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# **The path to new physics at the LHC and beyond**

## **XV Christmas Workshop**

**Windows on the Unknown**

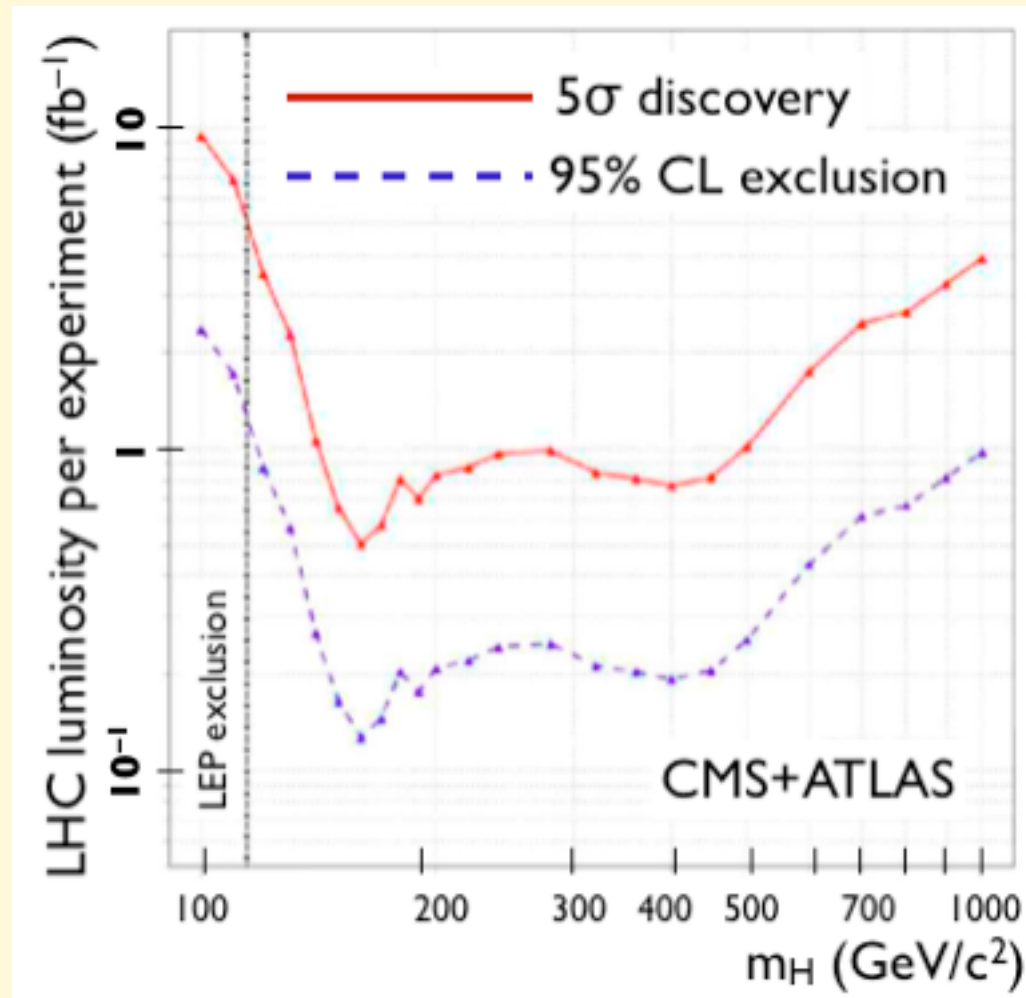
**Madrid, 16-18 Dec 2009**

**The LHC was designed  
to answer one question:**

**is electroweak symmetry broken as  
postulated in the SM Higgs mechanism?**

- SM production and decay rates well known
- Detector performance for SM channels well understood
- $115 < m_H < 200$  from LEP and EW fits in the SM

## Summary of SM Higgs discovery potential



**Within ~3 yrs from startup we should have an answer**

## IF Higgs seen with SM production/decay rates, but outside SM mass range:

- new physics to explain EW fits, or
- problems with LEP/SLD data

In either case,

- easy prey with low luminosity up to  $\sim 800$  GeV, but more lum is needed to understand why it does not fit in the SM mass range!

## IF NOT SEEN UP TO $m_H \sim 0.8$ -1 TeV GEV:

$\sigma < \sigma_{SM}: \Rightarrow$  **new physics**

or

$BR(H \rightarrow \text{visible}) < BR_{SM}: \Rightarrow$  **new physics**

or

$m_H > 800$  GeV: expect  $WW/ZZ$  resonances at  $\sqrt{s} \sim \text{TeV} \Rightarrow$  **new physics**

**•Sorting out non-SM scenarios may take longer than the SM H observation, and may well require LHC luminosity upgrades and/or a lepton collider, but the conclusion about the existence of a BSM origin of EWSB will come early and unequivocal**

**•Exposing the mechanism of EW symmetry breaking (EWSB) and identifying the Higgs boson or its alternatives is necessary to set the scene for what's next**

**We would also like the LHC to help us address the three key experimental shortcomings of the Standard Model:**

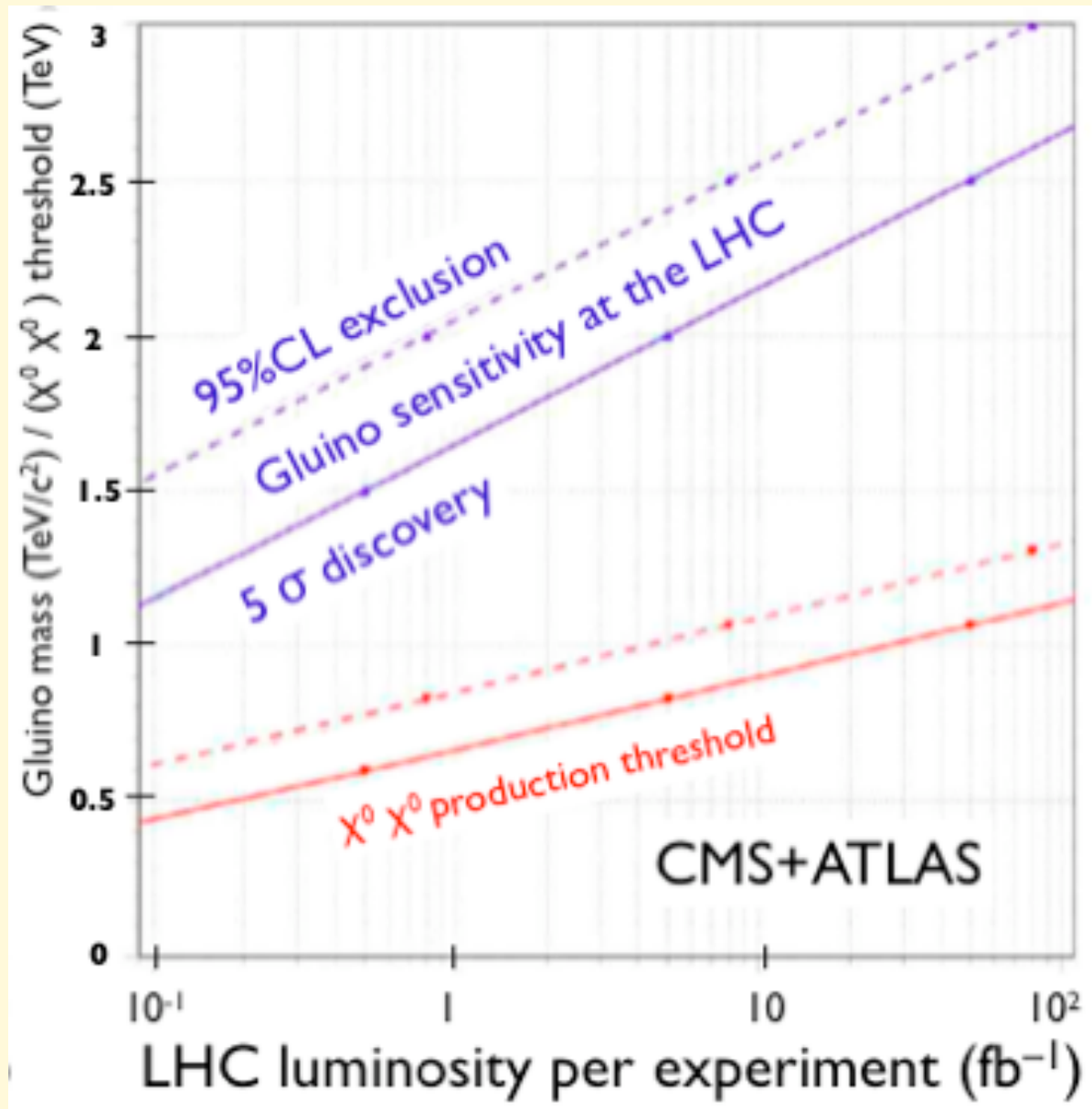
- **Neutrino masses**
- **Dark matter**
- **Baryon asymmetry of the universe**

**as well as its theoretical weakness, the **hierarchy problem****

Will the answers to these questions be related to each other?

Which experimental programme, at the LHC and beyond, will allow us to address them?

## Example: timeline for SUSY (MSSM) discovery/exclusion



**The discovery of SUSY (or whatever else) will only be the beginning of a new era of exploration, dominated by questions like**

**what is the mechanism of SUSY breaking?**

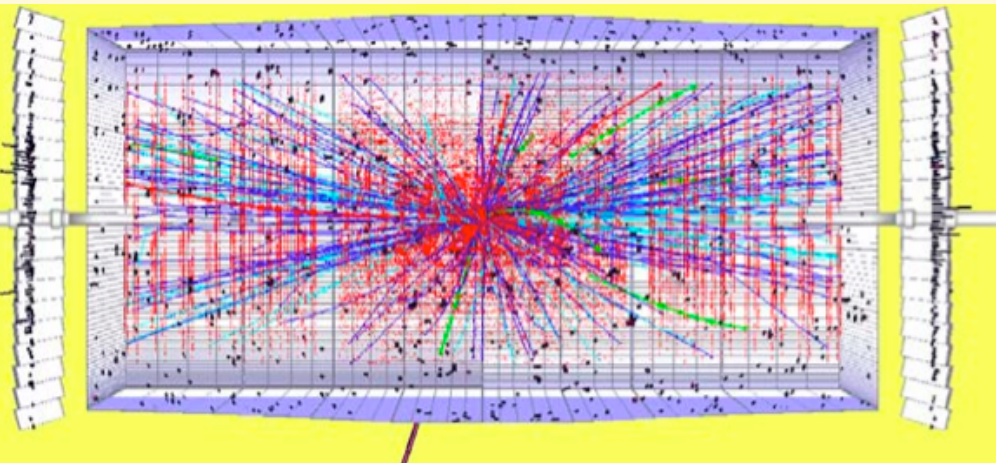
**New expt'l input will be needed to start addressing such issues. E.g.:**

- Chargino/gluino mass spectrum
- Squarks and sleptons masses and mixings
- CP structure of SUSY couplings
- .....

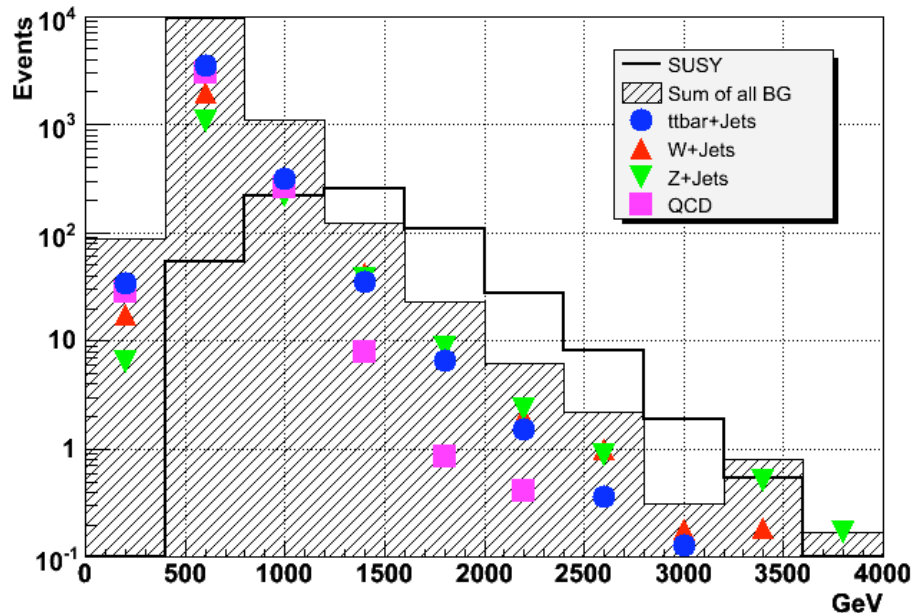


# The LHC inverse problem

Reconstruct the Lagrangian of new physics from the LHC data

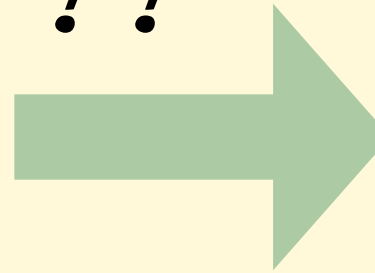


Effective Mass 0lepton SUSY



$$M_{\text{eff}}(\text{GeV}) = \sum_{i=1,4} E_T(i) + E_T^{\text{miss}}$$

??



??  
!

extra dims?

gluinos?

Little Higgs?

$\mathcal{L}$

- Very likely, the understanding of the new physics will emerge from a step-by-step consolidation of prominent features of the data, restricting more and more the class of models first, and their parameters later.
- Single key inputs, even if only partially accurate, can provide more valuable information than dozens of vaguely suggestive hints. For example, if SUSY:
  - the relation between gluino and chargino mass,
  - evidence for GMSB in the final states (prompt photons and MET),
  - the determination of the stop parameters and  $m_H$ , etc.

- **< 1973: theoretical foundations of the SM**

- renormalizability of  $SU(2) \times U(1)$  with Higgs mechanism for EWSB
- asymptotic freedom, QCD as gauge theory of strong interactions
- KM description of CP violation

- **Followed by 30 years of consolidation:**

- **technical theoretical advances** (higher-order calculations, lattice QCD, ...)
- **experimental** verification, via **discovery** of
  - **Fermions:** charm, 3rd family (USA)
  - **Bosons:** gluon, W and Z (Europe; .... waiting to add the Higgs ....)
- **experimental** consolidation, via **measurement** of
  - EW radiative corrections
  - running of  $\alpha_s$
  - CKM parameters

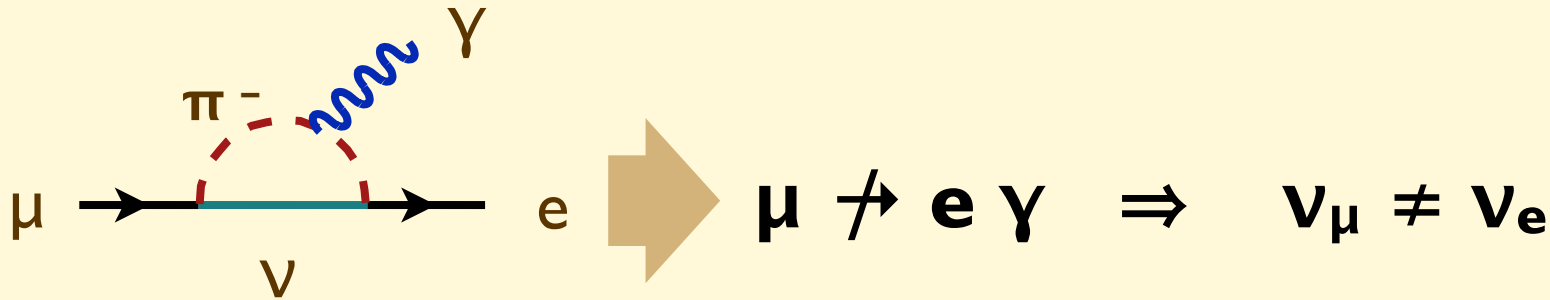
**It's unlikely it will take less than 30 yrs to clarify and consolidate the understanding of new phenomena to be unveiled by the LHC!**

Notice that of the 3 empirical proofs that the SM is incomplete:

- **Neutrino masses**
- **Dark matter**
- **Baryon asymmetry of the universe**

at least **two** are directly related to flavour .....

Flavour phenomena have contributed shaping modern HEP as much as the gauge principle



$K^0 \rightarrow \mu^+ \mu^-$   
 **$K$ - $\bar{K}$  mixing**

➡ **GIM, charm**

$\epsilon_K$ ,  **$\mathcal{CP}$**

➡ **KM**

**$B_d$ - $\bar{B}_d$  mixing**

➡ **large  $m_{\text{top}}$ , well before EW tests**

**$\nu$  masses**

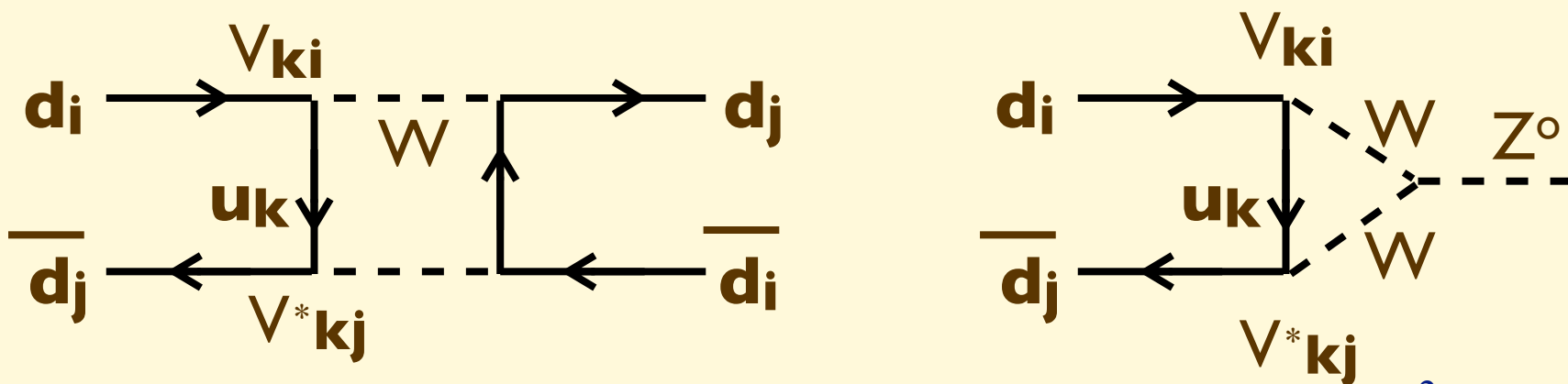
➡ **see-saw,  $SO(10)$  GUT, ... ?**

# What is “flavour physics” ?

- In the SM, flavour is what deals with the fermion sector (family replicas, spectra and mixings):
  - all flavour phenomena are encoded in the fermion Yukawa matrices.

# FCNC and CPV in the SM

- Suppression of FCNC and CPV are guaranteed in the SM by the following facts:
  - Quark sector:
    - unitarity of CKM (GIM mechanism)
    - small mixings between heavy and light generations



$$\Delta_{ij} \sim \sum_{k=u,c,t} V_{ki} V_{kj}^* f(m_k/m_W) \sim \sum_{k=c,t} V_{ki} V_{kj}^* m_k^2/m_W^2 \sim V_{ci} V_{cj}^* \frac{m_c^2}{m_W^2} + V_{ti} V_{tj}^*$$

- Lepton sector:
  - $m_\nu=0 \Rightarrow$  all phases and angles absorbed by field redefinitions, no mixings/CPV at all

# What is “flavour physics” ?

- In the SM, flavour is what deals with the fermion sector (family replicas, spectra and mixings):
  - all flavour phenomena are encoded in the fermion Yukawa matrices.
- Beyond the SM, “flavour” phenomena cover a wider landscape. E.g.
  - FCNC can be mediated by
    - gauge-sector particles, like charged higgses, gauginos, new gauge bosons, or by
    - SUSY scalar partners
  - New flavours in the form of new generations, exotic partners of standard quarks (e.g. Kaluza Klein excitations,  $T'$  in LH), etc.
  - CP violation can reside in gauge/Higgs couplings



# FCNC beyond the SM

S.Geer

- There is absolutely no guarantee that the suppression of FCNC and CPV is present in extensions of the SM
- As soon as these are released, effects are devastating!

Compare the to  $O(10 \text{ TeV})$  sensitivity w.r.t. modifications of the gauge/EW sector

	$B(K_L \rightarrow \mu e) < 4.7 \times 10^{-12}$	$M_X > 150 \text{ TeV}/c^2$
	$B(K^+ \rightarrow \pi^+ \mu^+ e^-) < 4 \times 10^{-11}$	$M_X > 31 \text{ TeV}/c^2$
	$B(K_L \rightarrow \pi^0 \mu^+ e^-) < 3.2 \times 10^{-10}$	$M_X > 37 \text{ TeV}/c^2$
	$B(\mu^+ \rightarrow eee) < 1 \times 10^{-12}$	$M_X > 86 \text{ TeV}/c^2$
	$B(\mu^+ \rightarrow e^+ \gamma) < 1.2 \times 10^{-11}$	$M_X > 21 \text{ TeV}/c^2$
	Normalized Rate $< 6.1 \times 10^{-13}$	$M_X > 365 \text{ TeV}/c^2$

N.B. Once coupling constants – say of EW size – and  $O(\theta_c)$  mixings, are included, these scales are not much bigger than the TeV scale accessible at the LHC  $\Rightarrow$

**great potential synergy between  
LHC and flavour observables**

# EWSB and flavour

- EWSB is intimately related to flavour:
  - No EWSB  $\Rightarrow$  fermions degenerate  $\Rightarrow$  no visible flavour effect
- In most EWSB models flavour plays a key role. E.g.:
  - Technicolor: tightly constrained by large FCNC
  - Supersymmetry: large value of top mass drives radiative EWSB
  - In several extra-dim models the structure of extra dimensions -- driven by the need to explain the hierarchy problem of EWSB -- determines the fermionic mass spectrum
  - Little Higgs theories  $\Rightarrow$  top quark partners
- Why  $m_{\text{top}} = g/\sqrt{2} m_W$  ( $\Leftrightarrow y_{\text{top}} = 1$ ) ?

# What will be the main driving theme of the exploration of the new physics revealed by the LHC?

the gauge sector  
(Higgs, EWSB)



## The High Energy Frontier

LHC  
SLHC  
VLHC  
ILC  
CLIC  
....

the flavour sector  
( $\nu$  mixings, CPV, FCNC,  
EDM, LFV)



## The High Intensity Frontier

Neutrinos:	Quarks:	Charged leptons:
super beams	B factories	stopped $\mu$
beta-beams	K factories	$\ell \rightarrow \ell'$ conversion
$\nu$ factory	n EDM	e/ $\mu$ EDM

+ Astrophysics and cosmology

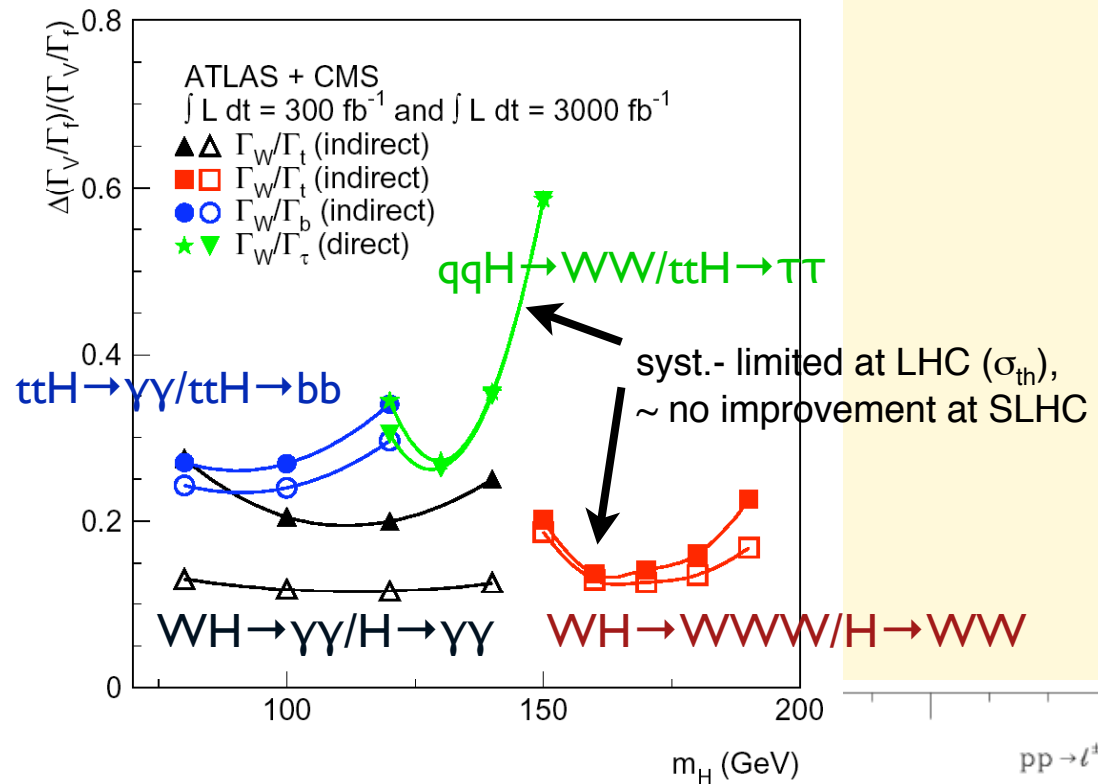
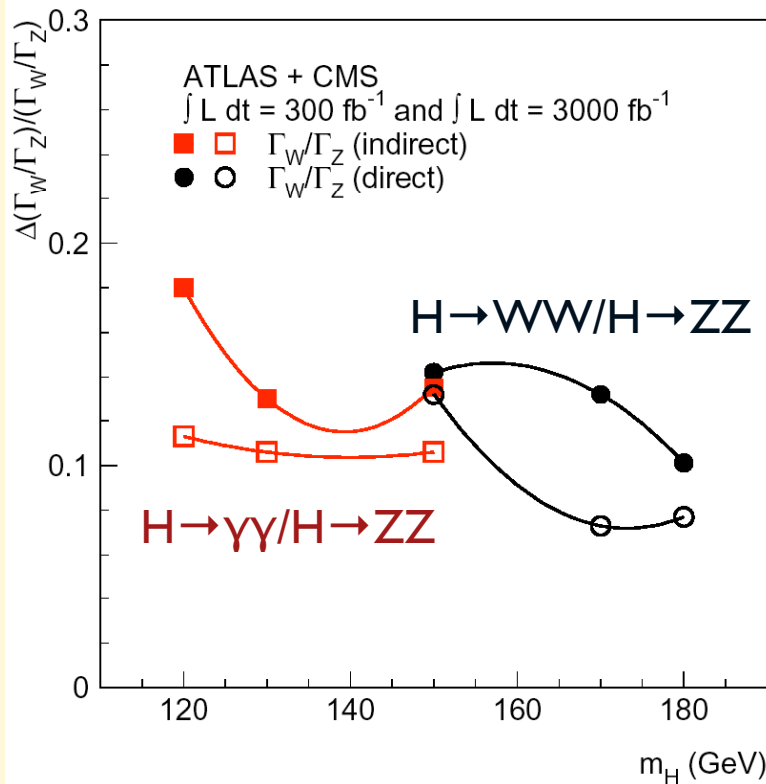
# What can we get from more integrated luminosity after LHC's first phase?

1. Improve measurements of new phenomena seen at the LHC. E.g.
  - Higgs couplings and self-couplings
  - Properties of SUSY particles (mass, decay BR's, etc)
  - Couplings of new  $Z'$  or  $W'$  gauge bosons (e.g. L-R symmetry restoration?)
2. Detect/search low-rate phenomena inaccessible at the LHC. E.g.:
  - $H \rightarrow \mu^+ \mu^-$ ,  $H \rightarrow Z \gamma$
  - top quark FCNCs
3. Push sensitivity to new high-mass scales. E.g.
  - New forces (  $Z'$ ,  $W_R$  )
  - Quark substructure
  - ....

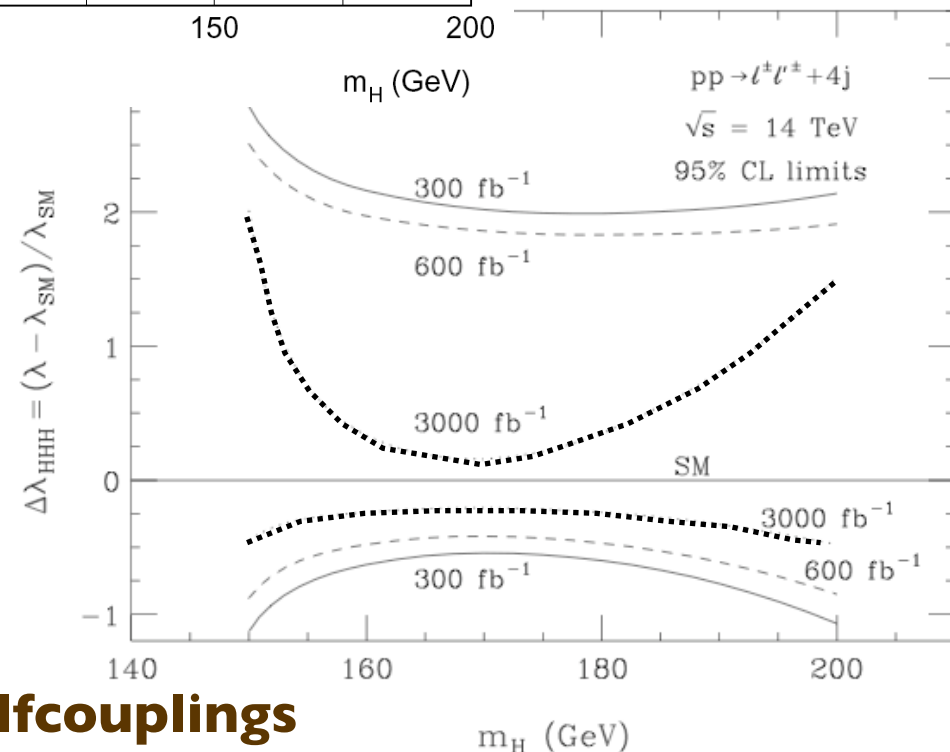
Energies/masses in the few-100 GeV range.  
Detector performance at SLHC should equal (or improve) in absolute terms the one at LHC

Very high masses, energies, rather insensitive to high-lum environment.  
Not very demanding on detector performance  
Slightly degraded detector performance tolerable

# Measurement of Higgs couplings - Accuracy goal: 10-20%

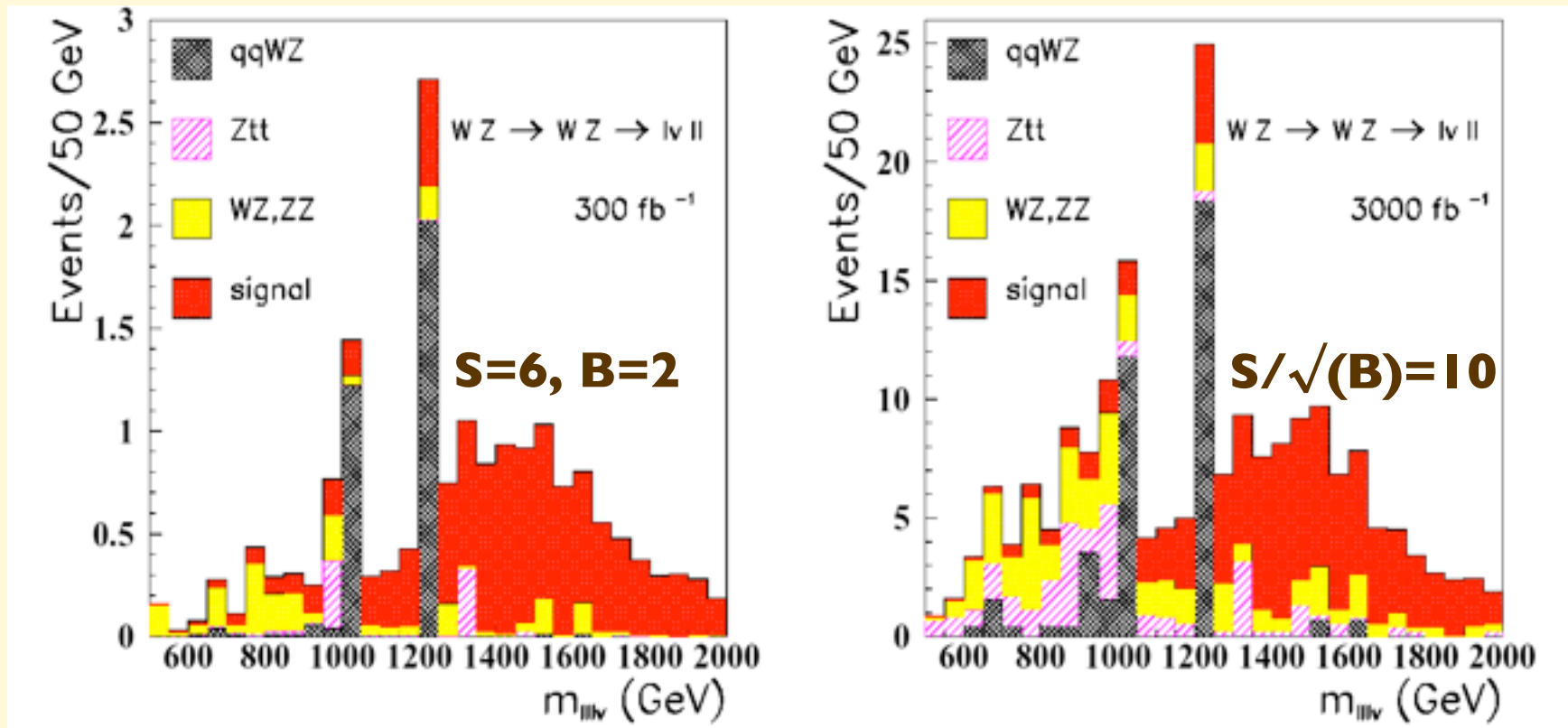
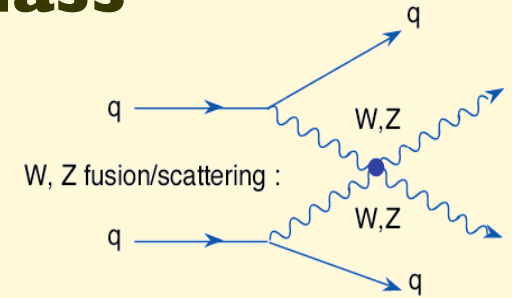


## Higgs boson couplings to fermions and gauge bosons



## Higgs boson selfcouplings

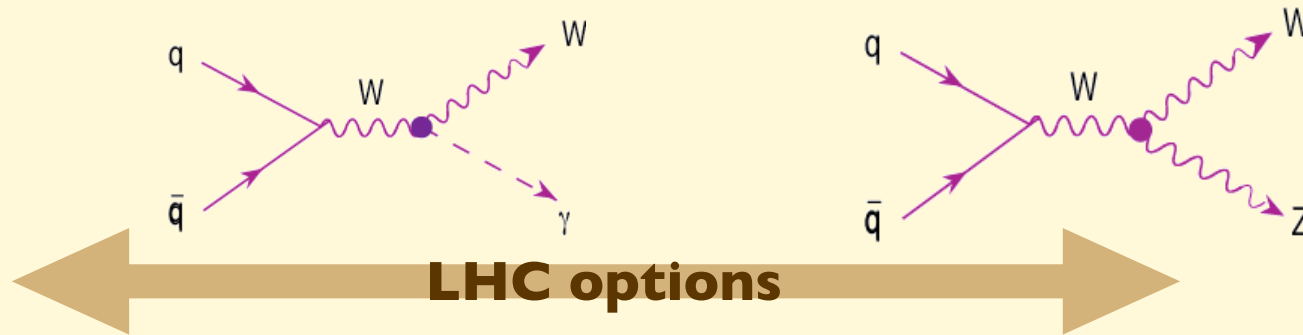
# Strong resonances in high-mass WW or WZ scattering



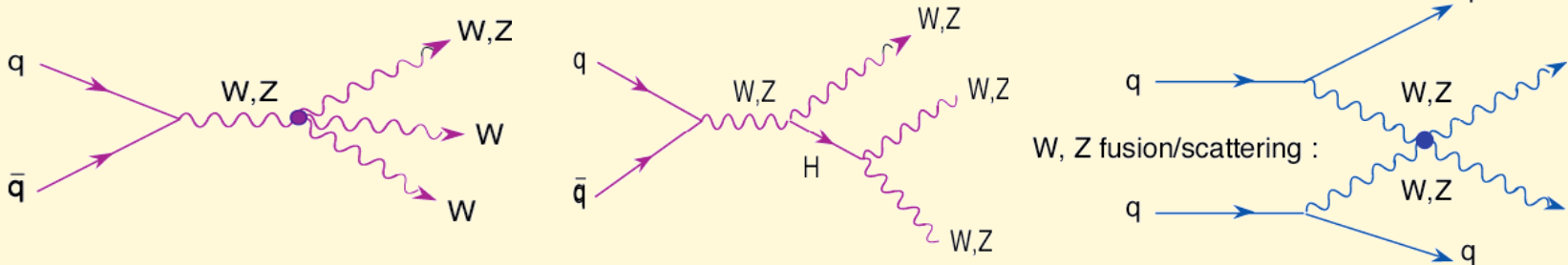
**Vector resonance** ( $\rho$ -like) in  $W_L Z_L$  scattering from Chiral Lagrangian model  
 $M = 1.5 \text{ TeV}$ , leptonic final states,  $300 \text{ fb}^{-1}$  (LHC) vs  $3000 \text{ fb}^{-1}$  (SLHC)

# Ex: Precise determinations of the self-couplings of EW gauge bosons

5 parameters describing weak and EM dipole and quadrupole moments of gauge bosons. The SM predicts their value with accuracies at the level of  $10^{-3}$ , which is therefore the goal of the required experimental precision



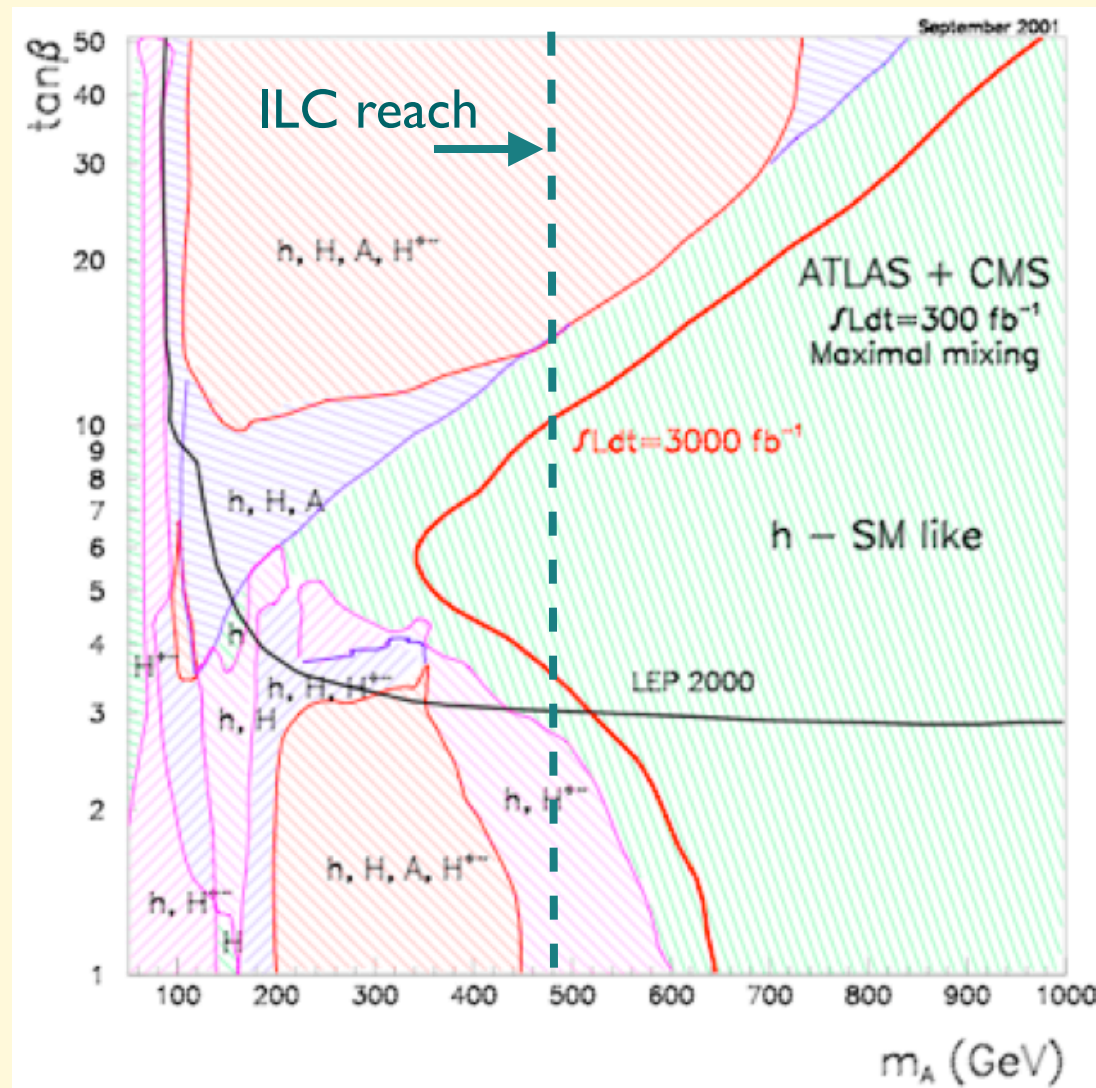
Coupling	14 TeV 100 fb <sup>-1</sup>	14 TeV 1000 fb <sup>-1</sup>	28 TeV 100 fb <sup>-1</sup>	28 TeV 1000 fb <sup>-1</sup>	LC 500 fb <sup>-1</sup> , 500 GeV
$\lambda_\gamma$	0.0014	0.0006	0.0008	0.0002	0.0014
$\lambda_Z$	0.0028	0.0018	0.0023	0.009	0.0013
$\Delta\kappa_\gamma$	0.034	0.020	0.027	0.013	0.0010
$\Delta\kappa_Z$	0.040	0.034	0.036	0.013	0.0016
$g_1^Z$	0.0038	0.0024	0.0023	0.0007	0.0050



(LO rates, CTEQ5M, $k \sim 1.5$ expected for these final states)						
Process	WWW	WWZ	ZZW	ZZZ	WWWW	WWWZ
$N(m_H = 120 \text{ GeV})$	2600	1100	36	7	5	0.8
$N(m_H = 200 \text{ GeV})$	7100	2000	130	33	20	1.6

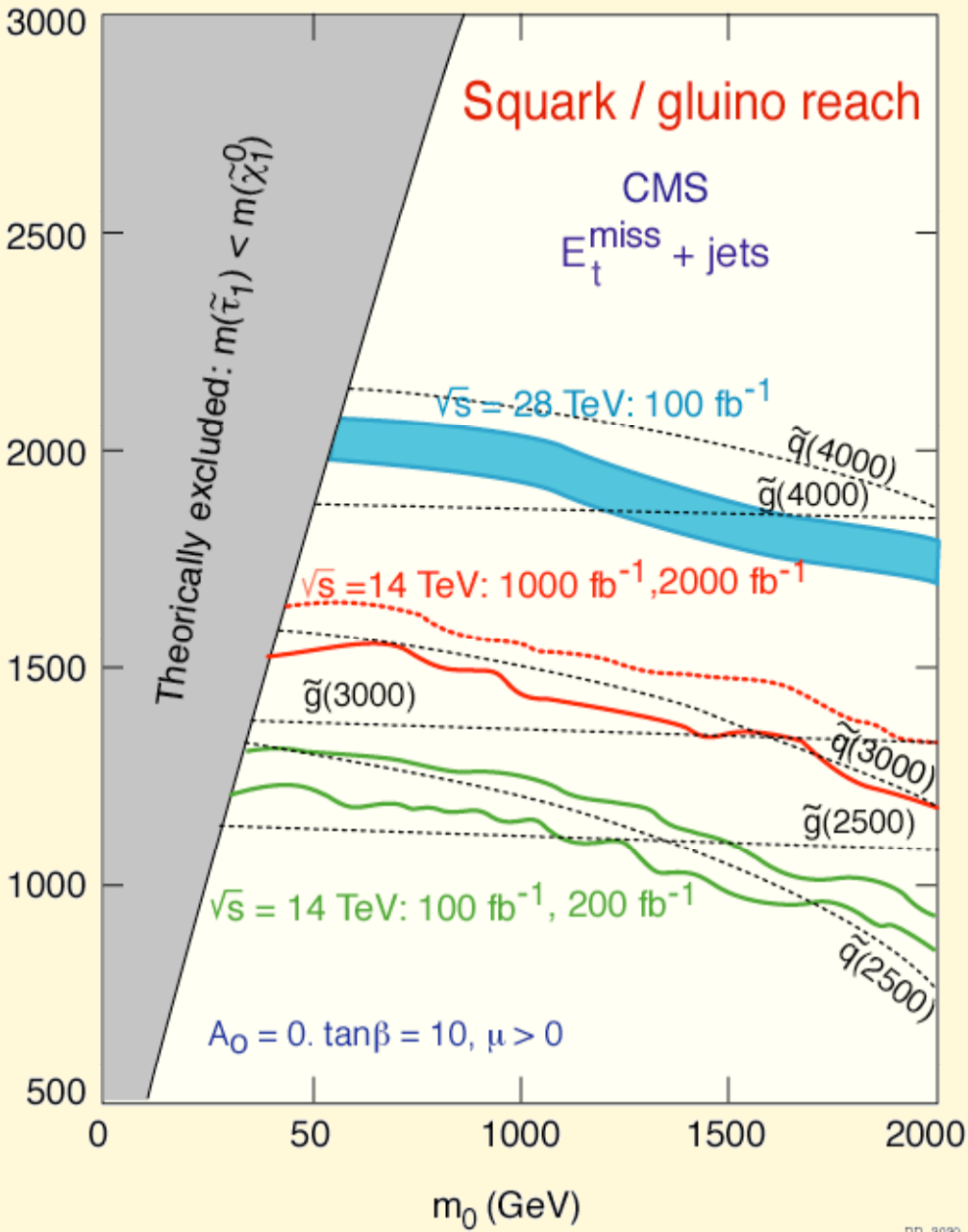


# Detecting the presence of extra H particles (as expected in SUSY)



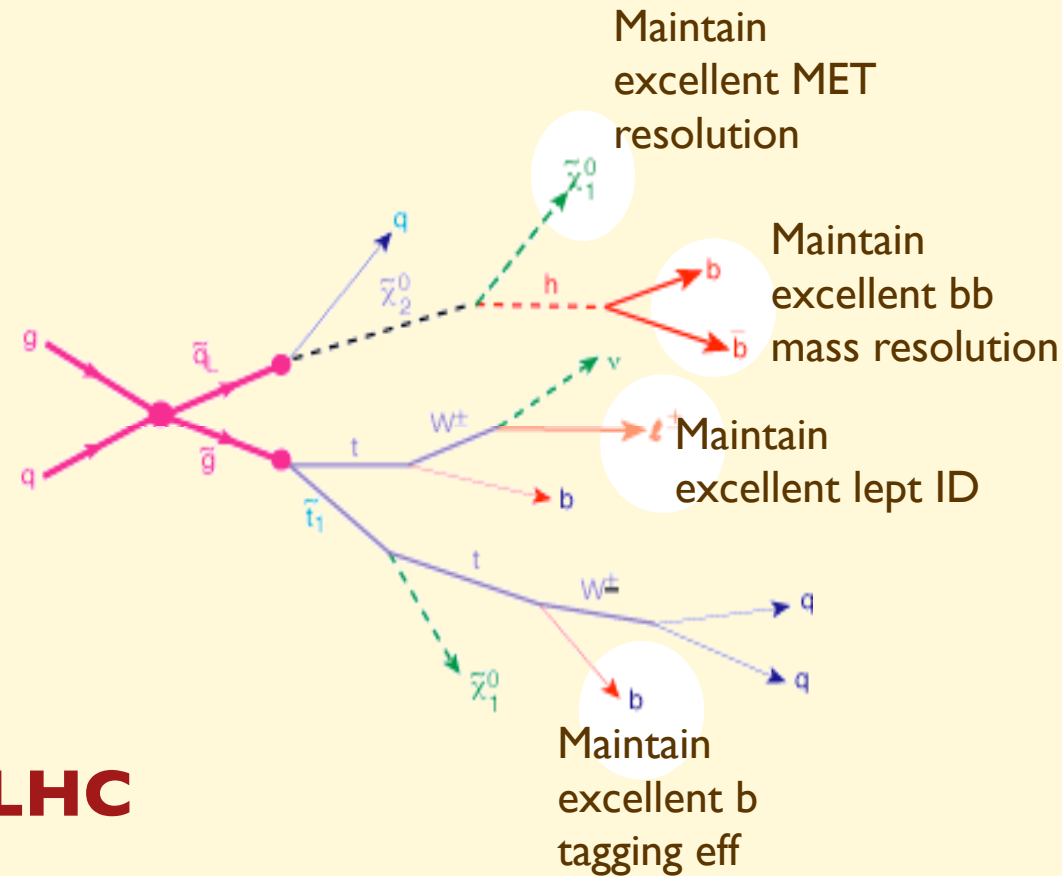


# SUSY reach and studies



**SLHC**

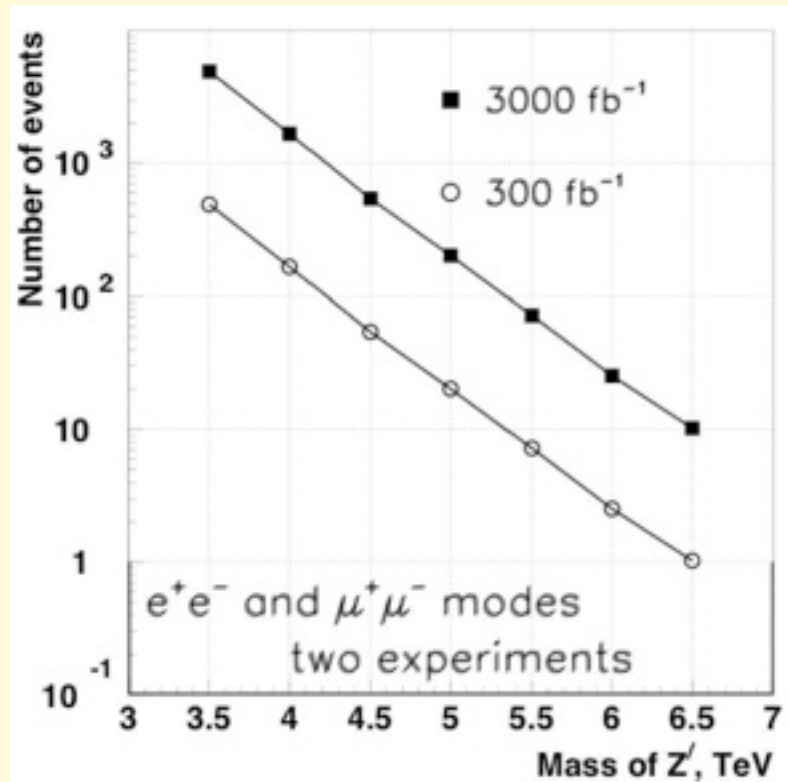
**LHC**



**Mass reach: squarks and gluinos up to 2.5-3 TeV**

# Searching new forces: $W'$ , $Z'$

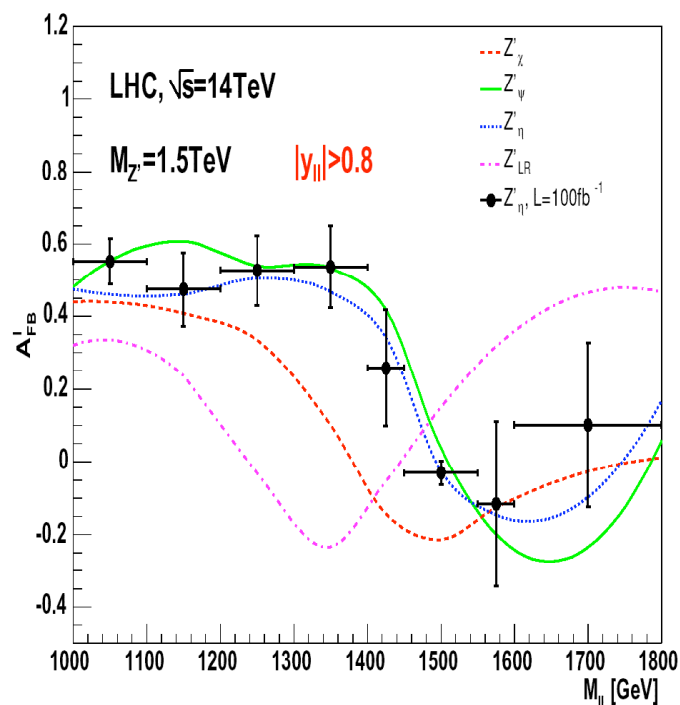
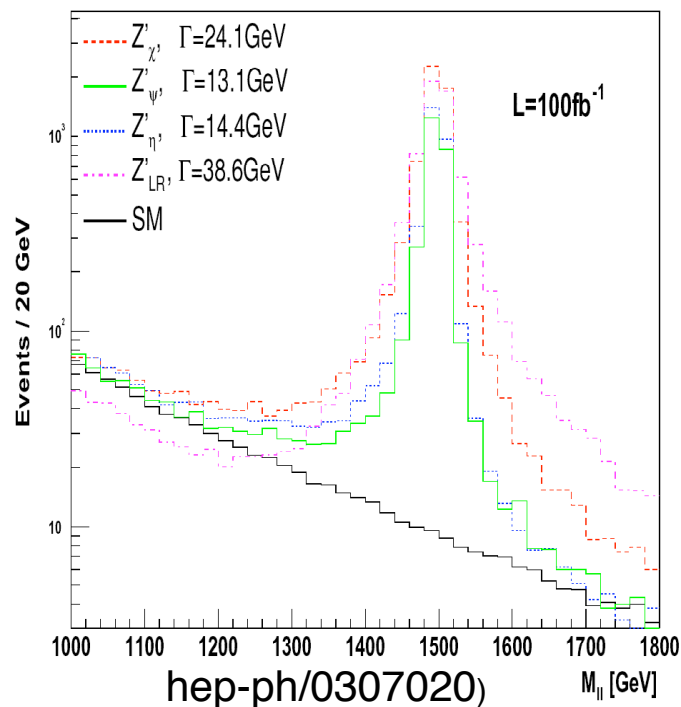
E.g. a  $W'$  coupling to R-handed fermions, to reestablish at high energy the R/L symmetry



**100  $\text{fb}^{-1}$   
discovery reach  
up to  $\sim 5.5$  TeV**

**but ....**

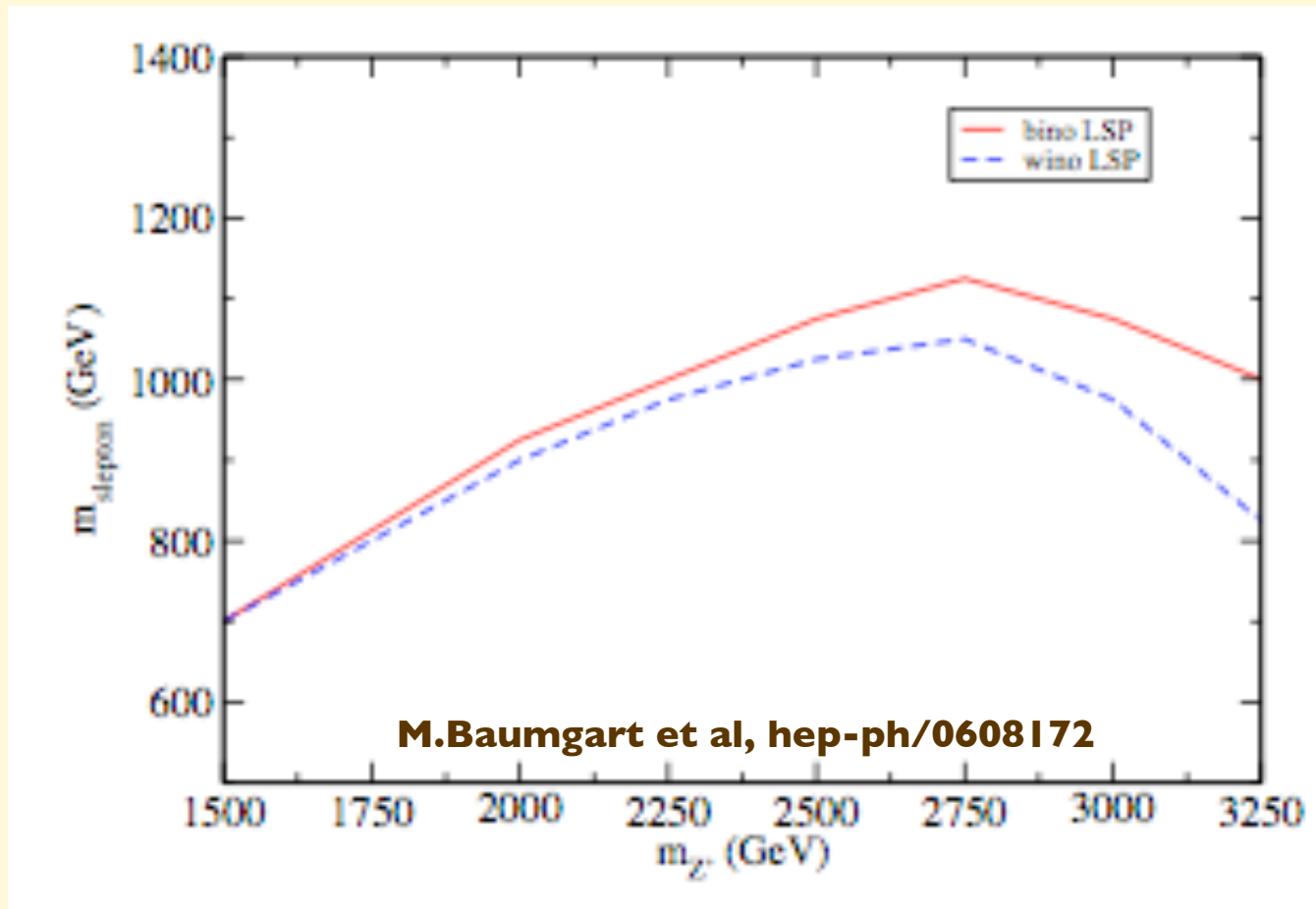
## Differentiating among different $Z'$ models:



**100  $\text{fb}^{-1}$  model  
discrimination  
up to 2.5 TeV**

# Z' and SUSY

Discovery potential ( $100\text{fb}^{-1}$ ) for sleptons, in presence of a Z'



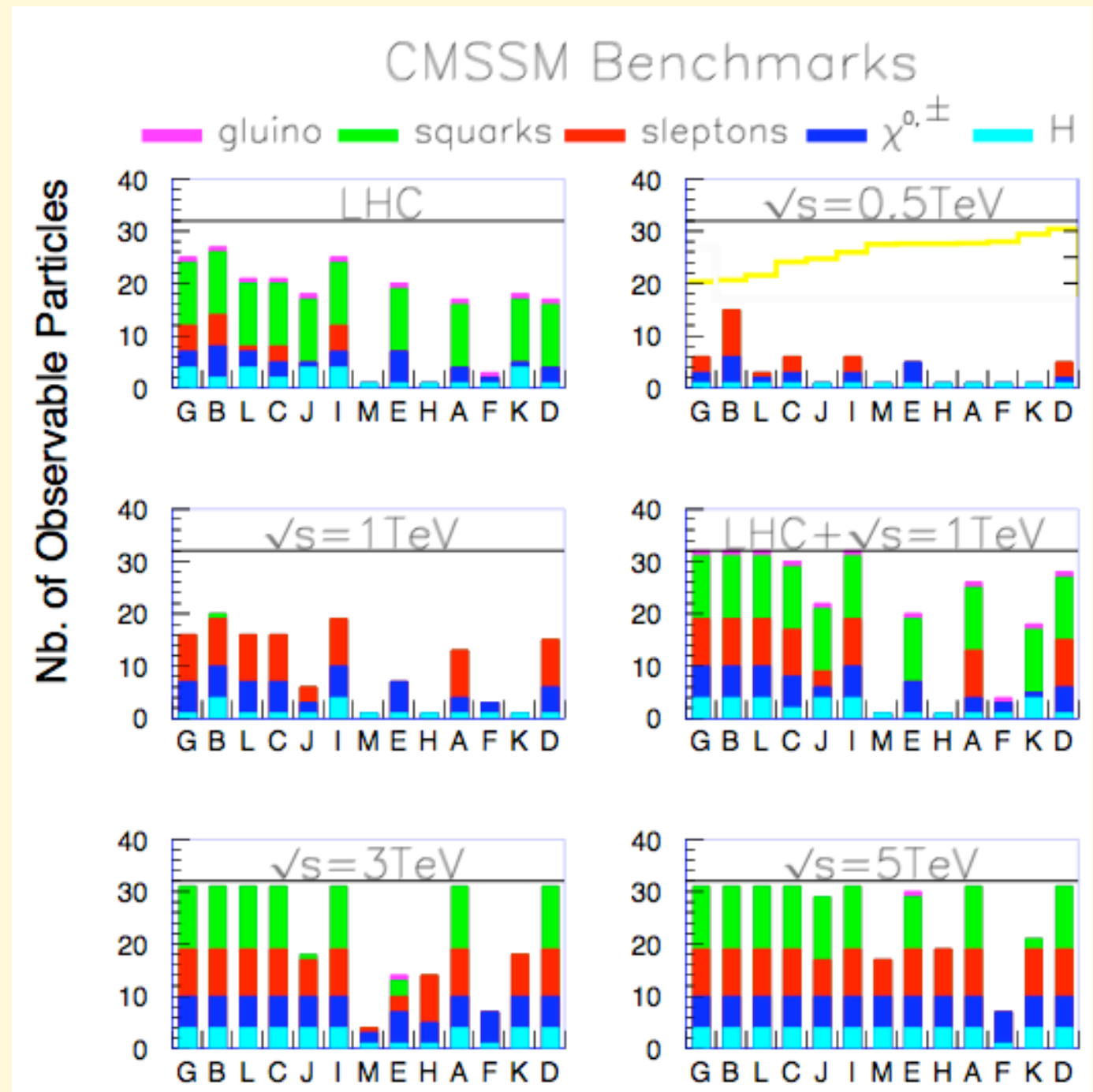
(without Z', can only access masses up to 2-300 GeV)

# SUSY Beyond the LHC: ILC/CLIC

## Example:

Exploration of the Supersymmetric particle spectrum, for 10 different SUSY models

Reference: Physics at CLIC,  
Battaglia, De Roeck, Ellis,  
Schulte eds.,  
hep-ph/0412251



# Neutrinos

- LEP: 3 weakly interacting neutrinos with  $m < M_Z/2$
- 2 relative masses, one absolute mass scale, 3 mixing angles, 1 CKM phase  $\delta$ , 2 extra relative phases if Majorana

$ \Delta m^2_{23} $	$\Delta m^2_{12}$	$m_1$	$\sin^2 \theta_{12}$	$\sin^2 \theta_{23}$	$\sin^2 \theta_{13}$	$\delta_i$
$\sim 2.6 \times 10^{-3}$	$\sim 7 \times 10^{-5}$	?	0.2-0.4	0.3-0.7	<b>&lt;0.05</b>	?

- Iff all  $\theta_{ij} \neq 0$  and at least one phase  $\delta \neq 0$ , then CPV
  - Leptogenesis (lepton-driven B asymmetry of the Universe)
- Dark Matter: WMAP  $\Rightarrow \Omega_\nu < 0.015$ ,  $m_\nu < 0.23$  eV

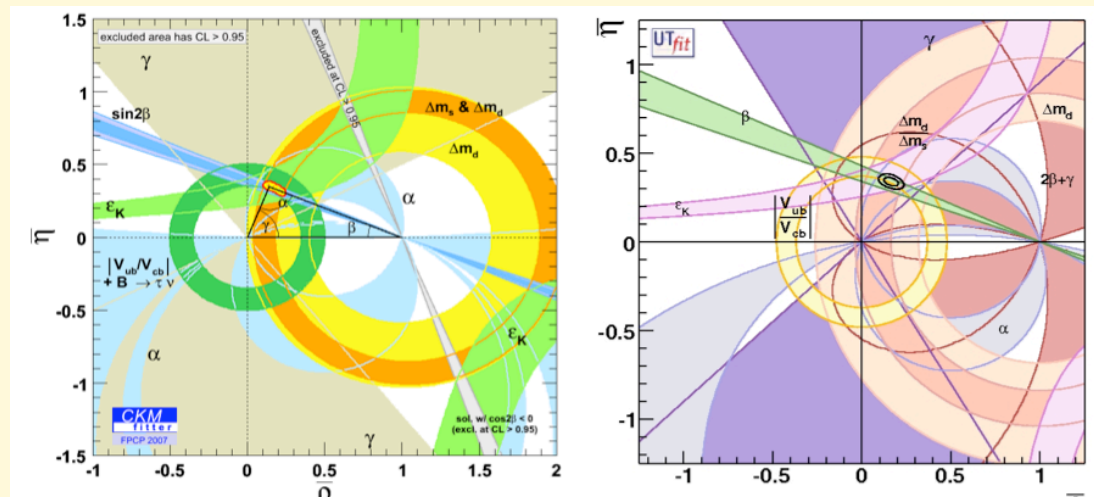
The completion of the neutrino programme, with the full determination of

- mass hierarchy**
- majorana vs dirac nature**
- full spectrum of masses and mixing angles**
- CPV phase(s)**

will “just” put us in the position we are today in the quark sector: we know masses and mixings, but have no idea where they come from.

**This is not enough.**

- To interpret these parameters we need to establish a **connection with the other sectors of the theory**
- We need a **redundancy of inputs** to expose deviations from the simple mixing picture. The equivalent of all redundant measurements of CKM offered by the many channels where we measure CKM angles and phases



# Neutrinos and SUSY

For details and refs, see:  
Masiero, Profumo,  
Vempati, Yaguna, hep-ph/  
0401138

The merging of neutrino masses, SUSY and GUT leads to very interesting constraints and consequences:

SUSY  $\Rightarrow$  Higgs field giving Dirac  $\nu$  mass = Higgs field giving up-quark masses

$$L_m \propto y_\ell H_d L_i L_i^c + y_\nu^{ij} H_u L_i N_j + M_N^{ij} N_i N_j$$

GUT (e.g. SO(10))  $\Rightarrow$  Yukawa  $\nu$ -mass matrix = Up-quark Yukawa matrix

$$L_m \propto y_{i,j}^{d,\ell} \mathbf{16}_i \mathbf{16}_j H_d + y_{i,j}^{u,\nu} \mathbf{16}_i \mathbf{16}_j H_u + y_{i,j}^R \mathbf{16}_i \mathbf{16}_j H_R^{126}$$

$$\text{where } \mathbf{16} = (u_L, d_L, u^c, e^c)_{10} + (d^c, L)_5 + N^c$$

$\Rightarrow$  one entry in the neutrino Yukawa matrix is of order of the top Yukawa coupling!

$$\Rightarrow m(N_R) = f(m_{\text{up}}, m_\nu) \approx (m_t^2 / m_\nu, m_c^2 / m_\nu, m_u^2 / m_\nu)$$

$$\Rightarrow m_\nu > m_t^2 / M_{\text{GUT}} \text{ to ensure that } m(N_R) < M_{\text{GUT}}$$

Even more interestingly, quark mixings induce charged **slepton** mixing via RG evolution from  $M_{GUT}$  to  $m(N_R)$ :

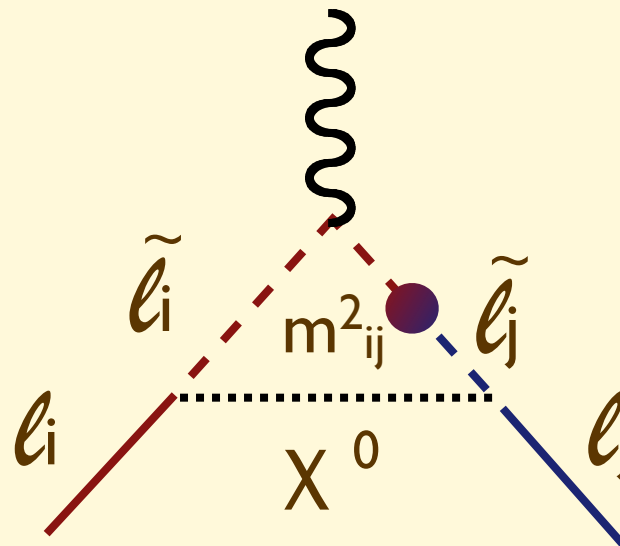
$$(m_{\tilde{L}}^2)_{ij} \sim -\frac{3m_0^2 + A_0^2}{8\pi^2} y_t^2 O_{ij} \log \frac{M_{GUT}}{M_{N_R}}$$

SUSY breaking param's

nu mixing param's

$$y_t^2 O_{ij} = \sum_k y_{ik}^{\nu} y_{jk}^{\nu*}$$

$\ell_i \rightarrow \ell_j \gamma$  transitions:



Possible scenarios:

$$O_{\mu e} = V_{td} V_{ts} \quad \text{“CKM scenario”}$$

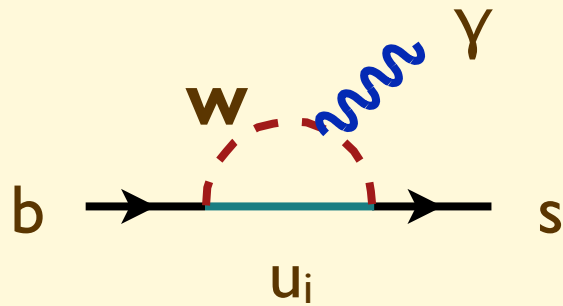
$$O_{\tau \mu} = V_{tb} V_{ts}$$

$$O_{\mu e} = U_{e3} U_{\mu 3} \quad \text{“MNS scenario”}$$

$$O_{\tau \mu} = U_{\tau 3} U_{\mu 3}$$



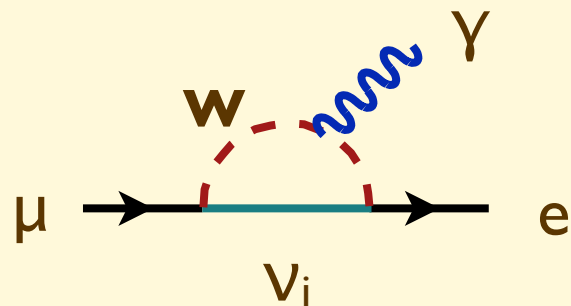
## In the SM



**GIM**

$$\propto \left| \frac{m_c^2 - m_u^2}{m_W^2} V_{cb} V_{cs}^* + \frac{m_t^2 - m_u^2}{m_W^2} V_{tb} V_{ts}^* \right|^2 \sim \frac{m_t^4}{m_W^4} |V_{tb} V_{ts}^*|^2$$

$O(1)$



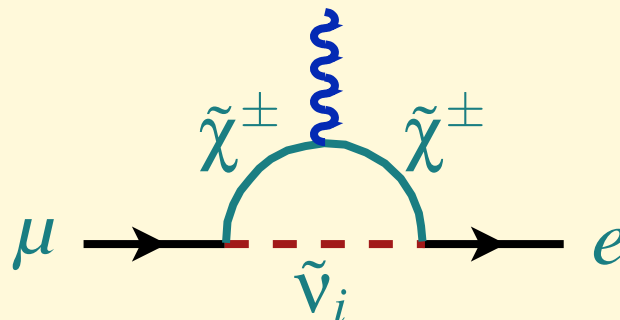
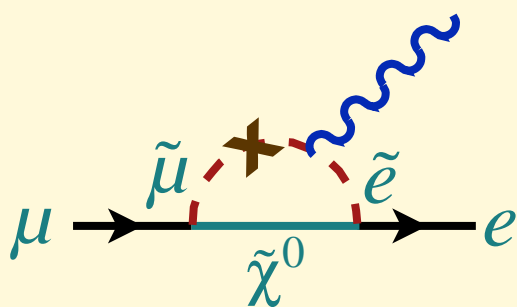
$$\propto \left| \frac{m_1^2 - m_2^2}{m_W^2} M_{12} M_{11}^* + \frac{m_3^2 - m_2^2}{m_W^2} M_{32} M_{31}^* \right|^2 \sim \frac{\Delta m_{23}^4}{m_W^4} s_{23}^2 c_{31}^2 s_{31}^2$$

$O(10^{-49})$

The smallness of  $B(\mu \rightarrow e \gamma)$  is entirely due to the smallness of  $\nu$  masses (and splittings)

The moment we have new states in the loop, the rates goes up!

## Example: SUSY



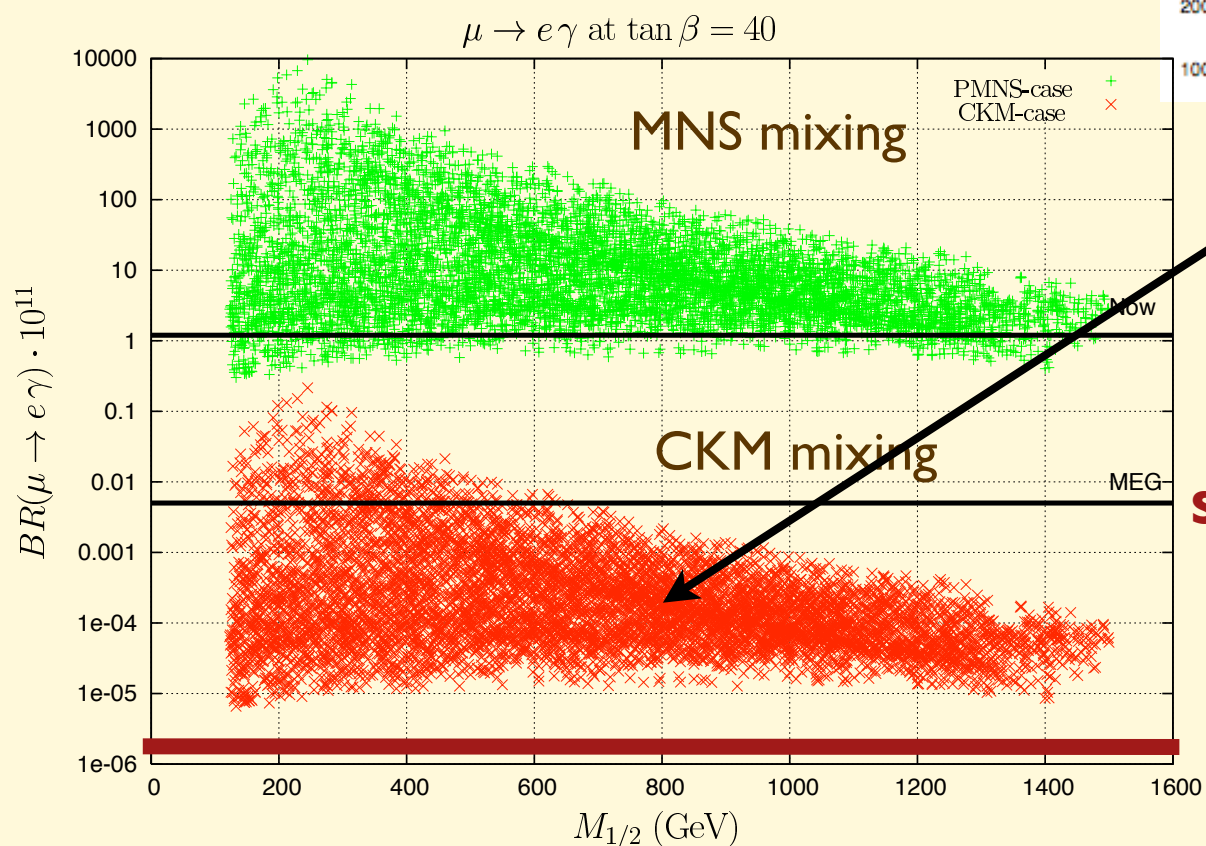
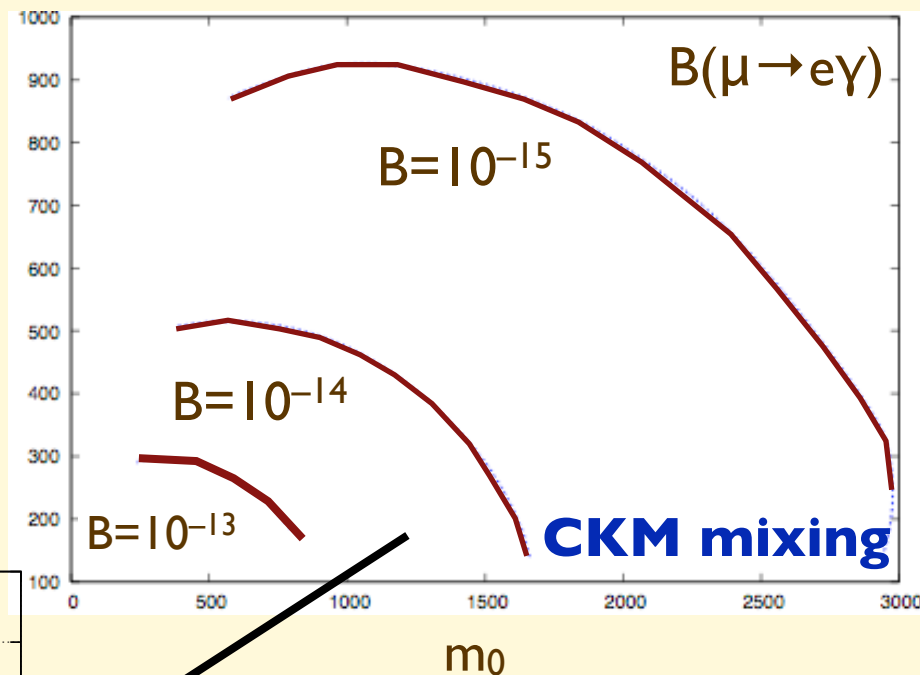
$$B \propto \left| \frac{\Delta m^2(\tilde{\nu})}{m_{\tilde{\chi}}^2} \times \epsilon_{12}^2 \right|^2$$

# Examples of LHC-( $\mu \rightarrow e\gamma$ ) synergy:

SO(10) GUT scenario, slepton mixing induced by RG evolution

To push to the ultimate LHC squark reach ( $m \sim 2.5\text{--}3\text{ TeV}$ ) may require sensitivity to  $B(\mu \rightarrow e\gamma) \sim 10^{-15}$

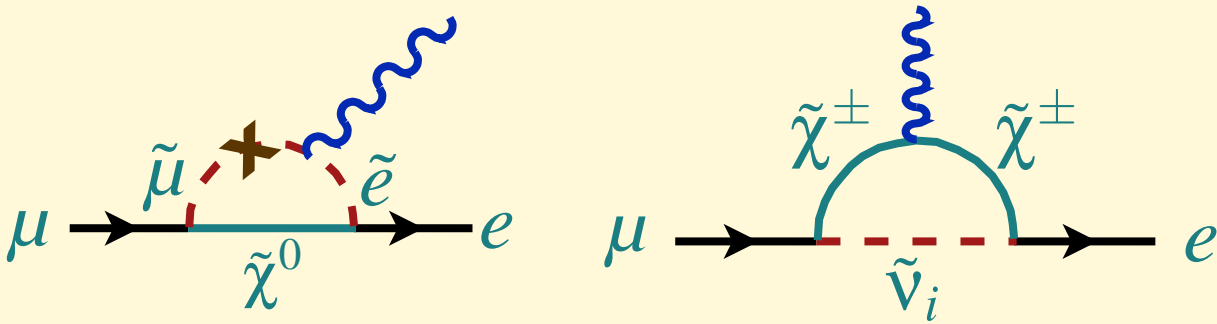
$m_{1/2}$



SO(10) mSUGRA scan  
with  $m(\text{squark}) < 2.5\text{ TeV}$   
Calibbi et al, hep-ph/0605139

**Sensitivity of MEG experiment**

**Sensitivity of Project-X mu2e conversion**



**Neglecting mixing, these diagrams are also responsible for  $(g-2)_\mu$**

**Assuming that the BNL data are explained by SUSY,**

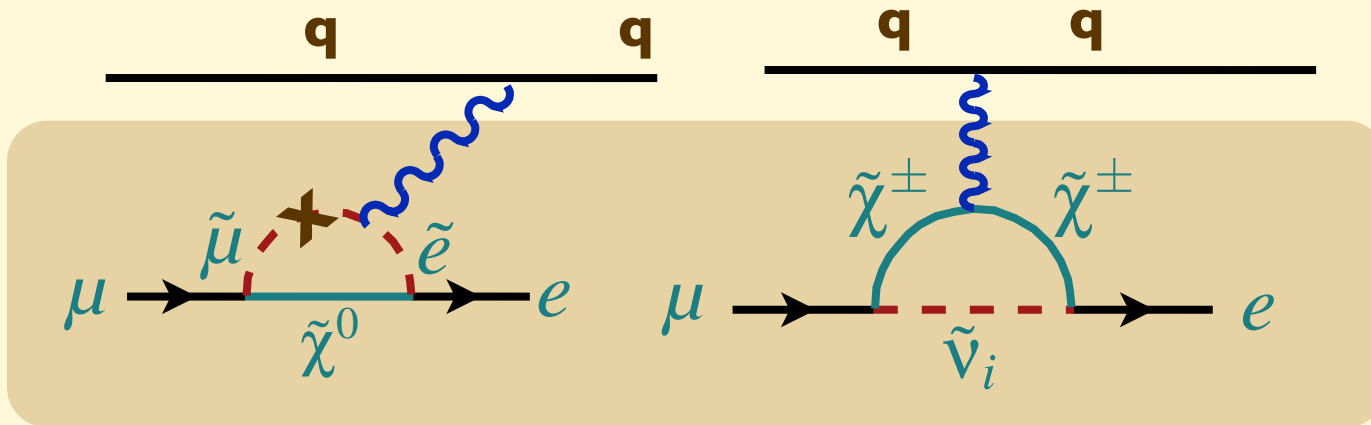
$$(g-2)_\mu^{\text{data}} - (g-2)_\mu^{\text{SM}} = (g-2)_\mu^{\text{SUSY}}$$

**sets a scale for  $m(\text{SUSY}) \sim 100 \text{ GeV}$**

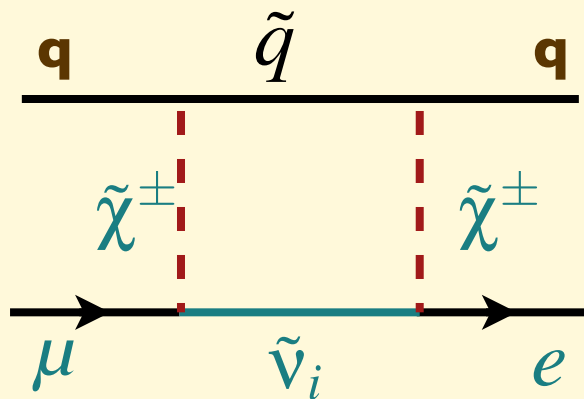
**Current  $B(\mu \rightarrow e\gamma)$  limits then indicate mass splittings in the slepton sector of few 10s MeV !!**

**Sensitive to natural mass splittings  $m(\tilde{\mu}) - m(\tilde{e}) \sim \mathcal{O}(m_\mu)$**

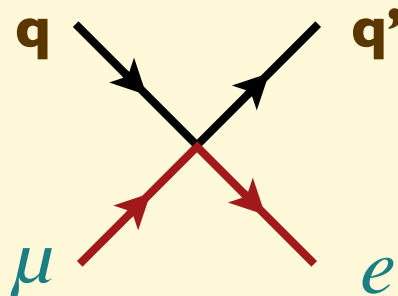
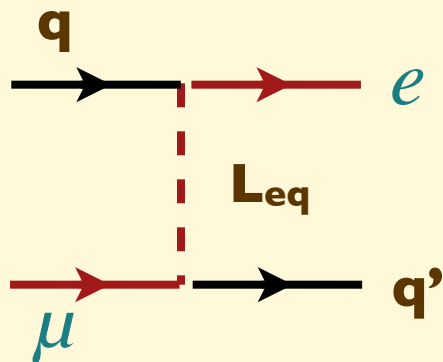
# $\mu \rightarrow e \gamma$ vs $\mu N \rightarrow e N$ complementarity



$\mu \rightarrow e \gamma$  diagrams

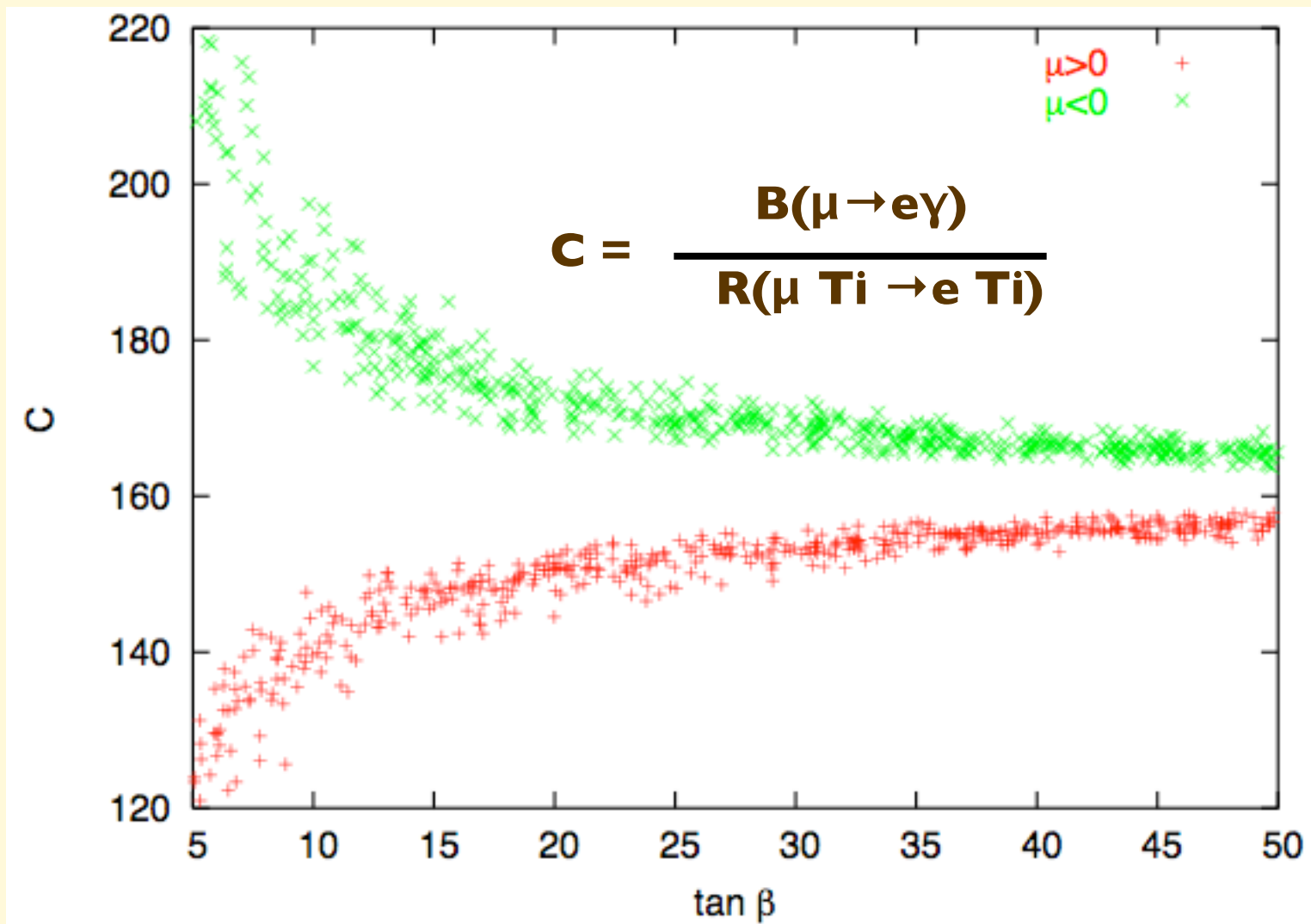


extra contributions,  
sensitive to additional  
model parameters



extra contributions,  
sensitive to other  
underlying dynamics

$K \rightarrow e \mu$ ?



C Yagouna, hep-ph/0502014

$$\mu \rightarrow e \gamma$$

# Current limits on $B(\mu \rightarrow e \gamma)$

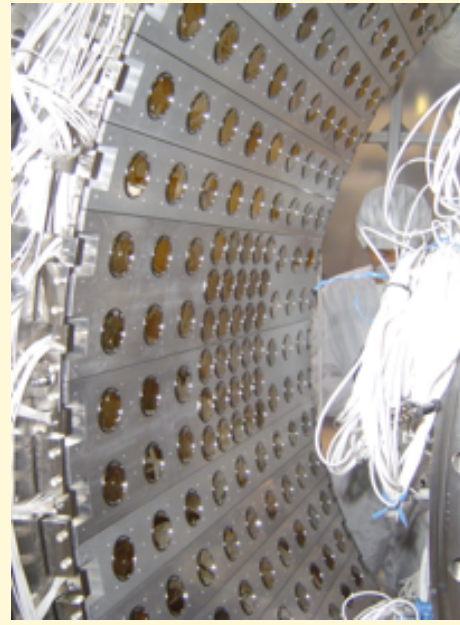
mode	upper limit (90% C.L.)	year	Exp./Lab.
$\mu^+ \rightarrow e^+ \gamma$	$1.2 \times 10^{-11}$	2002	MEGA / LAMPF
$\mu^+ \rightarrow e^+ e^+ e^-$	$1.0 \times 10^{-12}$	1988	SINDRUM I/ PSI
$\mu^+ e^- \leftrightarrow \mu^- e^+$	$8.3 \times 10^{-11}$	1999	PSI
$\mu^- \text{ Ti} \rightarrow e^- \text{ Ti}$	$6.1 \times 10^{-13}$	1998	SINDRUM II / PSI
$\mu^- \text{ Ti} \rightarrow e^+ \text{ Ca}^*$	$3.6 \times 10^{-11}$	1998	SINDRUM II / PSI
$\mu^- \text{ Pb} \rightarrow e^- \text{ Pb}$	$4.6 \times 10^{-11}$	1996	SINDRUM II / PSI
$\mu^- \text{ Au} \rightarrow e^- \text{ Au}$	$7 \times 10^{-13}$	2006	SINDRUM II / PSI

# Future:

near:  $\mu \rightarrow e \gamma$

**MEG at PSI** <http://meg.web.psi.ch/>

- o First data taking 2008, single-event sensitivity at  $BR < 5 \times 10^{-12}$
- o Run Fall 2009 underway, full detector => match current limits
- o New 2-yr run to start April 2010
- o ultimate sensitivity:  $BR < 1 \times 10^{-14}$  at 90%CL by 2011



far:  $\mu \rightarrow e \text{ conv}$

**MU2e at Fermilab**

- o 1st stage DoE approval achieved Fall 09. Could be taking data by 2016
- o sensitivity:  $R(\mu \rightarrow e) < 6 \times 10^{-17}$  @90%CL  $\rightarrow 10^{-18}$  with Project-X



(i.e.  $BR < \sim 10^{-14}$  if only  $(\mu \rightarrow e \gamma)$  diagrams contribute)

# More physics with charged leptons

- $\mu \rightarrow eee$  (typically  $O(\alpha)$ , but  $O(1)$  in LH models)
- $\tau \rightarrow \mu \gamma$   $\tau \rightarrow e \gamma$  : model-dependent correlations with  $\mu \rightarrow e \gamma$
- $\tau \rightarrow \mu \mu \mu$  (LHCb ?)
- CP violation in SM-allowed  $\tau$  decays?
  - $O(10^{-3})$  CP asymmetry in  $\tau \rightarrow \nu K \pi \Rightarrow B(\tau \rightarrow \mu \gamma) \sim O(10^{-9})$
- .....

# Example of correlations between $\nu$ and quark-sector observables

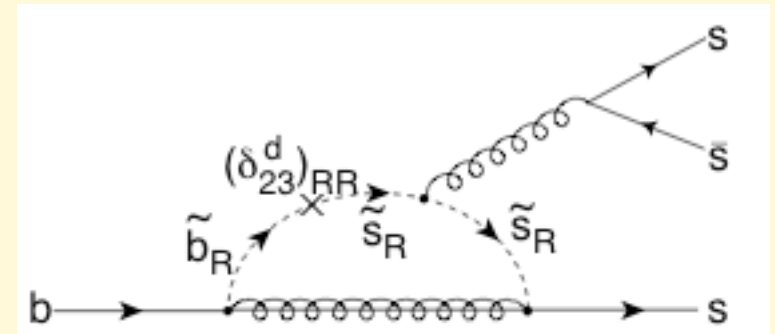
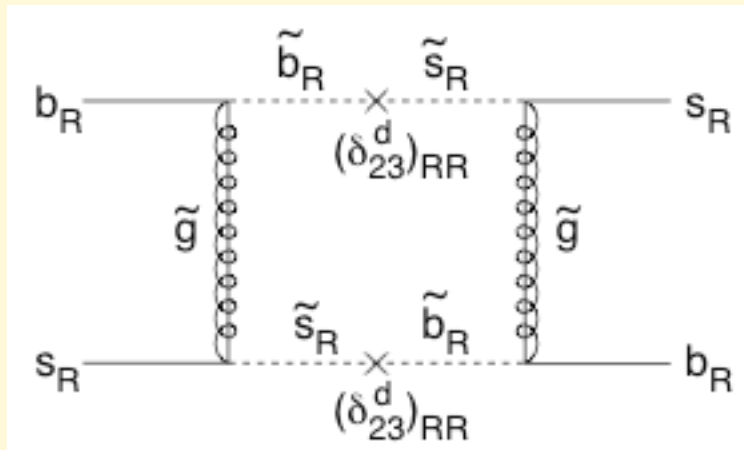
$$L_m \propto y_{i,j}^{d,\ell} \mathbf{16}_i \mathbf{16}_j H_d + y_{i,j}^{u,\nu} \mathbf{16}_i \mathbf{16}_j H_u + y_{i,j}^R \mathbf{16}_i \mathbf{16}_j H_R^{126}$$

$$\mathbf{16} = (u_L, d_L, u^c, e^c)_{10} + (d^c, L)_5 + N^c$$

A large mixing between  $\nu_\mu$  and  $\nu_\tau$  implies a large mixing between

$$(b_R, \bar{\nu}_\tau, \tau^+) \quad (s_R, \bar{\nu}_\mu, \mu^+)$$

This has no direct impact on phenomenology, since right-handed quarks do not couple to weak interactions. However it leads to a large mixing between the scalar partners of R-handed squarks, and to interactions like



with potentially large contributions to:

**$B_s$  mixing, CP violation in  $B_s \rightarrow \phi \psi$  ( $\sim 0$  in the SM)**

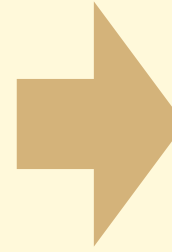
$$\sin 2\beta(B \rightarrow \phi K_s) \neq \sin 2\beta(B \rightarrow \psi K_s)$$



# EDMs

## ● Flavour-conserving CPV

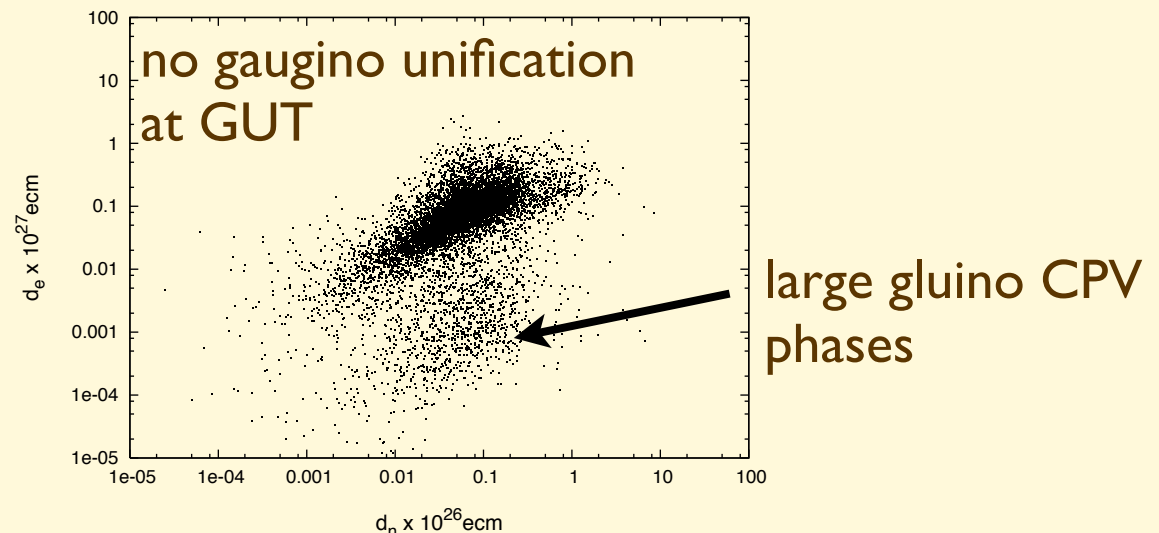
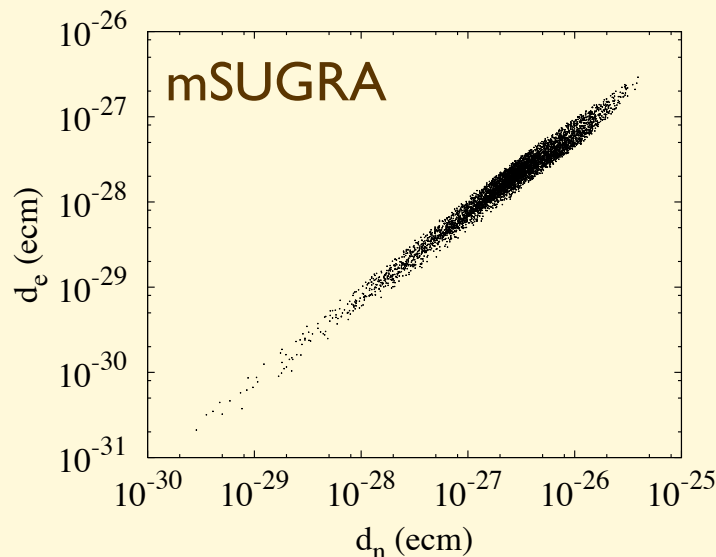
- Sensitive probes of CPV in extended gauge sectors (e.g. SUSY gluinos, gauginos, higgsinos)



Probes of mechanisms to generate the antimatter asymmetry of the universe

$d_e / d_n$  correlations:

**SUSY:  $d_e / d_n \sim m_e/m_q \sim 0.1$**



**Extra-dim, 2HDM:  $d_e / d_n \ll 1$**

## Atoms:

- paramagnetic (Tl):**
  - fundamental electron EDM
  - CPV  $eeqq$  interactions
- diamagnetic (Hg):**
  - fundamental electron EDM
  - fundamental quark EDM and  $\theta_{\text{QCD}}$
  - CPV  $eeqq$  interactions
- heavy molecules with unpaired electrons (YbF):**
  - fundamental electron EDM

## Neutron:

- fundamental quark EDM and  $\theta_{\text{QCD}}$
- higher-dim CPV  $qq$  operators ( $\text{int}^{\text{ns}}$  with gluinos, etc)

# Neutron EDM

Current limit:  $d_{\text{neutron}} = 3 \times 10^{-26} \text{ e cm}$

C.A. Baker et al, (RAL, Sussex, ILL Grenoble)  
<http://arxiv.org/pdf/hep-ex/0602020>

## Forthcoming experiments with ultracold neutrons:

### ILL (Grenoble) and PSI

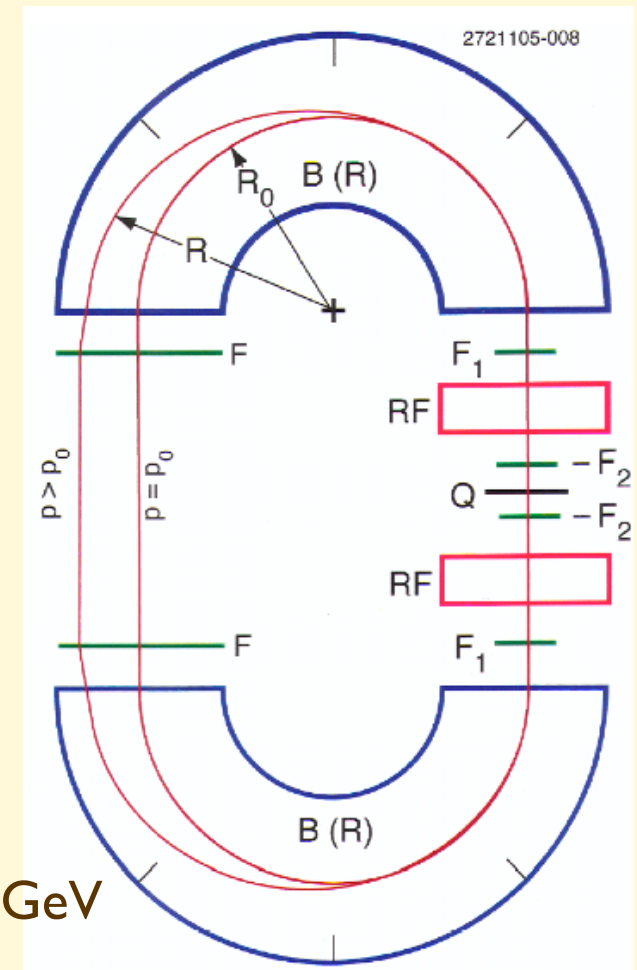
- o R&D and construction of new detectors/beamline
- o new runs 2009-2011 (ILL) and 2011-2014 (PSI)
- o Goal:  $d_{\text{neutron}} < \sim 2 \times 10^{-28} \text{ e cm/yr}$

⇒ **probe SUSY CPV phase of  $O(10^{-4})$**

## Deuteron EDM in a storage ring

Orlov, Morse, Semertzidis,  
<http://arxiv.org/pdf/hep-ex/0605022>

- o Inject deuterons from LEIR, CERN's low-energy ion ring used to prepare heavy ion beams for the LHC
- o Sensitivity:  $\sigma_d = 2.5 \times 10^{-29} \text{ e cm/yr}$



1.5 GeV

# Rare K decays

$$K_L^0 \rightarrow \pi^0 \nu \nu \quad B(K_L^0 \rightarrow \pi^0 \nu \nu)_{\text{SM}} = 2.8 \pm 0.4 \times 10^{-11}$$

**E391 at KEK**, final result:

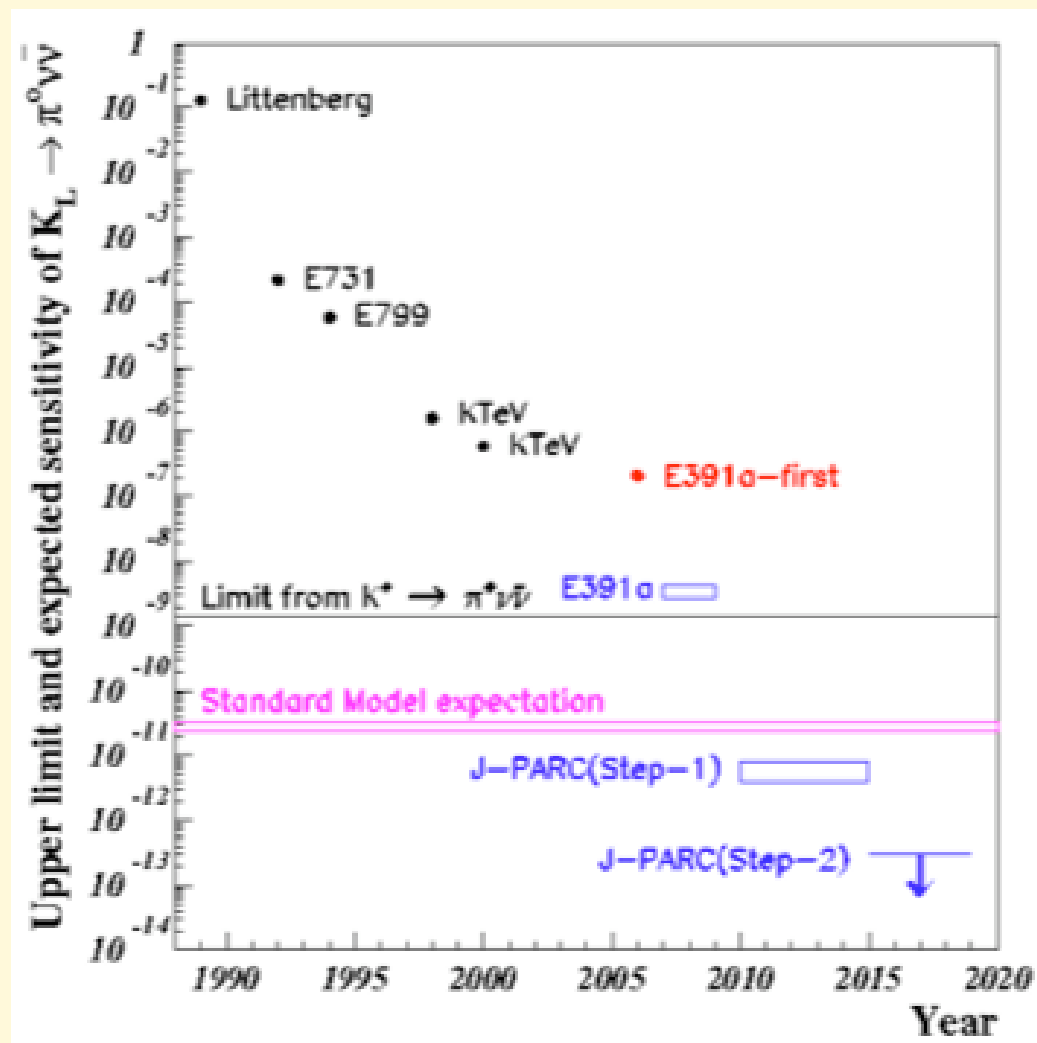
$$B(K_L^0 \rightarrow \pi^0 \nu \nu) < 2600 \times 10^{-11} \text{ @ 90\% CL}$$

arXiv:0911.4789

## KOTO (E14) at JPARC

[http://www-ps.kek.jp/jhf-np/NuclPart/0701/Day2\\_AM/E14.ppt.pdf](http://www-ps.kek.jp/jhf-np/NuclPart/0701/Day2_AM/E14.ppt.pdf)

- o Aim to 3000 x rate[E391]
- o Detector completion: 2010
- o Beam construction/survey: 2009. Status:  $K_L$  observed, Dec 7
- o Data: 2012-20
- o Goal:  $O(10^{-13})$ ,  $\Delta BR \sim 10\%$



# Rare K decays, CERN

$$\mathbf{K^+ \rightarrow \pi^+ \nu \nu} \quad B(K^+ \rightarrow \pi^+ \nu \nu)_{E787/949 \text{ BNL}} = \mathbf{1.73 \pm 1 \times 10^{-10}} \quad (3 \text{ events, arXiv:0808.2459})$$

$$B(K^+ \rightarrow \pi^+ \nu \nu)_{SM} = \mathbf{0.85 \pm 0.07 \times 10^{-10}}$$

Expected reduction to 4% error via NNLO  
+better input parameters ( $m_{top}$ , etc)

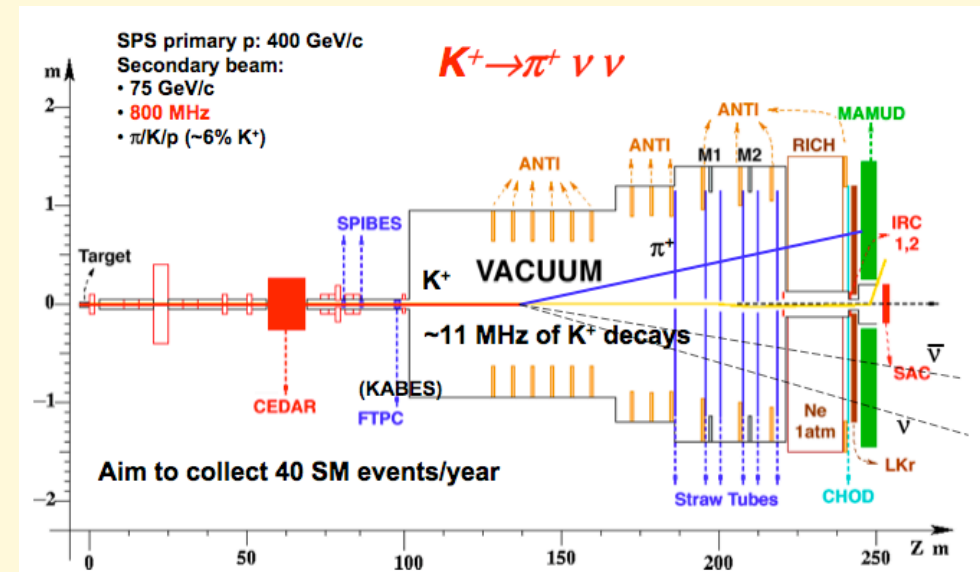
## NA62

<http://na48.web.cern.ch/NA48/NA48-3/>

o R&D ongoing, with 2007 run for  
 $R_{e/\mu} = \Gamma(K \rightarrow e \nu) / \Gamma(K \rightarrow \mu \nu)$  to 0.3%

o Ready for beam: **Fall 2012**

o Goal: 80 events (@SM rate) in 2 yrs  
of run,  $S/B=10/1 \Rightarrow \delta|V_{td}|=10\%$



$$\mathbf{K_L^0 \rightarrow \pi^0 e^+ e^-} \quad \mathbf{K_L^0 \rightarrow \pi^0 \mu^+ \mu^-}$$

$$\mathbf{K_L^0 \rightarrow \pi^0 \nu \nu}$$

**NA48/4**

**NA48/5**

**Require more protons  
than available from  
the SPS today**

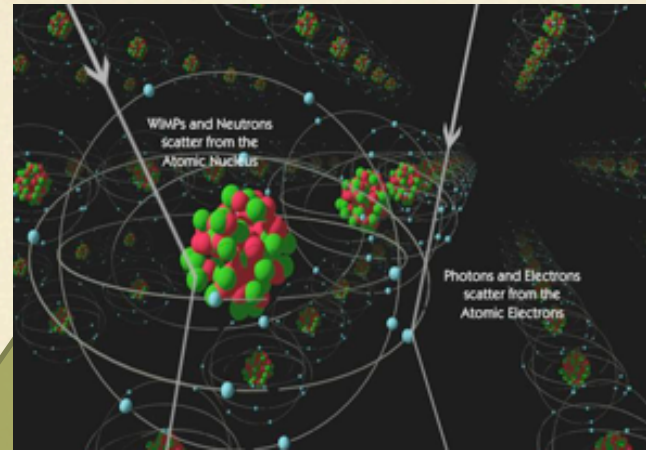
# 2015, a dream scenario

- SUSY is seen at LHC, with squarks/gluinos at  $\sim 1$  TeV, weak sparticles at  $\sim 0.1$  TeV
- Observe in parallel:
  - LHC:  $B_s \rightarrow \mu^+ \mu^-$  at rates  $> \text{SM}$ , and NA62:  $K^+ \rightarrow \pi^+ \nu \nu$  at rates  $> \text{SM}$
  - MEG:  $\mu \rightarrow e \gamma$
  - ILL/PSI: neutron EDM
  - large  $\theta_{13} \rightarrow$  measurable CP violation in nu mixing
- $Z'$  at 2–2.5 TeV seen at LHC:
  - open decays to all SUSY sparticles  $\Rightarrow$  very accurate studies
  - the LHC turns into a  $Z'$  factory
  - CLIC is above threshold to further study it in the future
- Direct DM detection underground fits well neutralino properties
- .....

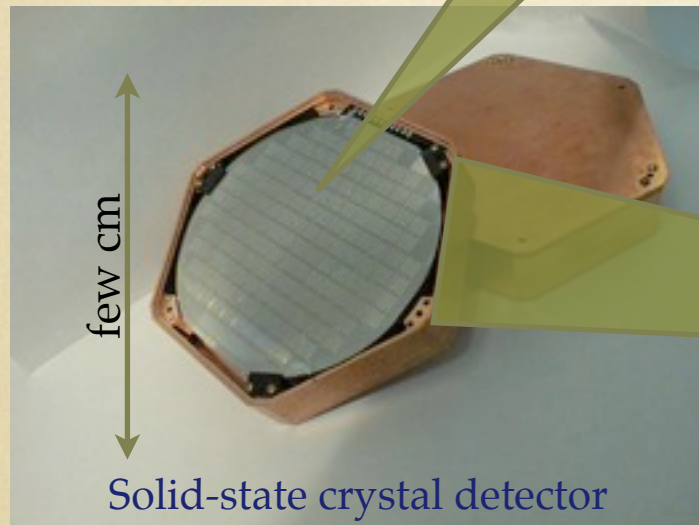
**It's a dream, but not an impossible one !**



# Search for Direct detection of Dark Matter particles flying through the galaxy



CDMS experiment,  
Stanford and Soudan mine  
See <http://cdms.berkeley.edu/>



Cryogenic box

Dec 17 2009:

- reported results of 2007-08 exposure (2xprevious results)
- **2 events**  $\sim \rightarrow$  **25% probability of being bg**
- will have 3 x the current detector volume available by Summer 2010

# Conclusions

- **Progress in the field will be 100% driven by new and better experimental data.** We are running out of ideas and tools to make progress based on first principles only.
- Nevertheless, we created scenarios for BSM physics which, in addition to addressing the most outstanding **theoretical puzzles** and the **established deviations** from the SM (DM, BAU, nu mixing), predict galore of new phenomena at energy and accuracy scales just behind the corner
- Maintaining **diversity** in the exp'l programme is our best investment for HEP.