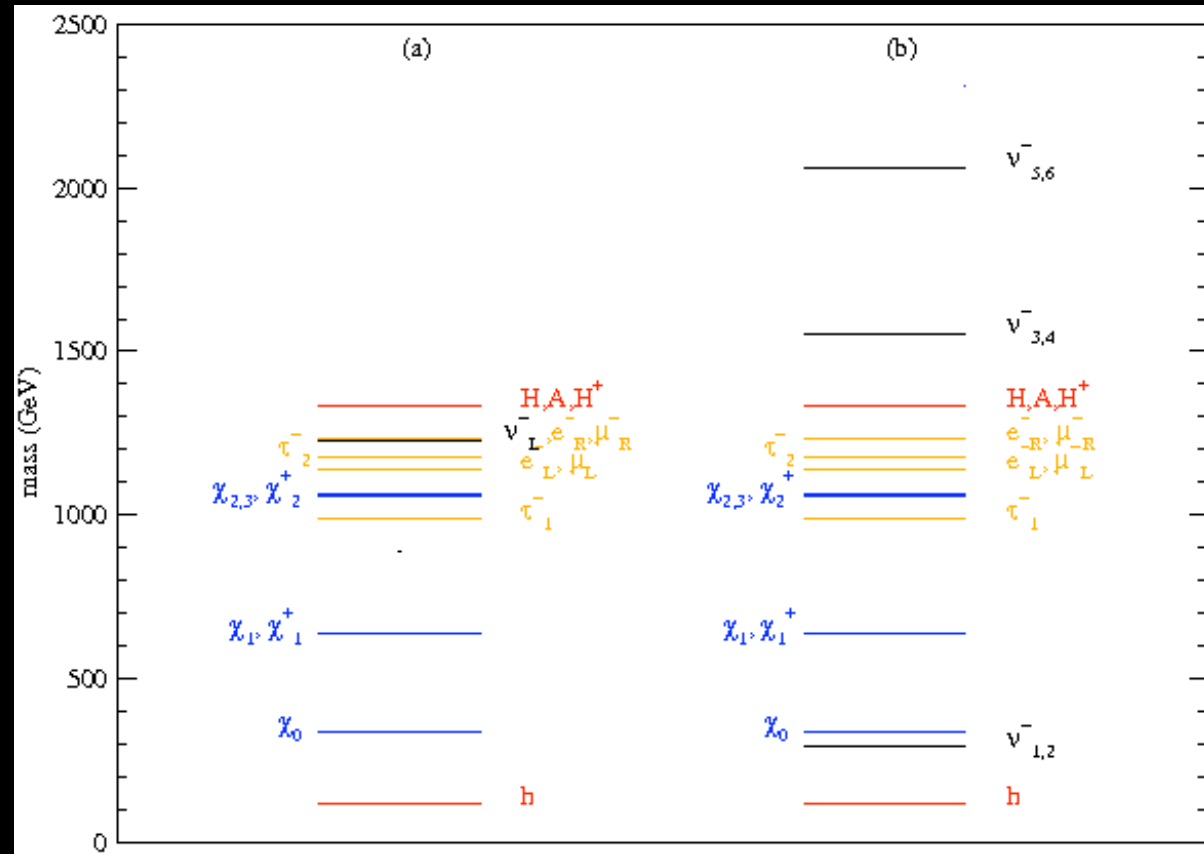
$$\tan \beta = 35, A_0 = 0, \mu > 0, B_{\mu_S} = 10 \text{ GeV}^2$$

allowed & neutralino LSP

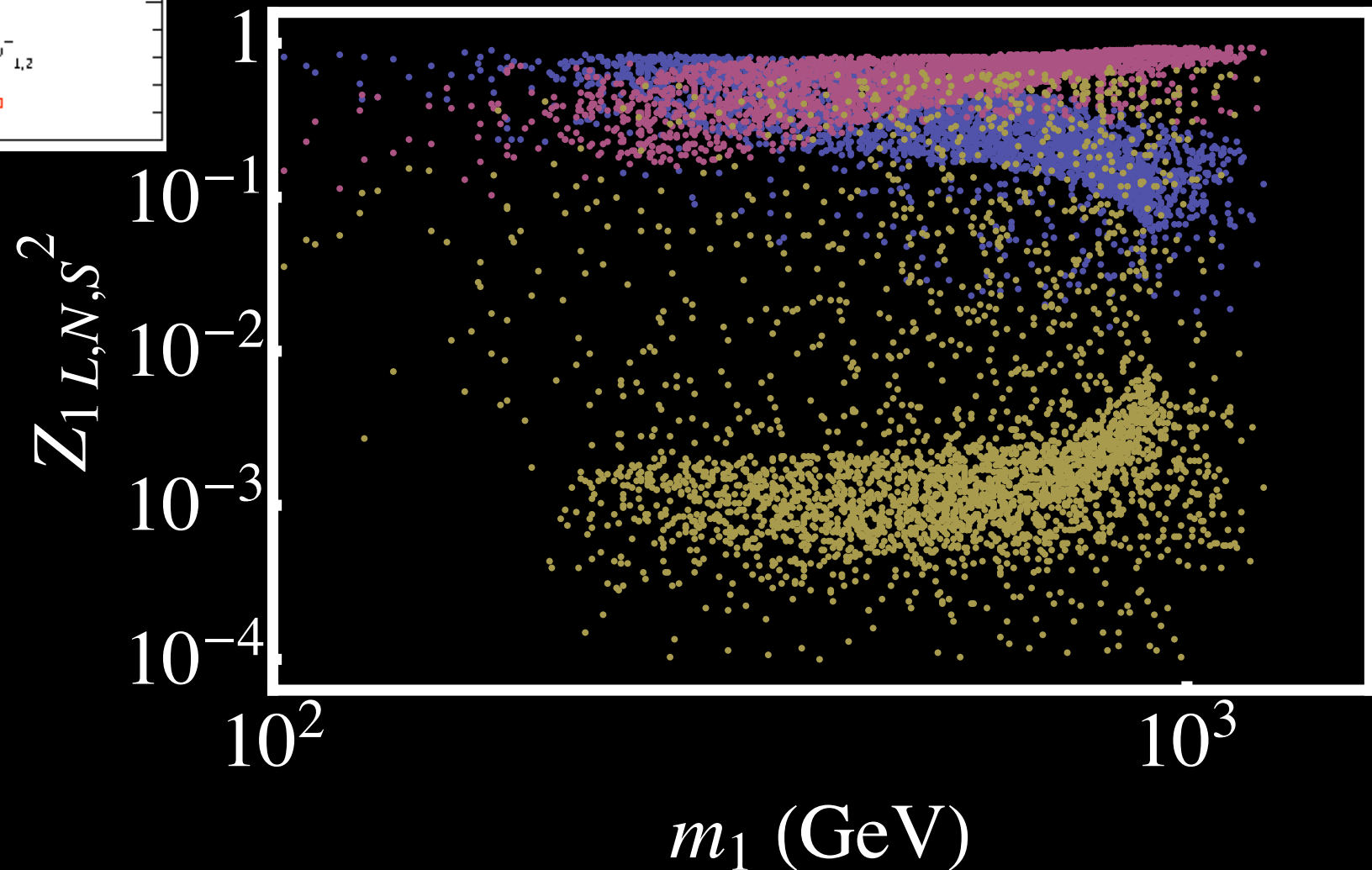


the spectrum

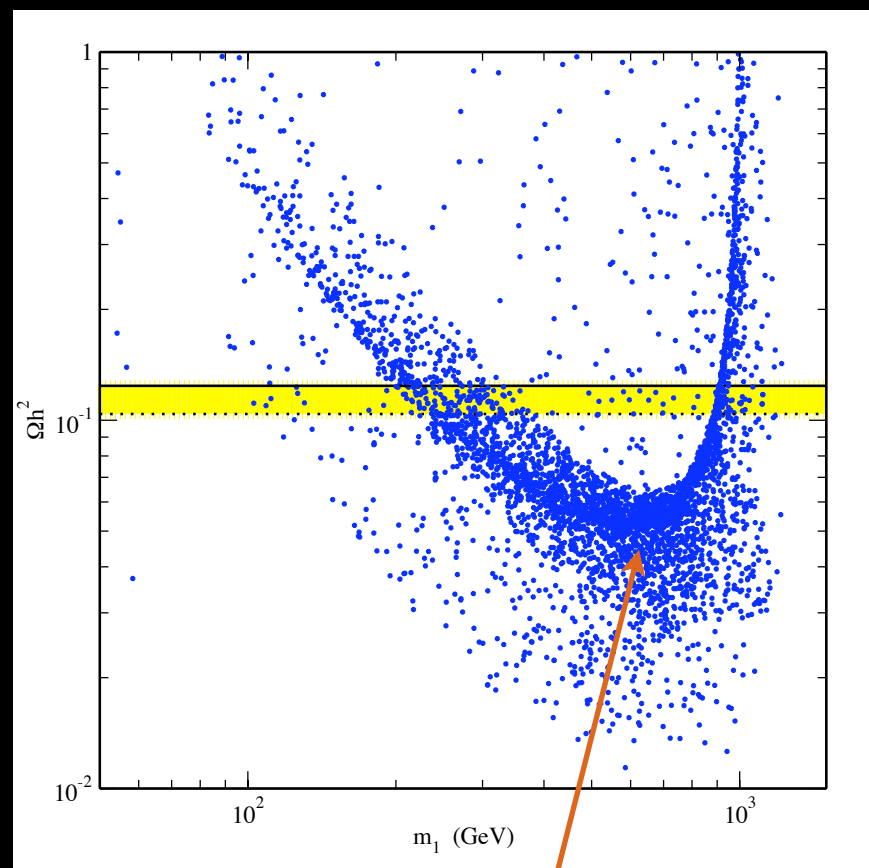
$$\tan \beta = 35, A_0 = 0, m_0 = 358 \text{ GeV}, m_{1/2} = 693 \text{ GeV}, \mu > 0, B_{\mu_S} = 10 \text{ GeV}^2$$



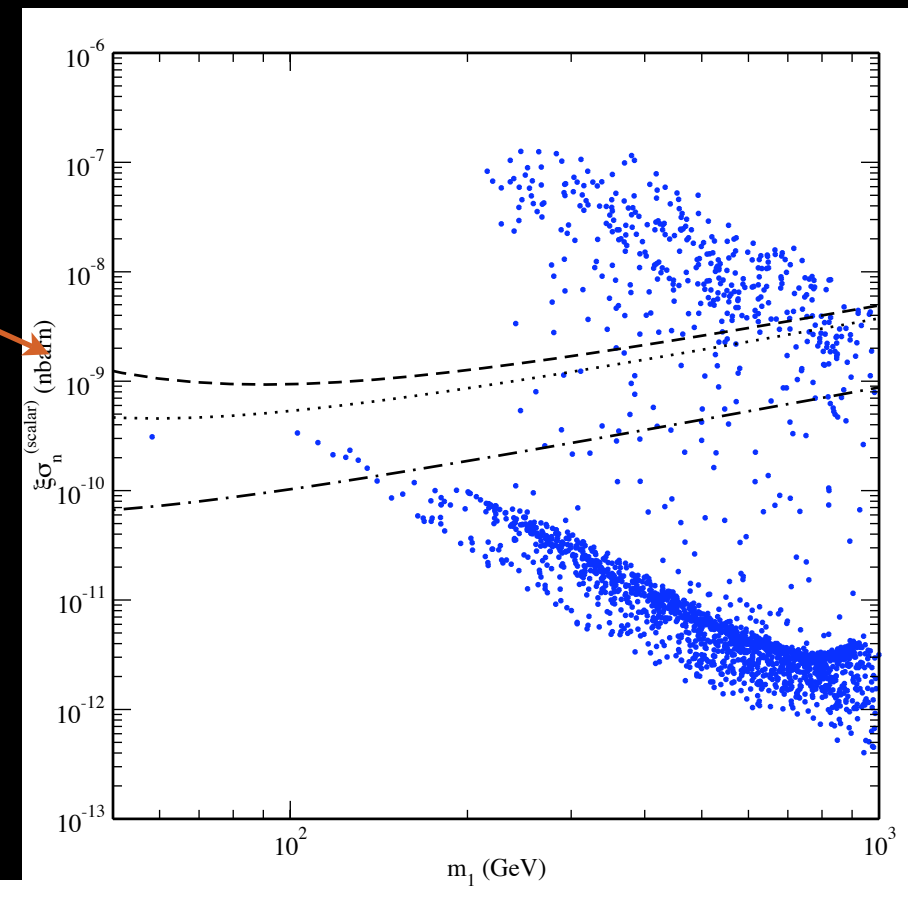
$$\begin{aligned} m_{\tilde{\nu}_{1,2}} &= 290 \text{ GeV} \\ m_{\chi_0} &= 340 \text{ GeV} \\ m_{\chi_1, \chi_1^+} &= 650 \text{ GeV} \\ m_{\tau_1} &= 950 \text{ GeV} \\ m_{\chi_{2,3}, \chi_2^+} &= 950 \text{ GeV} \\ m_{\tilde{e}_L, \tilde{\mu}_L} &= 1150 \text{ GeV} \end{aligned}$$



$0.1 < m_{0,1/2}/\text{TeV} < 3, , A_0 = 0, 3 < \tan \beta < 50$
 $10^{-9} < \mu_S/\text{GeV} < 10^{-6}, \mu > 0, 1 < B_{\mu_S}/\text{GeV}^2 < 100 \text{ GeV}^2$
 relic abundance direct searches

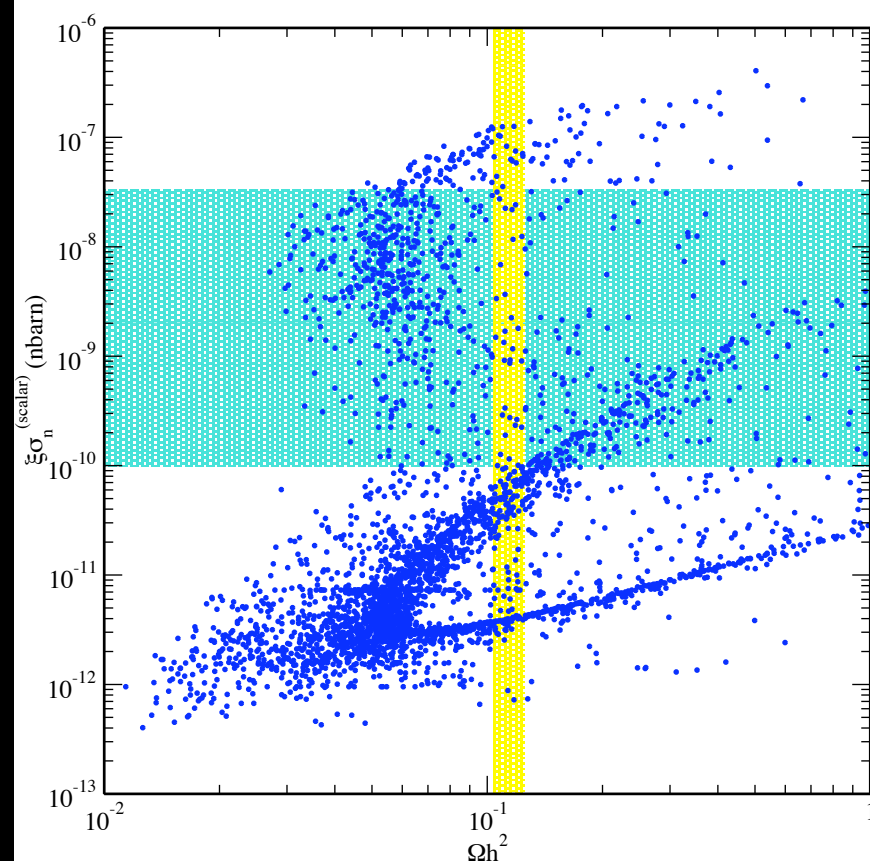


upper limits from CDMS
(3 galactic halo)



$$m_{\tilde{\nu}} - \Omega h^2$$

annihilation through the
H channel



$$m_{\tilde{\nu}} - \xi \sigma$$

$$\Omega h^2 - \xi \sigma$$

crucial ingredient of the InMSSM

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$$\mu_{SS}$$

small !
(keV-MeV)

$$m_L^\nu \sim m_D \frac{1}{M_{RS}} \mu_S \frac{1}{M_{RS}^T} m_D^T$$

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try to find a dynamical way to justify it!

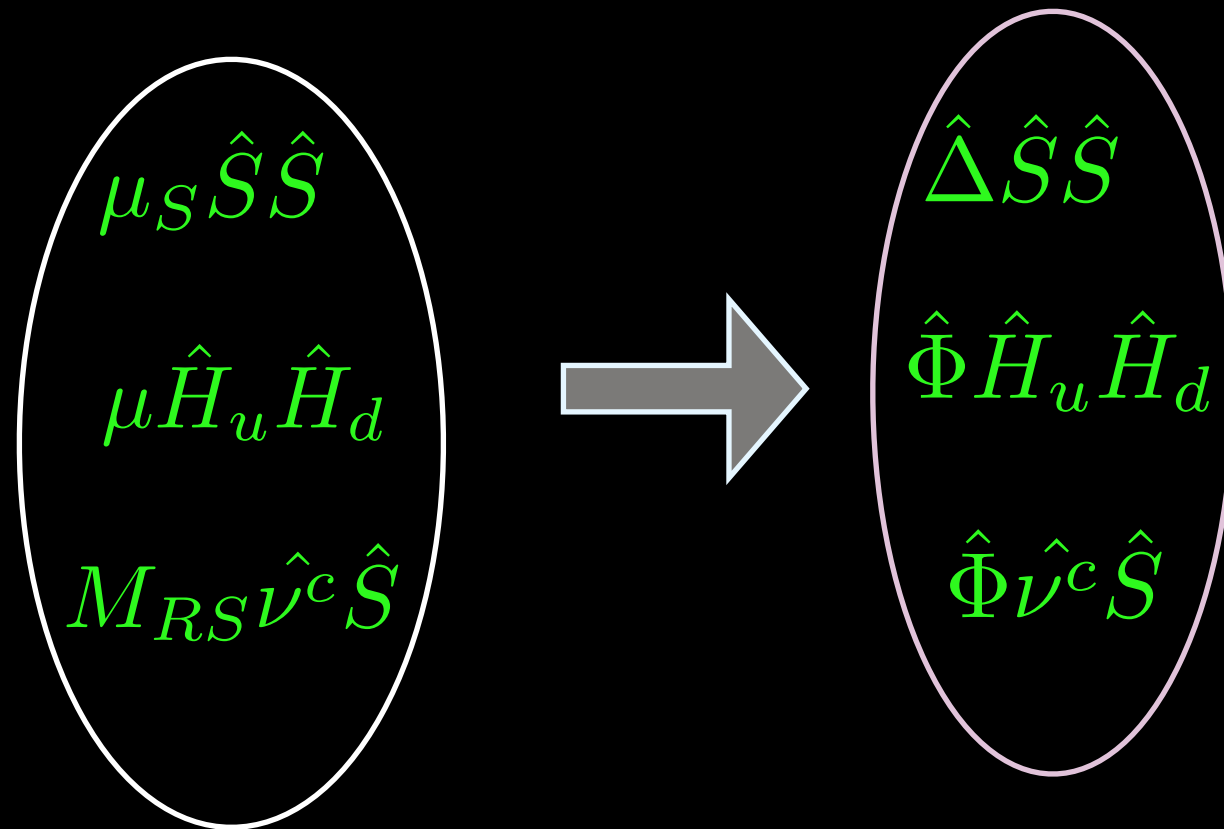
dynamical InMSSM μ_S, μ, M_{RS}

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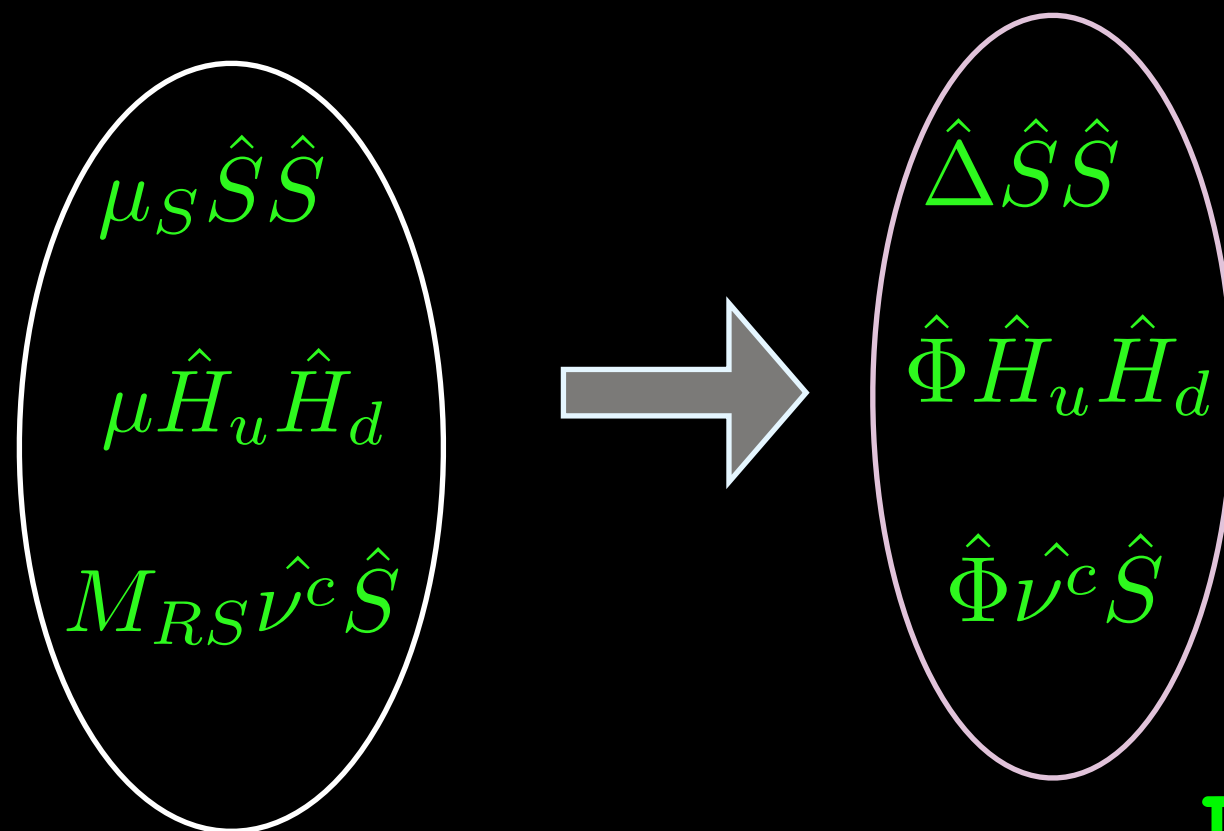


Cerdeno et al. PRD79(2009)

$$\mu_S \sim \langle \Delta \rangle$$

$$\mu, M_{RS} \sim \langle \Phi \rangle$$

dynamical InMSSM μ_S, μ, M_{RS}



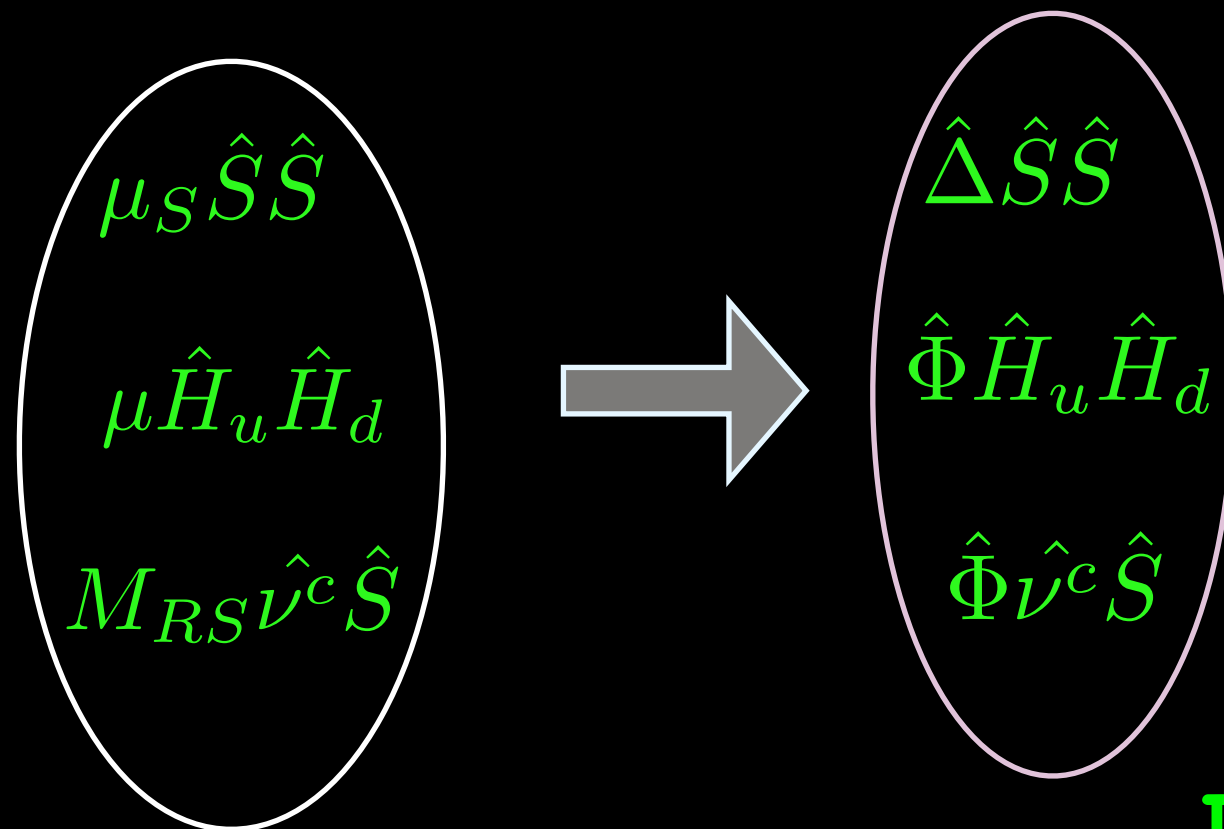
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InNMSSM : **NMSSM** + $\hat{\nu}^c, \hat{S}, \hat{\Delta}$
L (-1,1,-2)

dynamical InMSSM μ_S, μ, M_{RS}



Cerdeno et al. PRD79(2009)

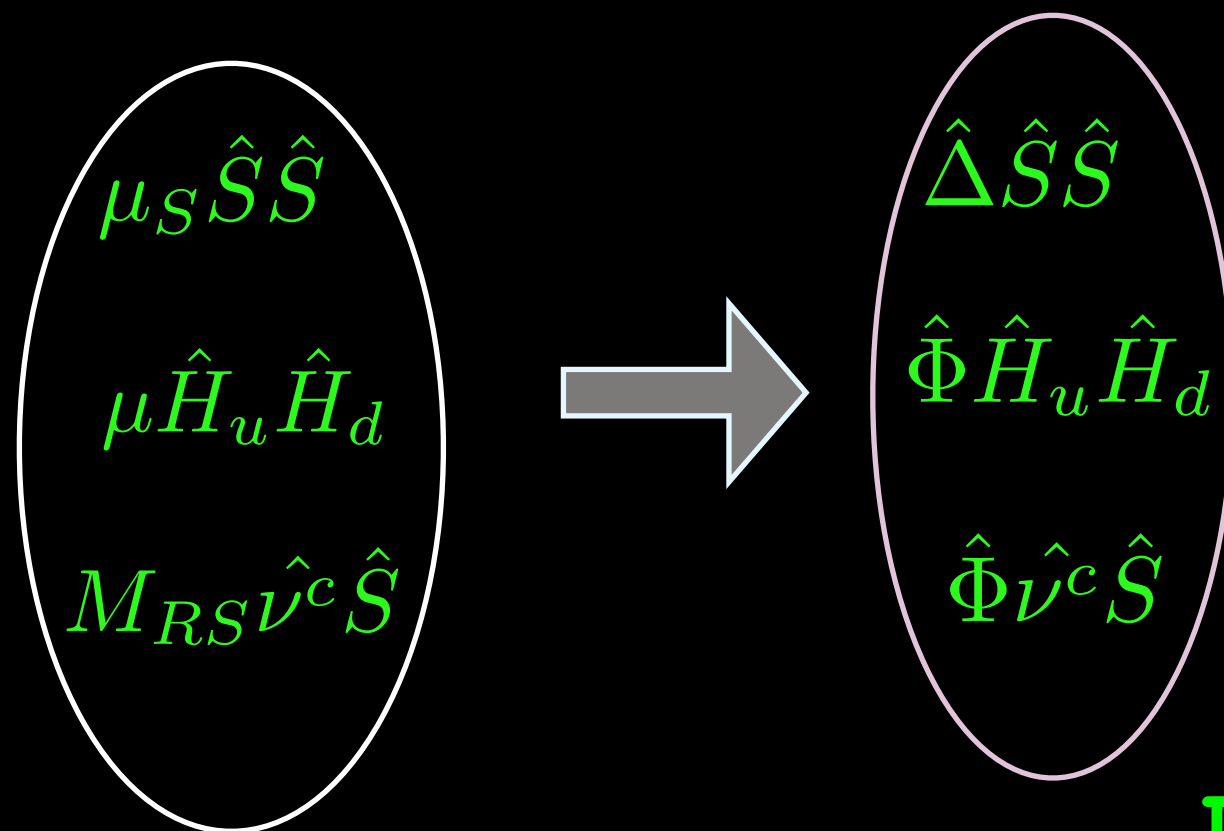
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$$\langle \Delta \rangle = 0$$

dynamical InMSSM μ_S, μ, M_{RS}



Cerdeno et al. PRD79(2009)

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$$\langle \Delta \rangle = 0$$

$$m_\nu \propto \mu_S \sim \langle \Delta \rangle$$

neutrinos massless!

InNMSSM : NMSSM + $\hat{\nu}^c, \hat{S}, \hat{\Delta}, \hat{\tilde{\Delta}}$
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$$\mathcal{W} \longrightarrow \frac{1}{2} \xi \hat{\Delta} \hat{\tilde{\Delta}} \hat{\tilde{\Delta}}$$

$$V_{soft} \longrightarrow \frac{1}{2} a_{\Delta} \Delta \tilde{\Delta} \tilde{\Delta}$$

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$$\langle \tilde{\Delta} \rangle \sim \langle \Phi \rangle$$

$$\langle \Delta \rangle \propto a_{\Delta}$$

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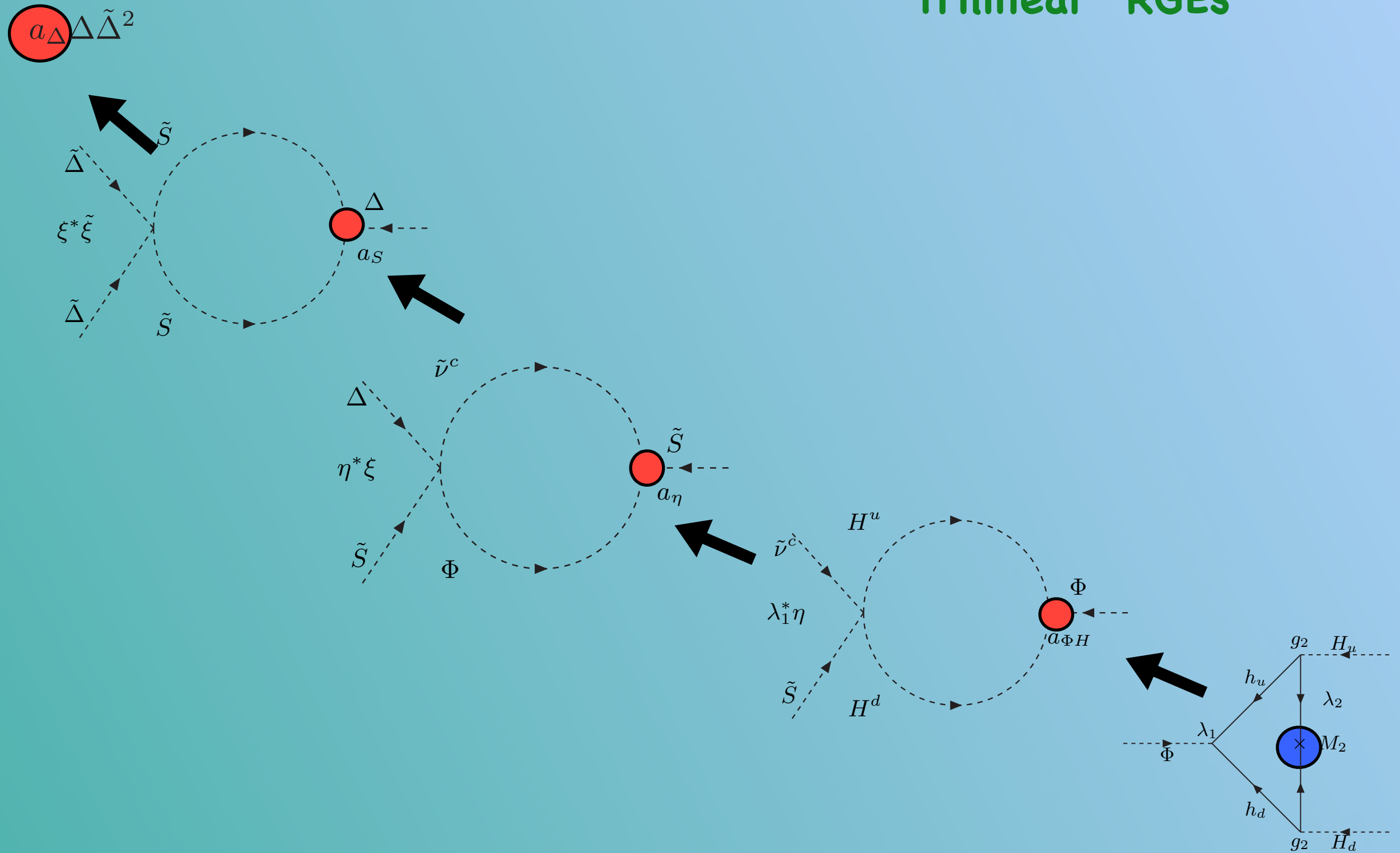
$$\langle \Delta \rangle \propto a_{\Delta}$$

justify its smallness

$$m_{\nu} \propto \mu_S \sim \langle \Delta \rangle$$

**assumption : all the SUSY trilinear soft terms
 vanish at GUT**

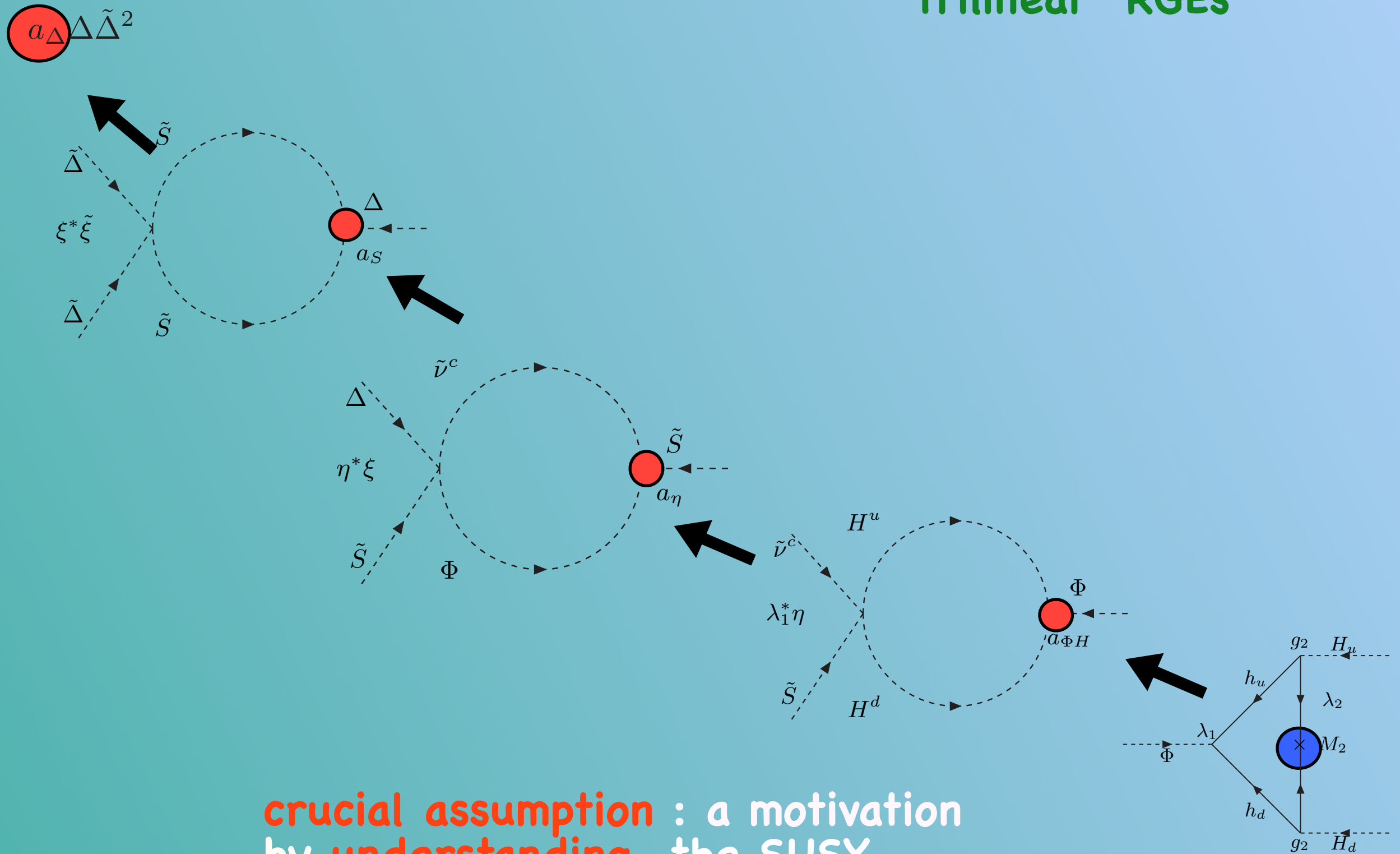
trilinear RGEs



$$a_0 = 0, M_{1/2} > 0$$

$$m_0^2 < 0, \tilde{m}_0^2 > 0$$

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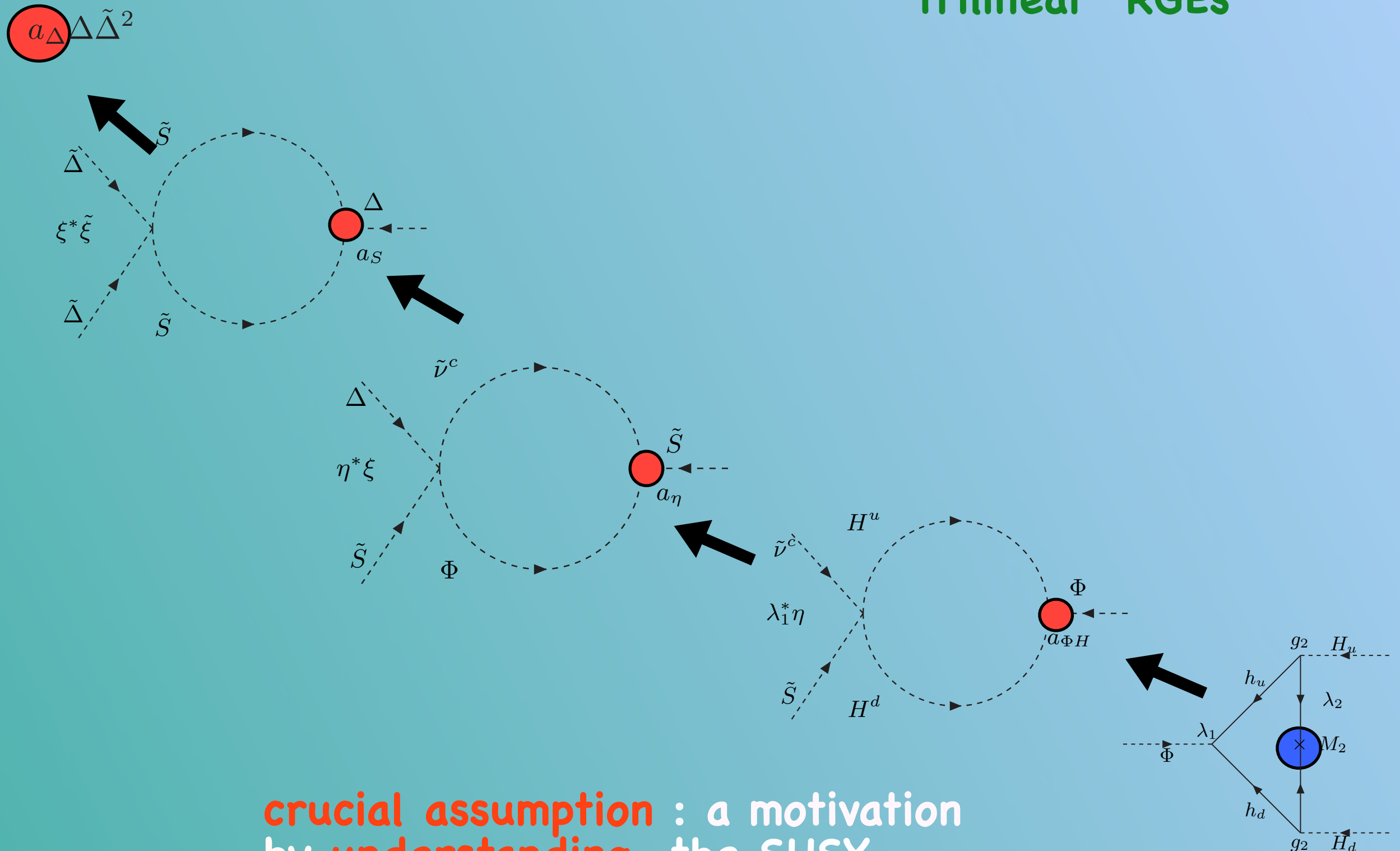


crucial assumption : a motivation
by **understanding** the SUSY
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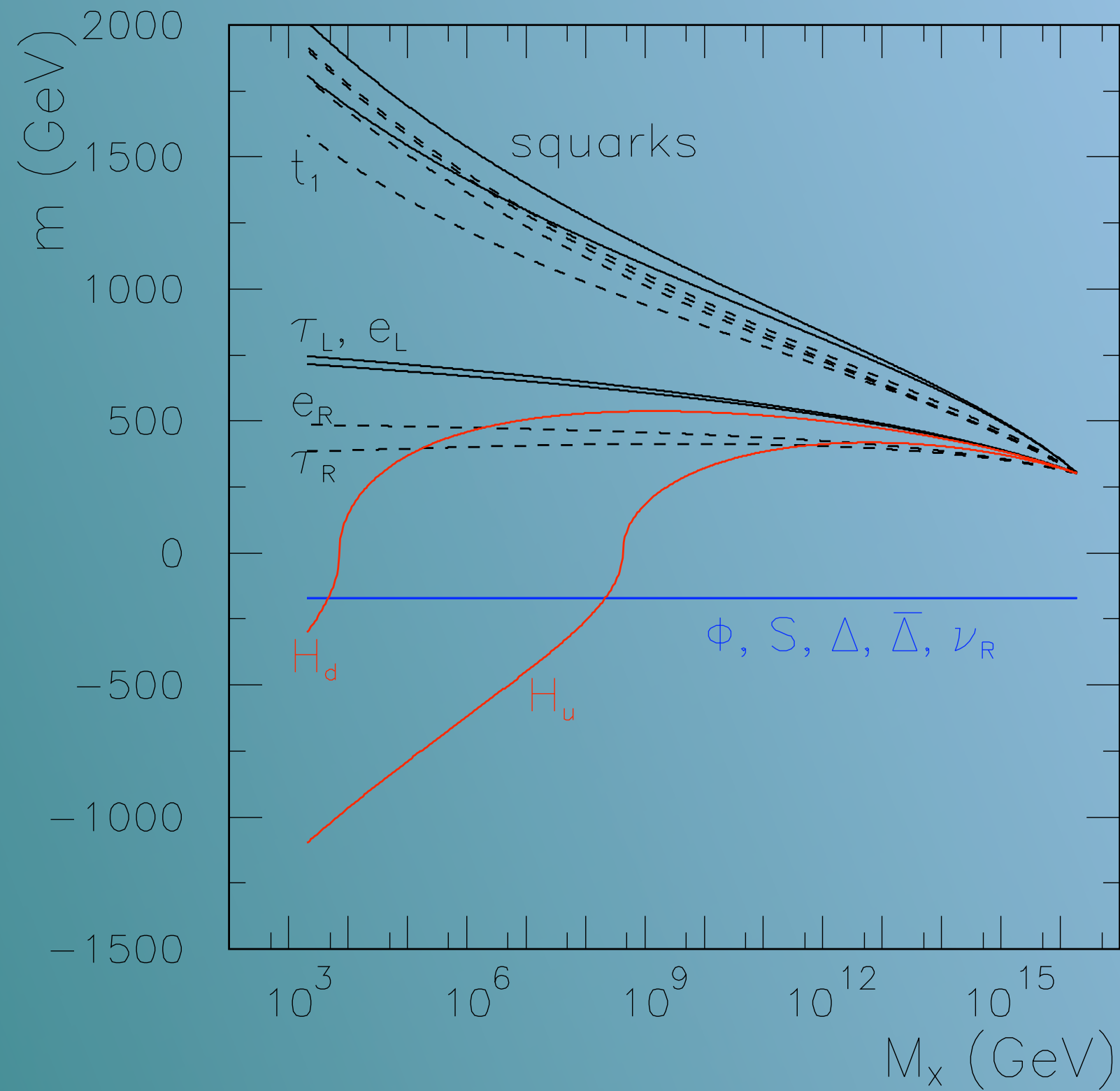
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(next step?)

$$a_0 = 0, M_{1/2} > 0$$

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mass RGEs



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CP-eigenstates
basis

$$a_\Delta \sim 0$$

$$M_{\tilde{\nu}_\pm}^2 \sim \begin{pmatrix} m_L^2 & 0 & 0 \\ 0 & m_0^2 + \alpha_\pm v^2 & \pm \delta v^2 \\ 0 & \pm \delta v^2 & m_0^2 + \beta_\pm v^2 \end{pmatrix}$$

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$$m_L^2 > 0, m_0^2 < 0$$

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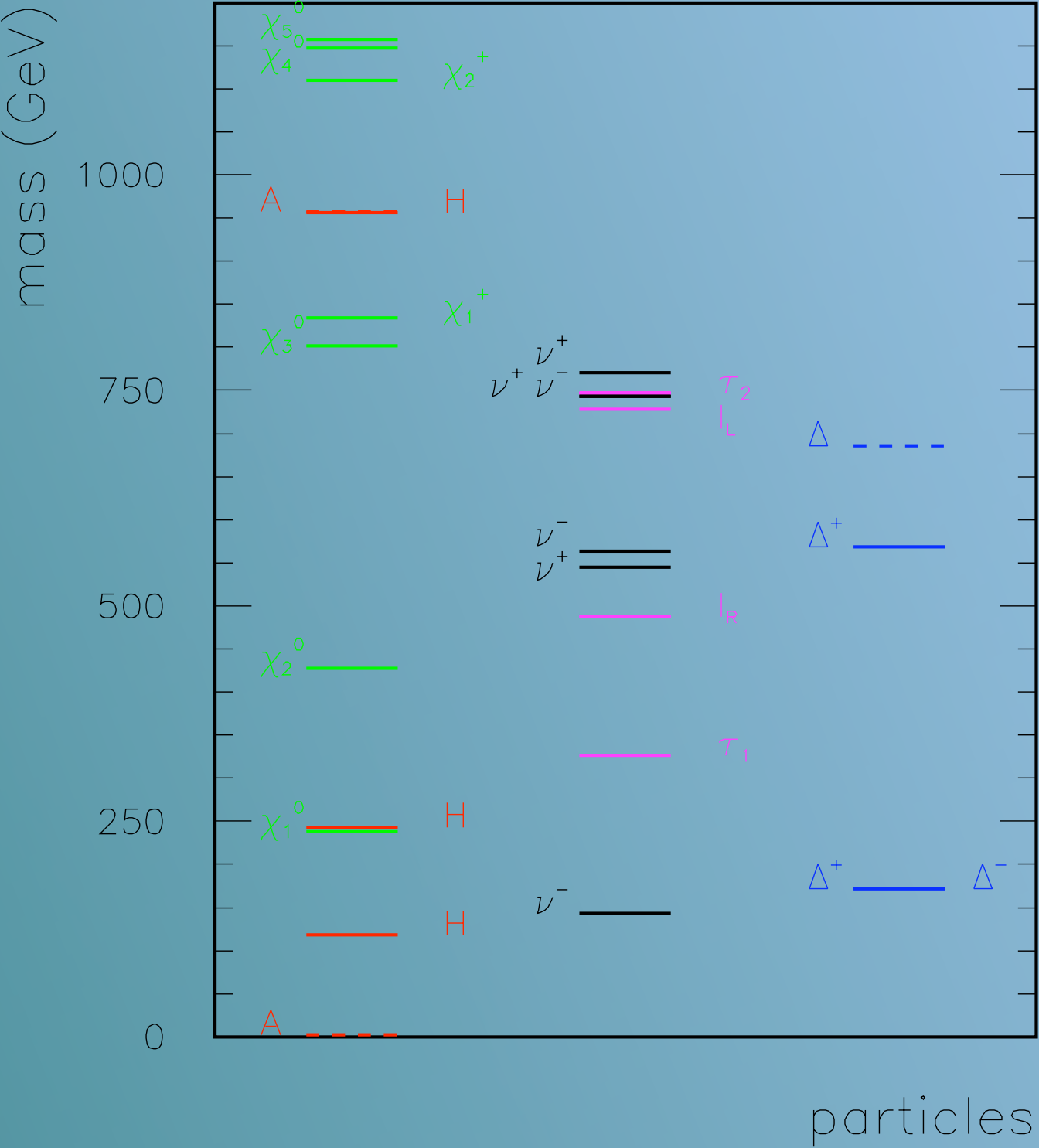
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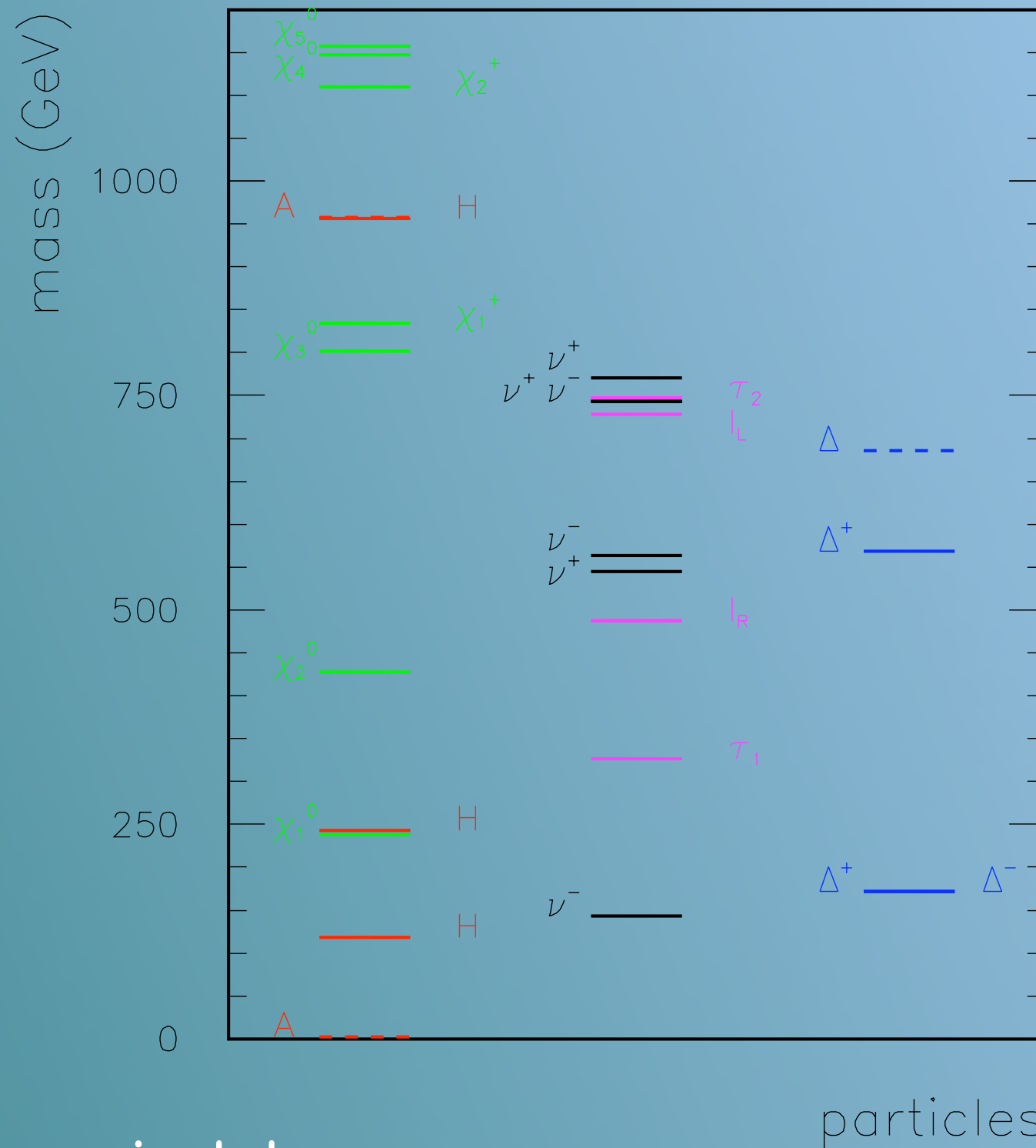
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pure gauge singlet, mixing right handed
and S scalar neutrino

spectrum

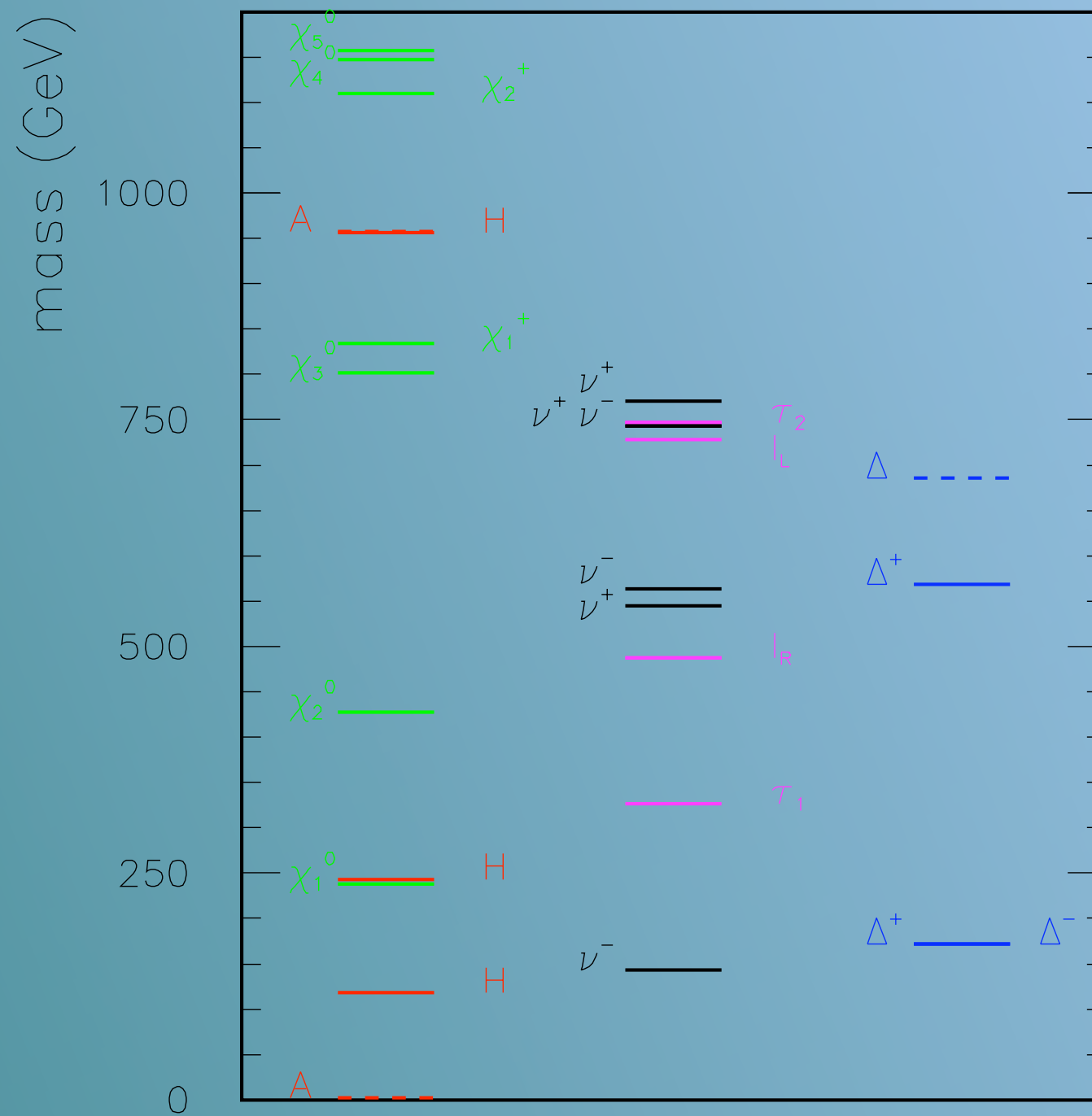


spectrum



if mainly gauge singlet,
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spectrum



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Cerdeno et al. PRD79(09)

new Higgs channel through
 $\hat{\Phi} \hat{H}_u \hat{H}_d$ $\hat{\Phi} \hat{\nu}^c \hat{S}$

Conclusions

- ★ different kind of extensions of MSSM leads the sneutrino to be a natural LSP as the neutralino
- ★ inverse see-saw mechanism is a very appealing mechanism to give mass to neutrinos (low scale)
- ★ in the InMSSM sneutrino LSP is also a valid CDM (indeed many configurations have the correct relic abundance & satisfy the direct search bounds)
- ★ 'standard prejudice' to justify the smallness of the μ -parameter
- ★ built a model (InNMSSM) in which both the ' μ ' parameters arise dynamically
- ★ neutrino masses linked to soft SUSY breaking terms
- ★ region where the singlet sneutrino is the LSP
- ★ new possible DM scenario
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